

V Signals of Dark Matter

Roni Harnik, Fermilab

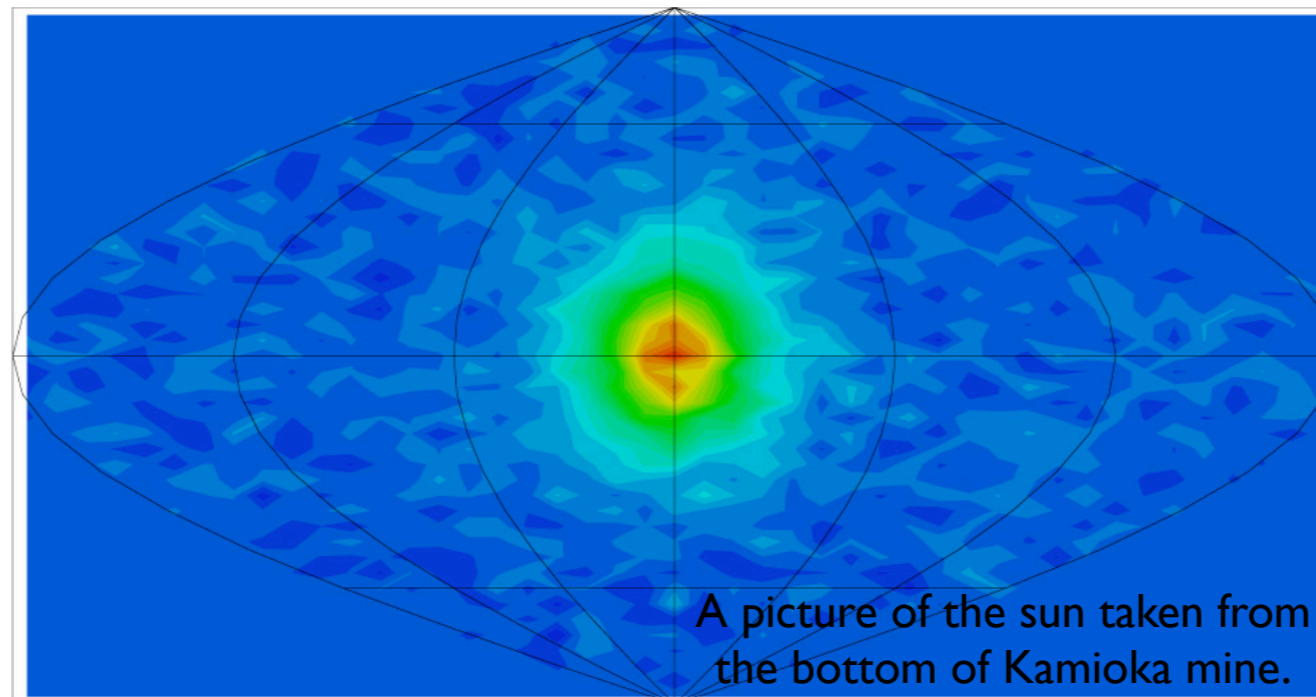
with Joachim Kopp and Pedro Machado
1202.6073

See related works-
Pospelov 1103.3261
Pospelov and Pradler 1203.0545

yesterday's talk

Direct Detection

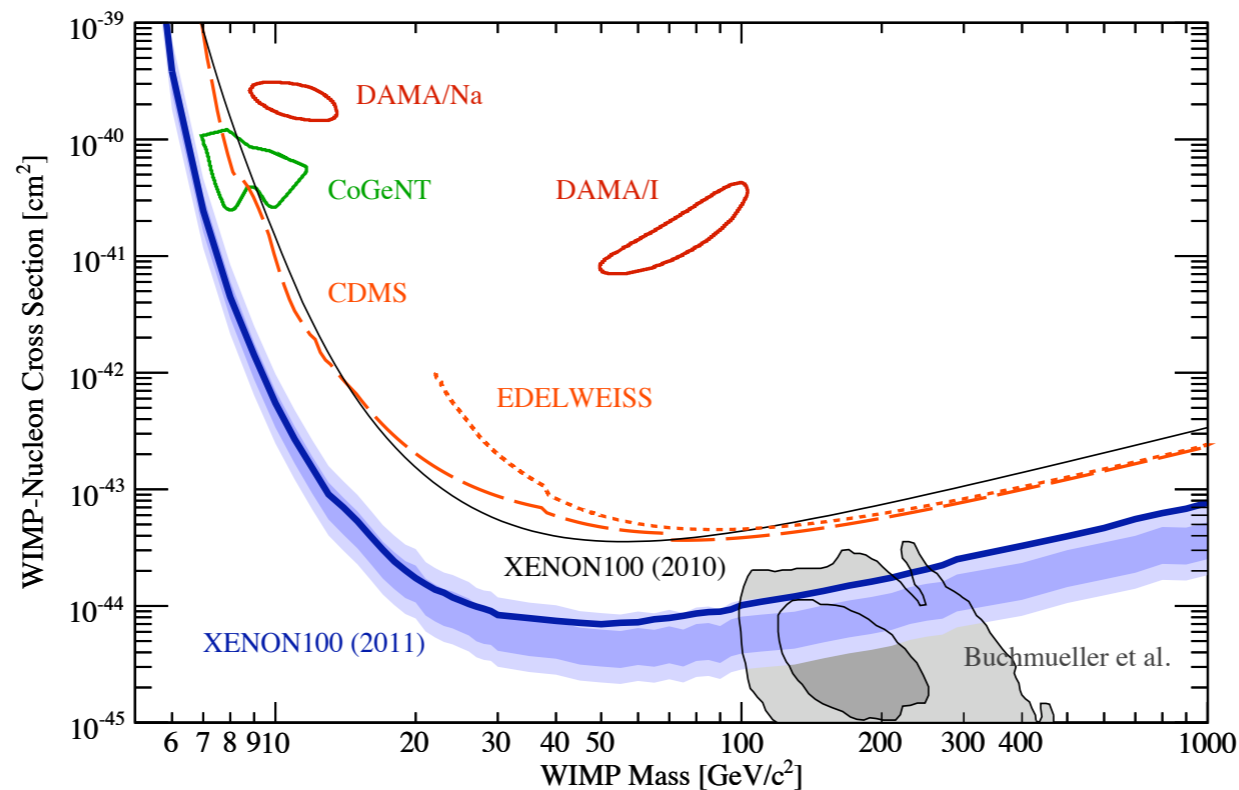
- * In past decades direct detection collaborations have made a heroic (and successful!) effort in background reduction.
- * Try as you might, these experiments are still completely exposed to the sun....



...in neutrinos!!!

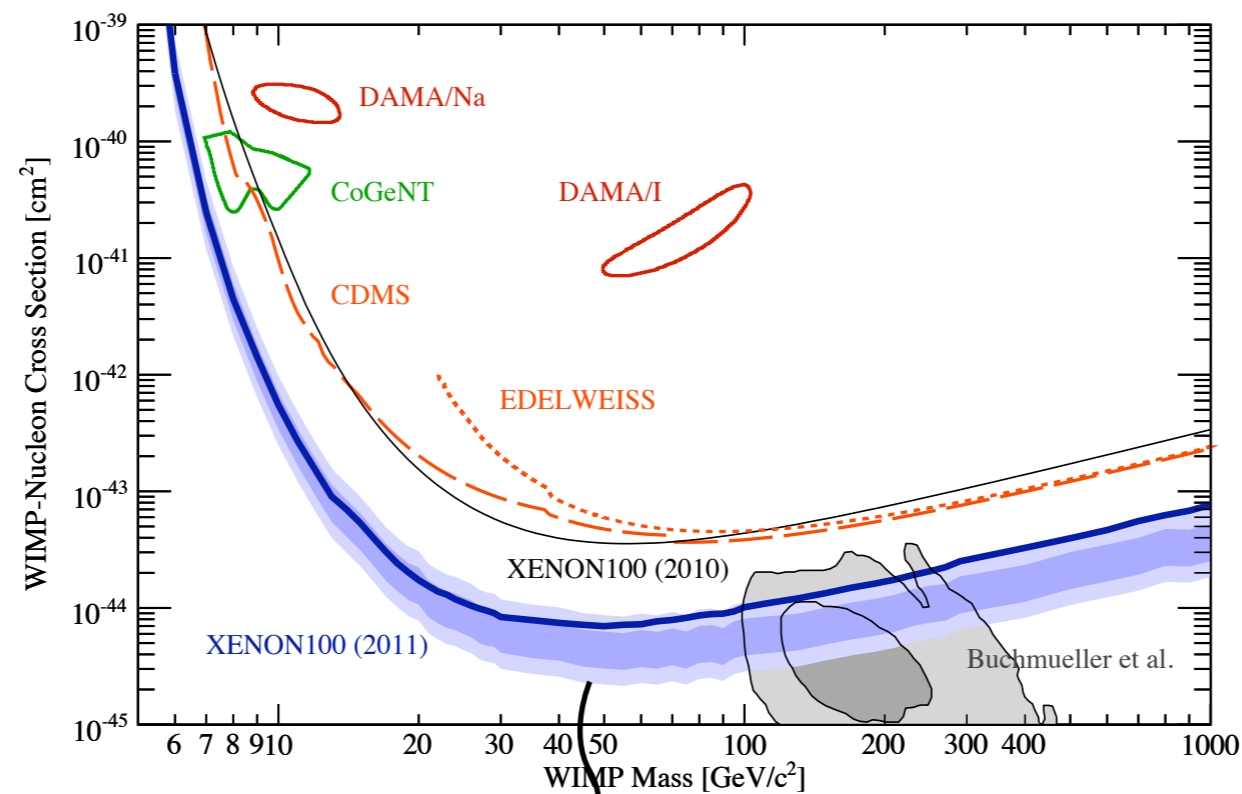
Direct Detection

- * The irreducible background from solar and atmospheric neutrinos will be the eventual end-game of direct detection.



Direct Detection

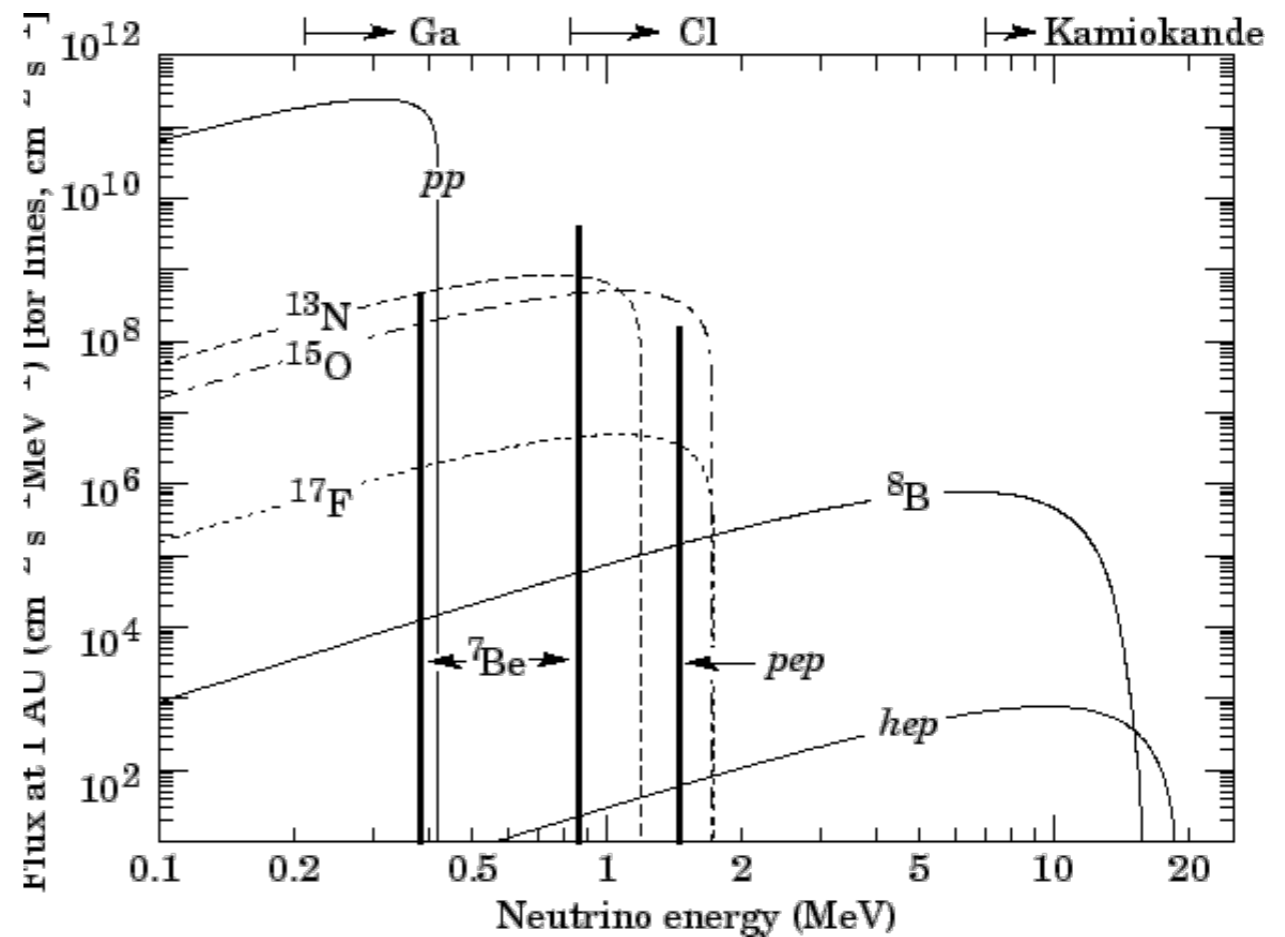
- * The irreducible background from solar and atmospheric neutrinos will be the eventual end-game of direct detection.



can this
Background be
"raised" to be a
signal?



What does this look like in a direct detection experiment?



$$\frac{dR}{dE_r} = N_T \int_{E_\nu^{\min}}^{\infty} \frac{d\Phi}{dE_\nu} \frac{d\sigma}{dE_r} dE_\nu$$

with $E_\nu^{\min} = \frac{1}{2} \left(E_r + \sqrt{E_r^2 + 2E_r m_T} \right)$

(relativistic version of $\frac{dR}{dE_R} = \frac{M_N N_T \rho_\chi \sigma_n}{2m_\chi \mu_{ne}^2} F^2 \frac{(f_p Z + f_n (A - Z))^2}{f_n^2} \int_{v_{min}} \frac{f(v)}{v}$)

Spectrum

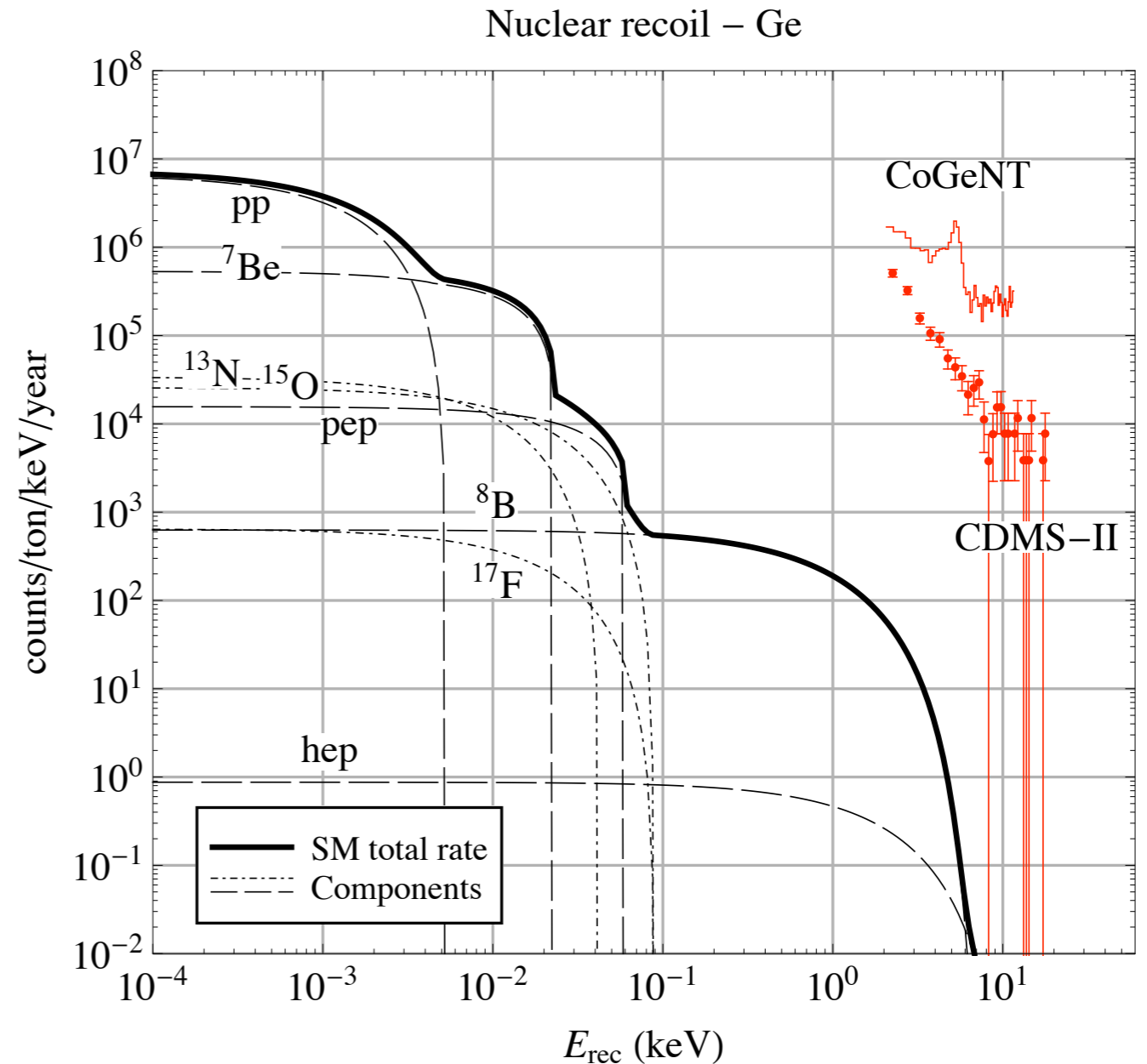
* For nuclear recoil:

see Joseph Pradler's talk

* The scattering is coherent, $\propto A^2$.

* The flux is low above threshold.

$$E_r^{\max} = \frac{2E_\nu^2}{m_T + 2E_\nu}$$



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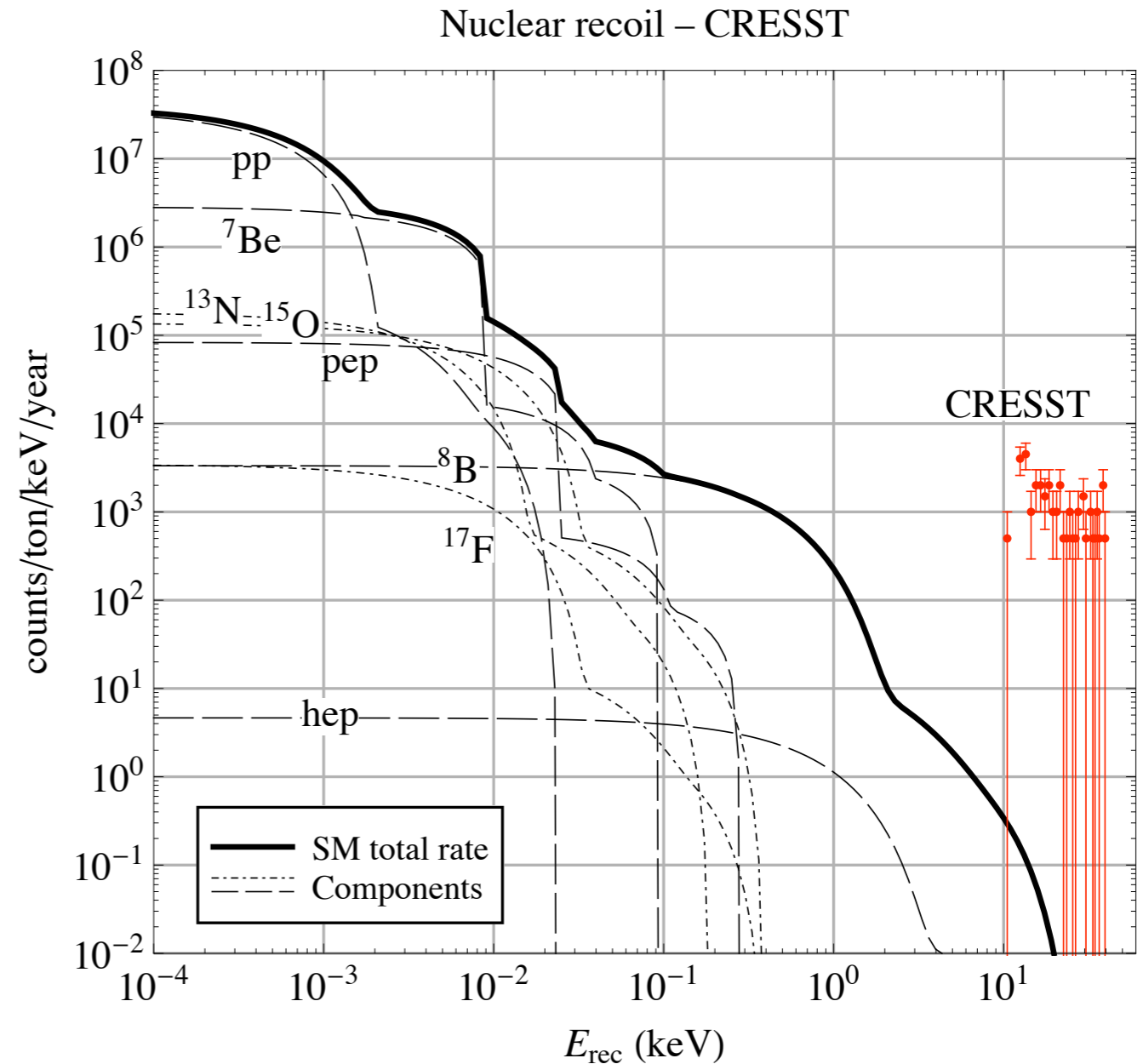
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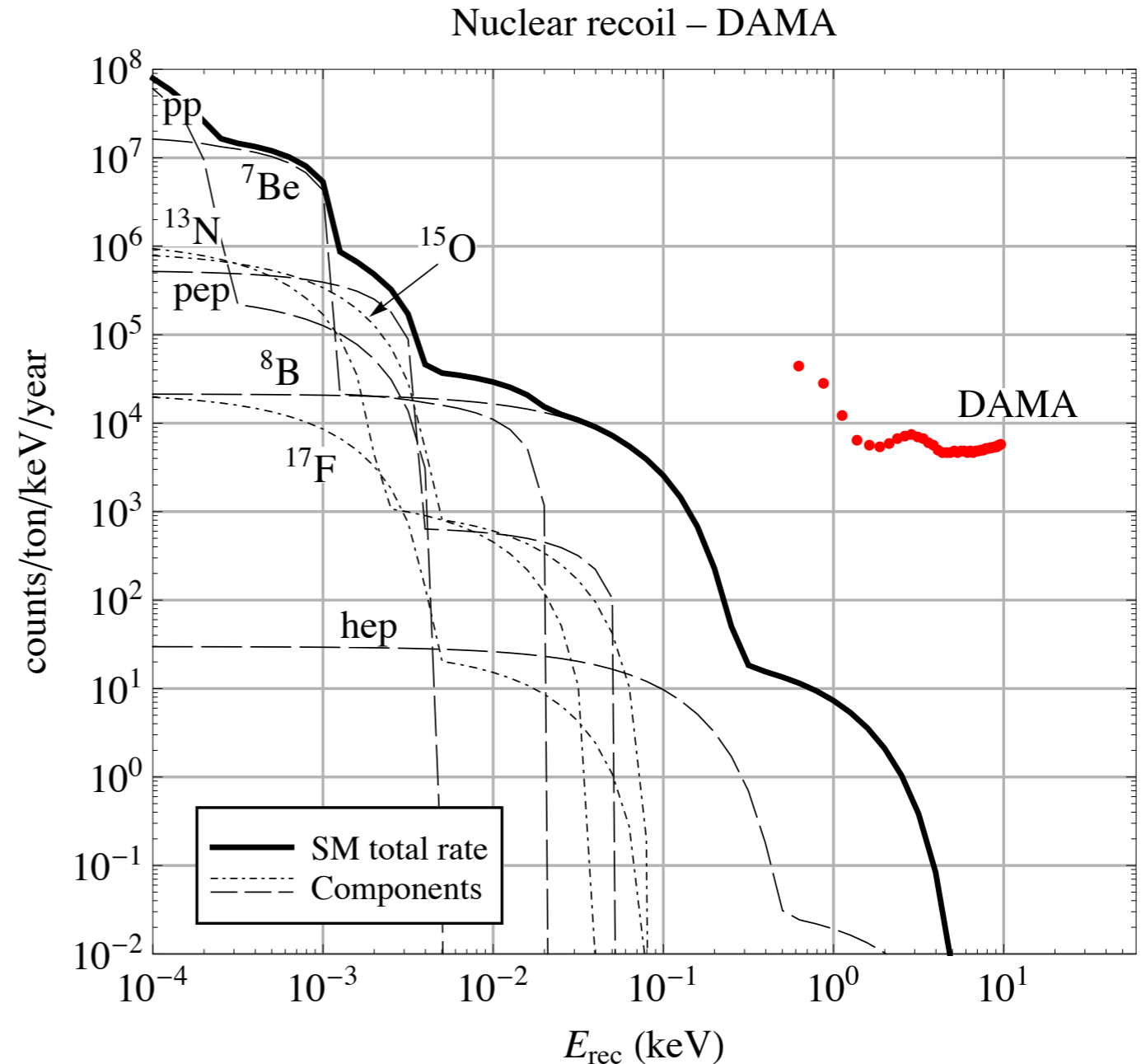
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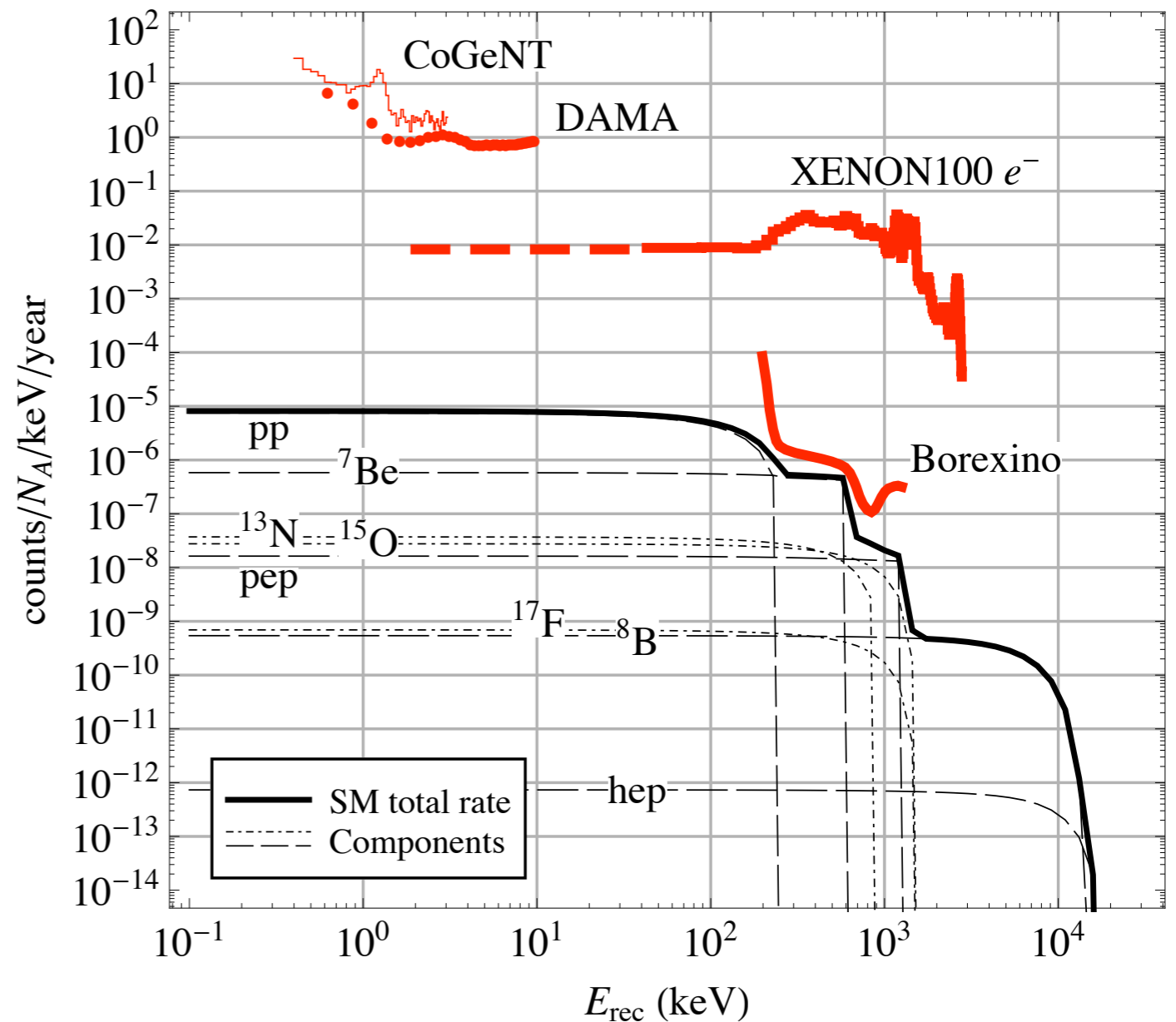
Spectrum

* Electron recoil:

- pp neutrinos are above threshold. High flux.

$$E_r^{\max} = \frac{2E_\nu^2}{m_T + 2E_\nu}$$

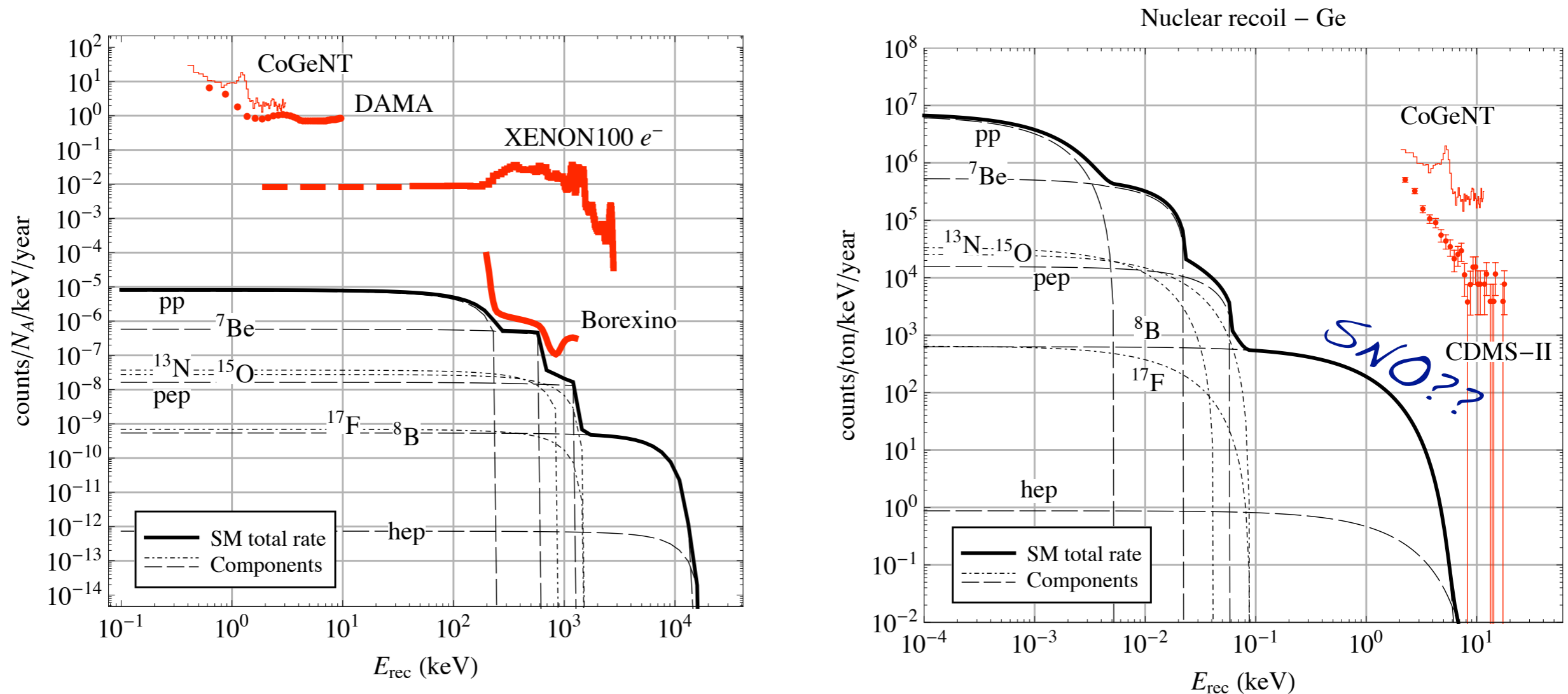
- But, lower cross section, (phase space and coherence).



*scattering on effectively free electrons.
All experiments in one plot!*

New Physics?

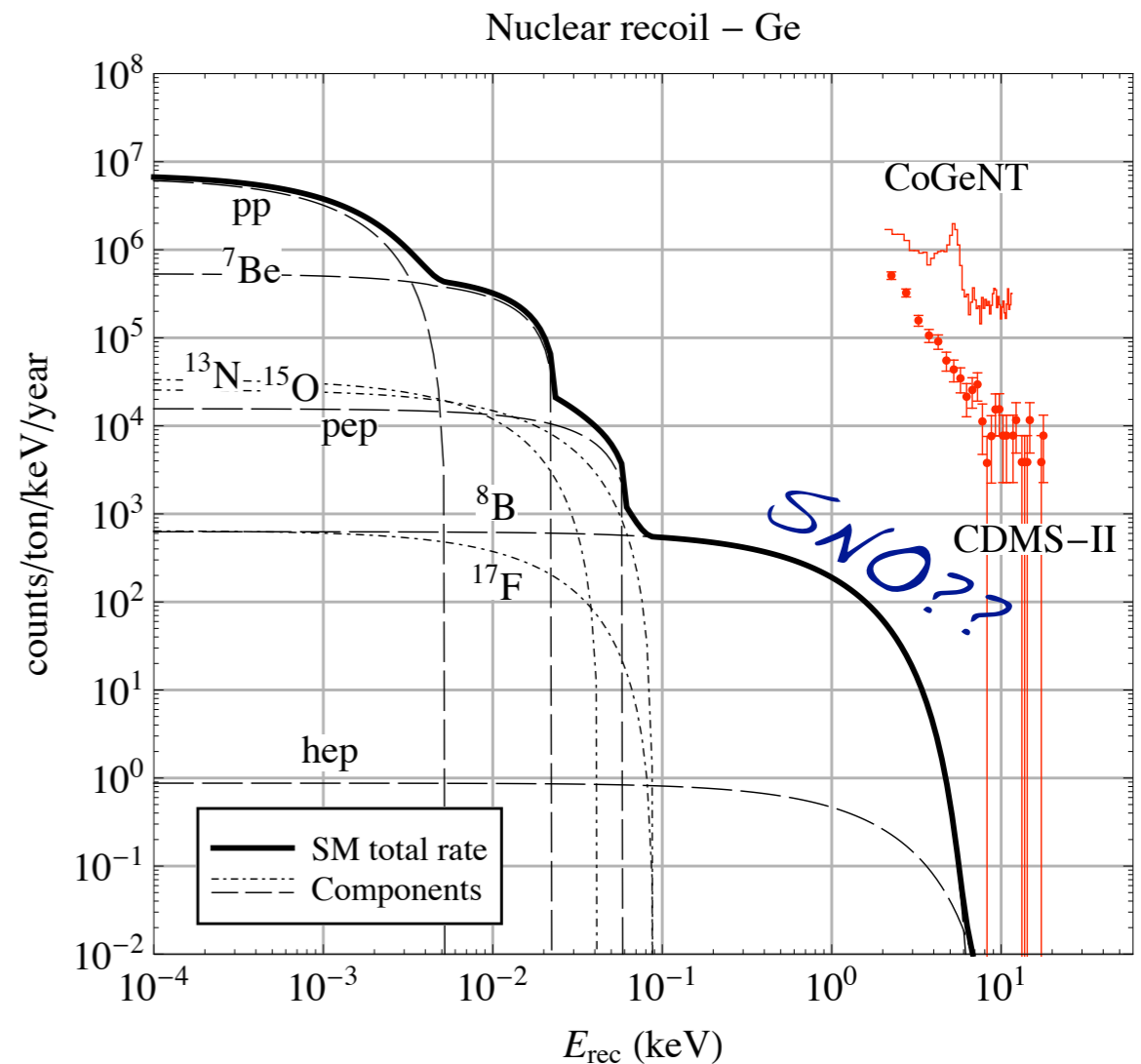
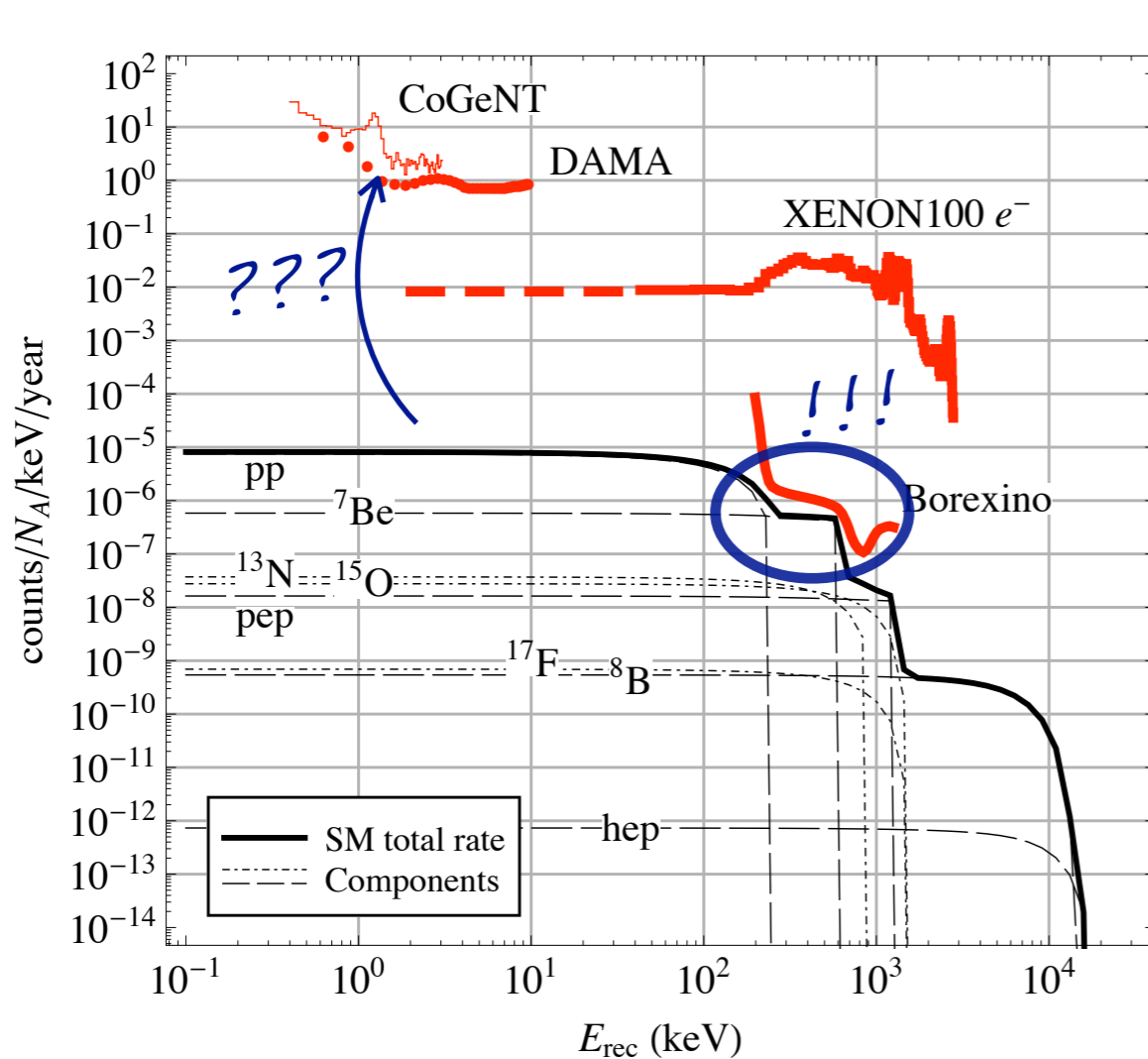
- * Can new physics in the neutrino sector “raise” this background and give interesting signals?



Since Joseph cover nuclear, I'll focus on e-recoil.

New Physics?

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Outline

- * New light gauge bosons (A')
- * Electron recoil
 - o SM neutrinos or light steriles
 - o Heavy steriles
- * Potential sources of modulation-
 - o Solar distance
 - o Matter effects
 - o Channeling
- * Constraints on A' 's
- * Conclusion.

New Physics Models

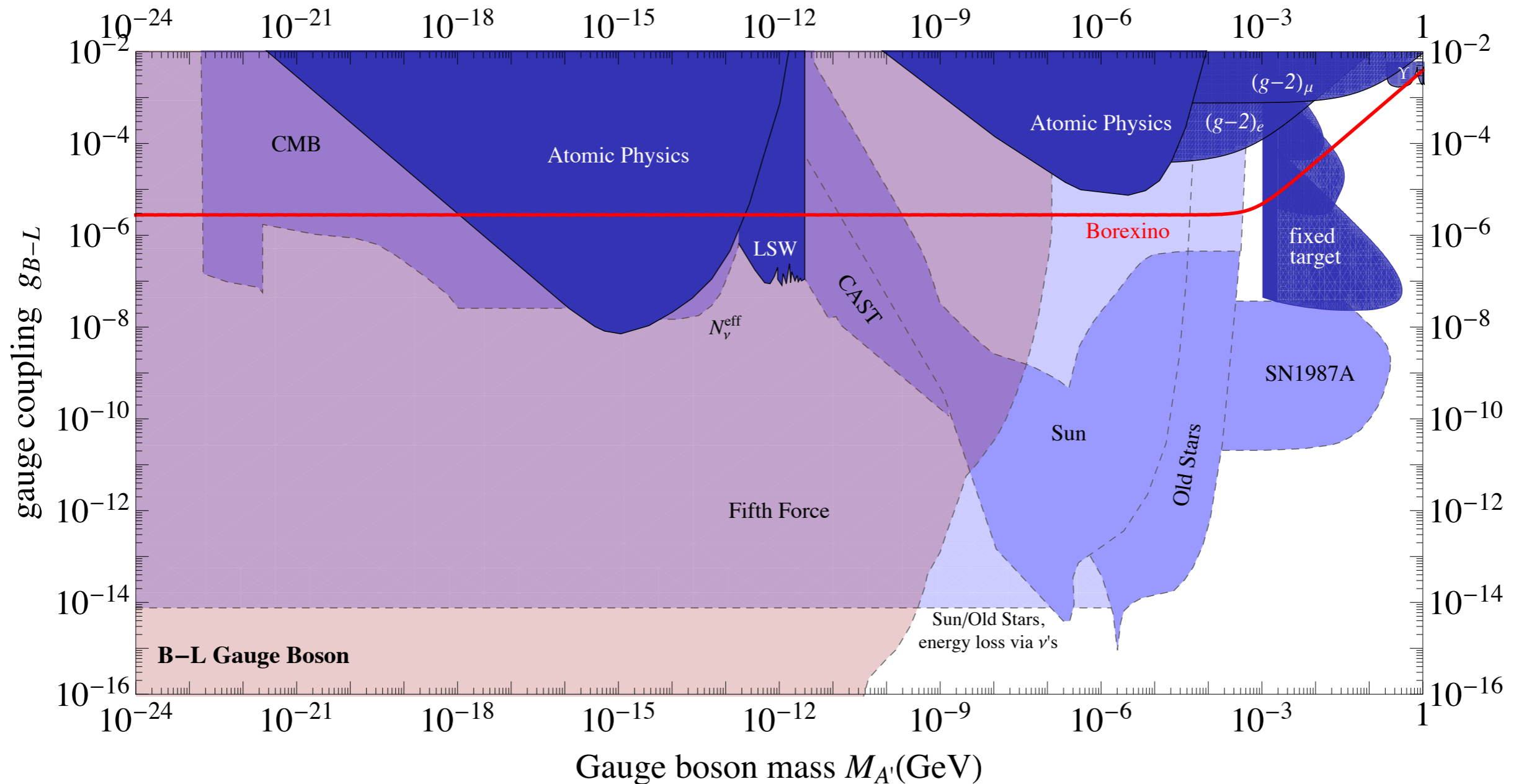
- * Many models with new light gauge boson:
 - o A light **B-L** gauge boson
 - o Kinetically mixed U(1) (a.k.a **dark photon**).
- * New **sterile** neutrinos can also come in handy.
Can be emitted by the sun via mixing or oscillation.

-
- * Another “new” gauge boson that can couple to neutrinos is the **photon**.

a magnetic dipole moment: $\mu_\nu \bar{\nu} \sigma^{\mu\nu} \nu F_{\mu\nu}$

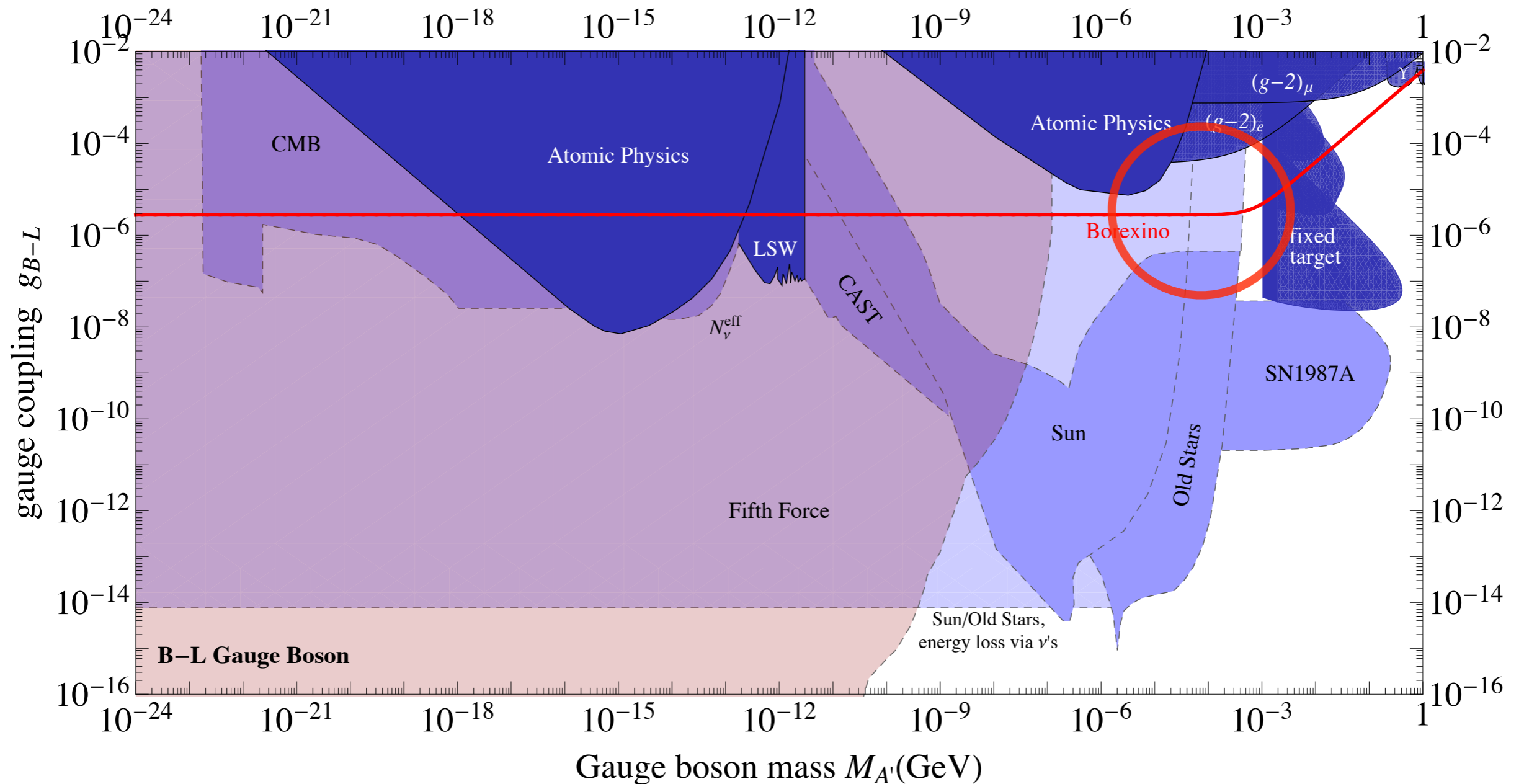
Parameter Space

* B-L is constrained. There are holes:



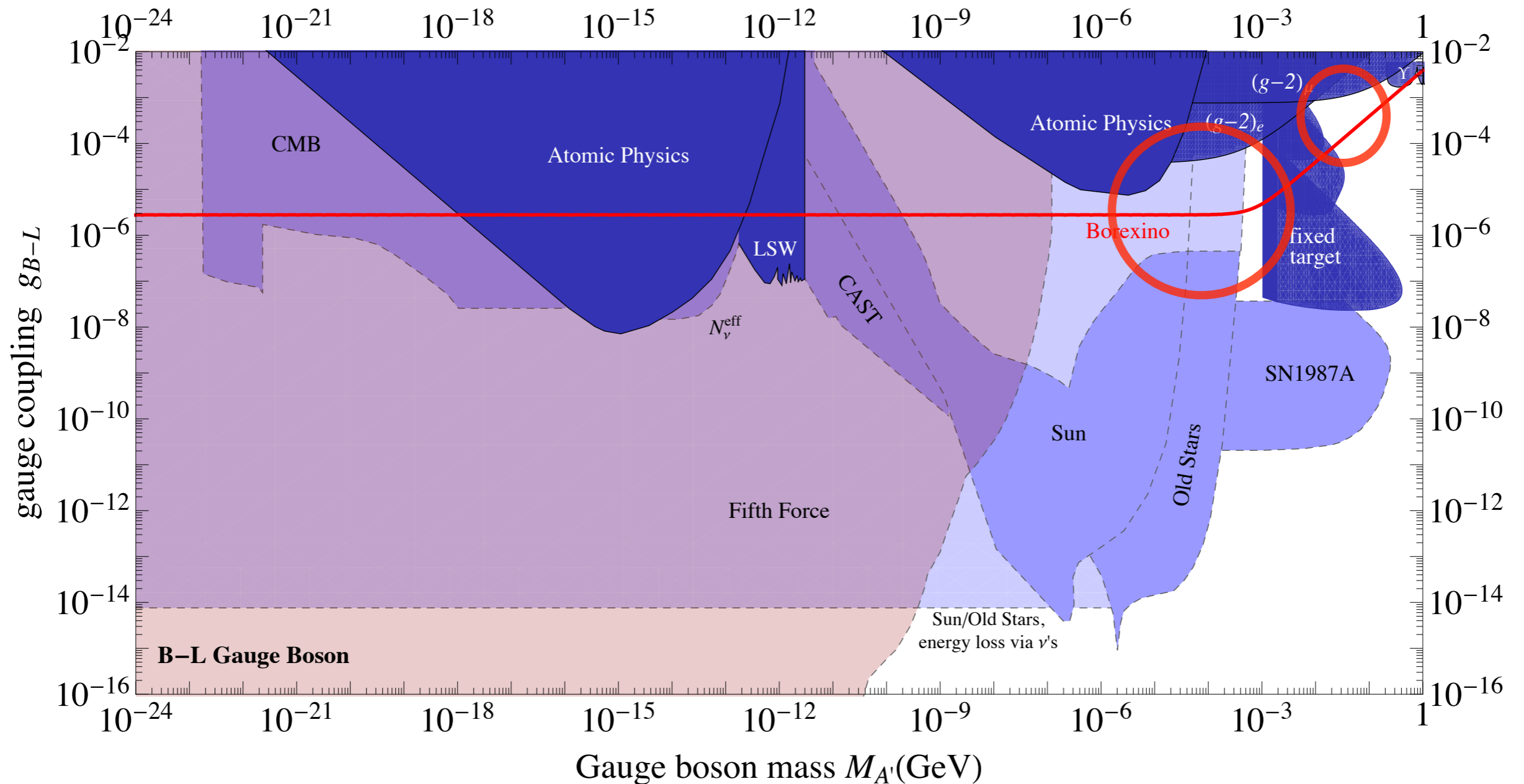
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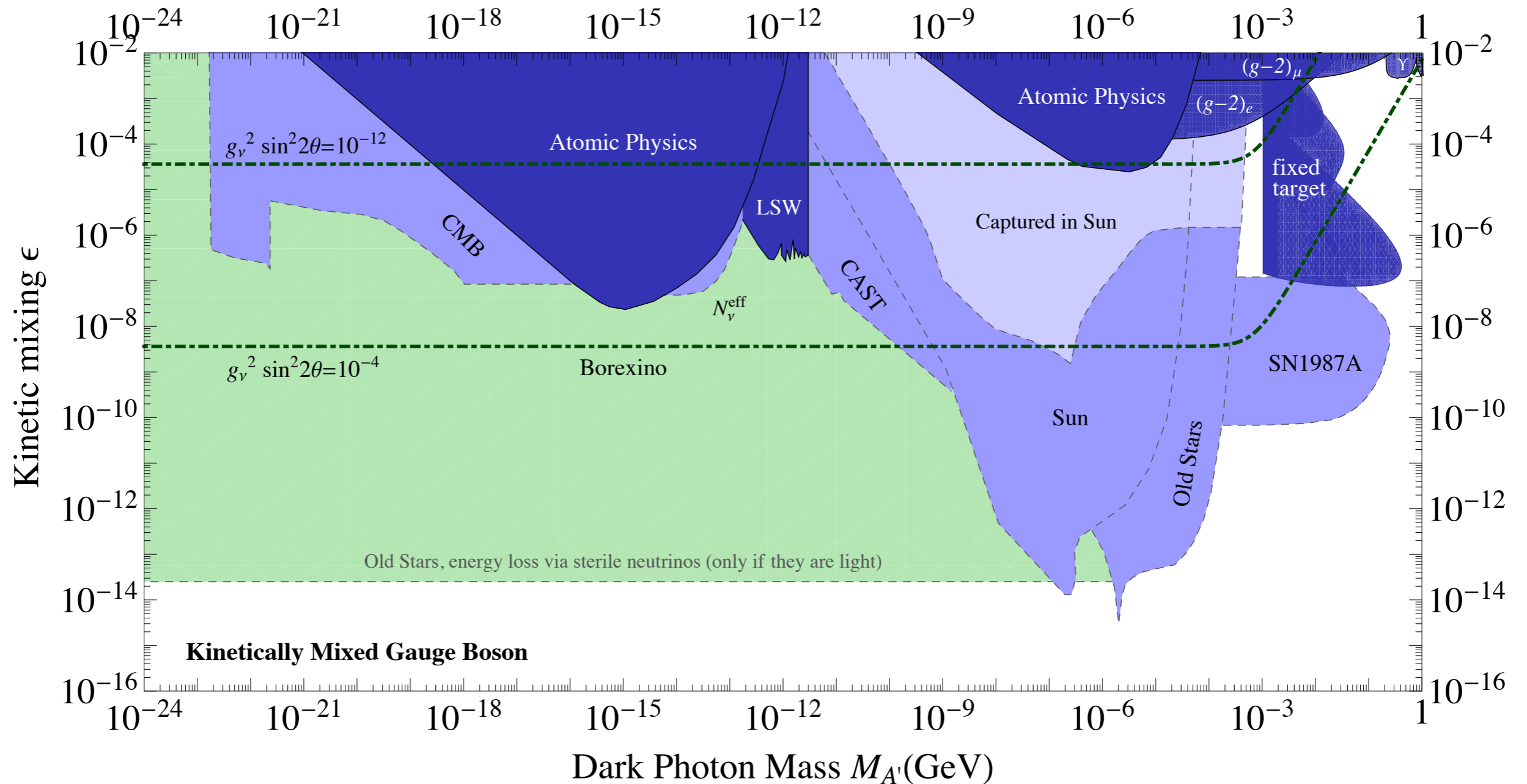
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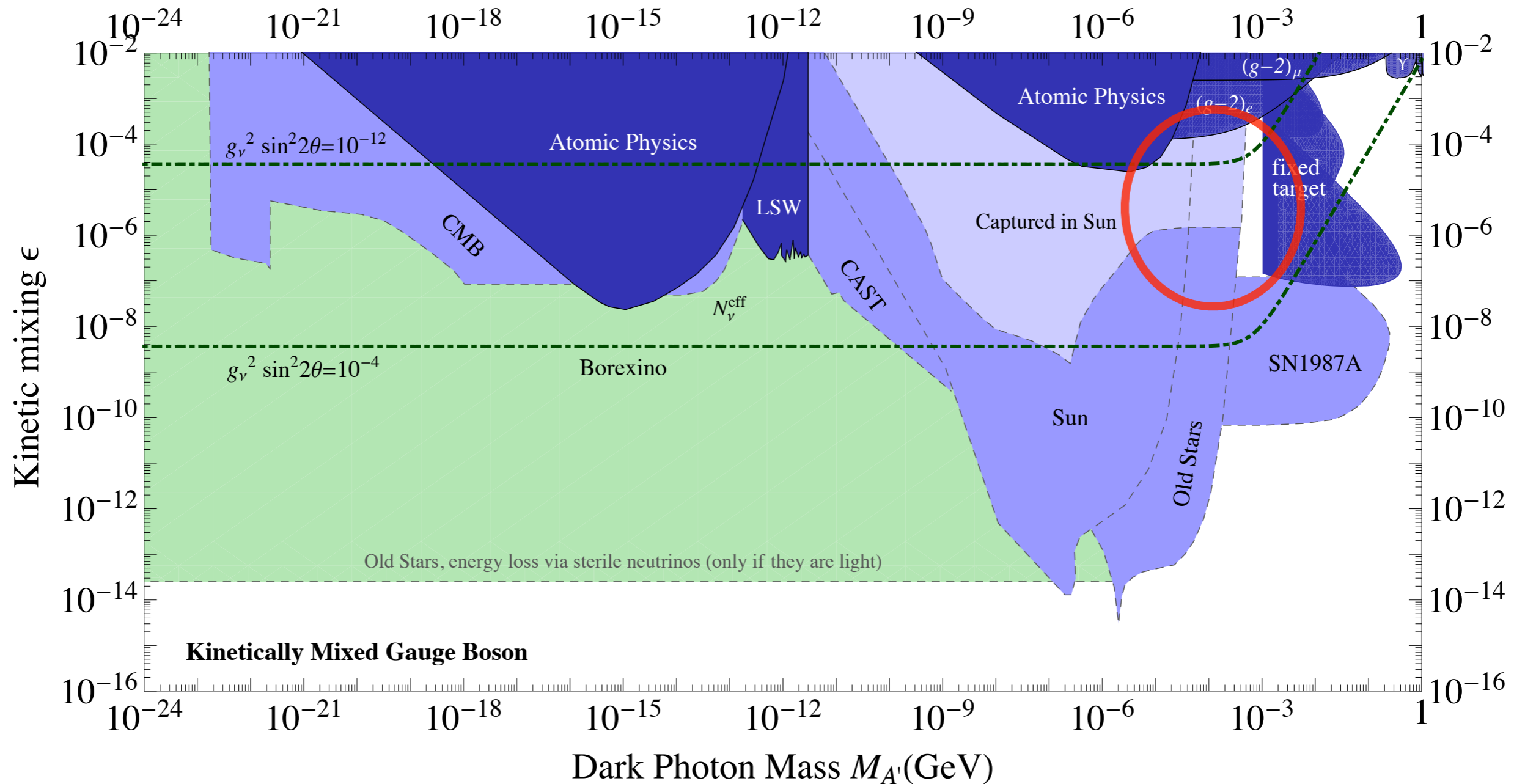
Parameter Space

* Dark photons are a bit more open.



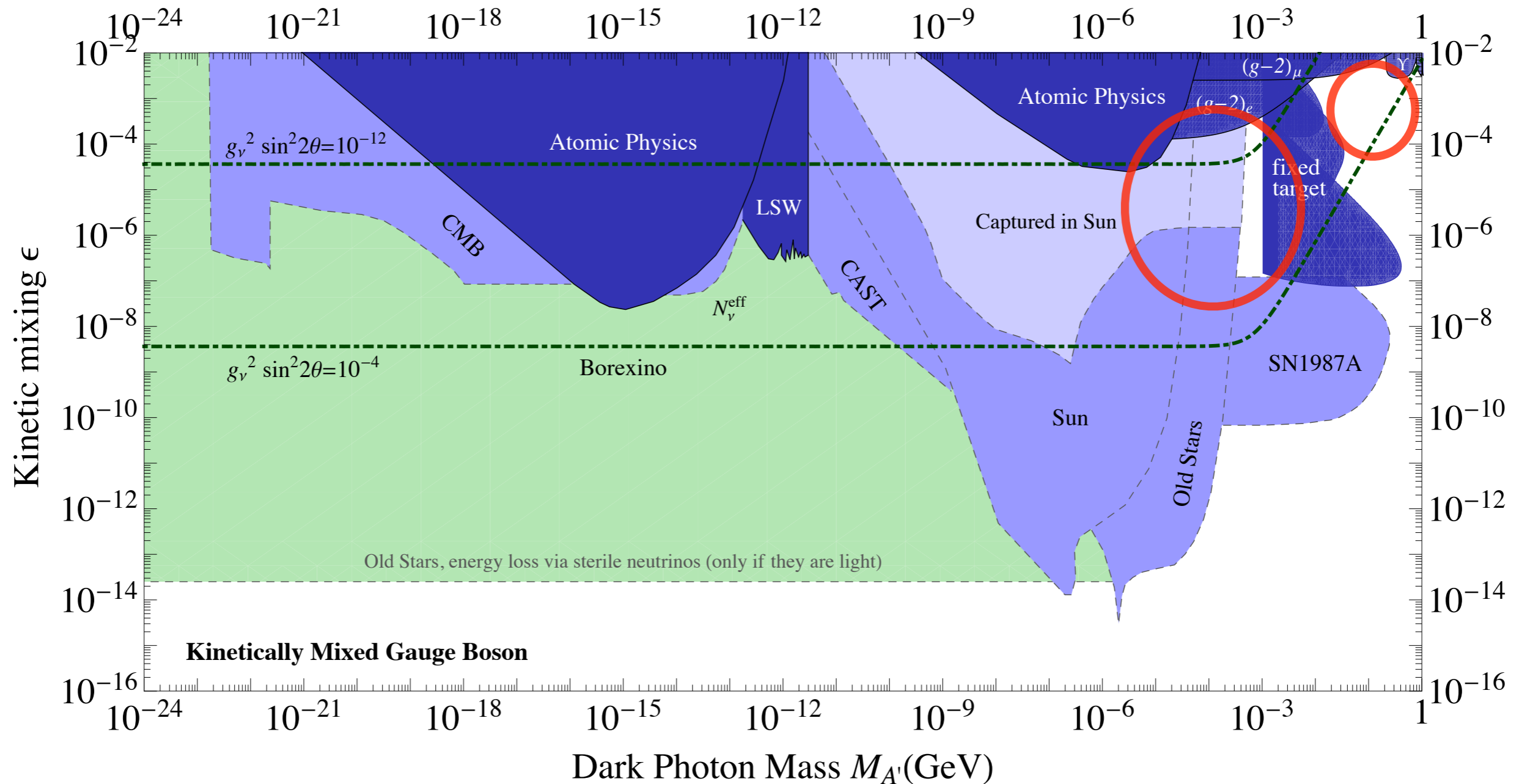
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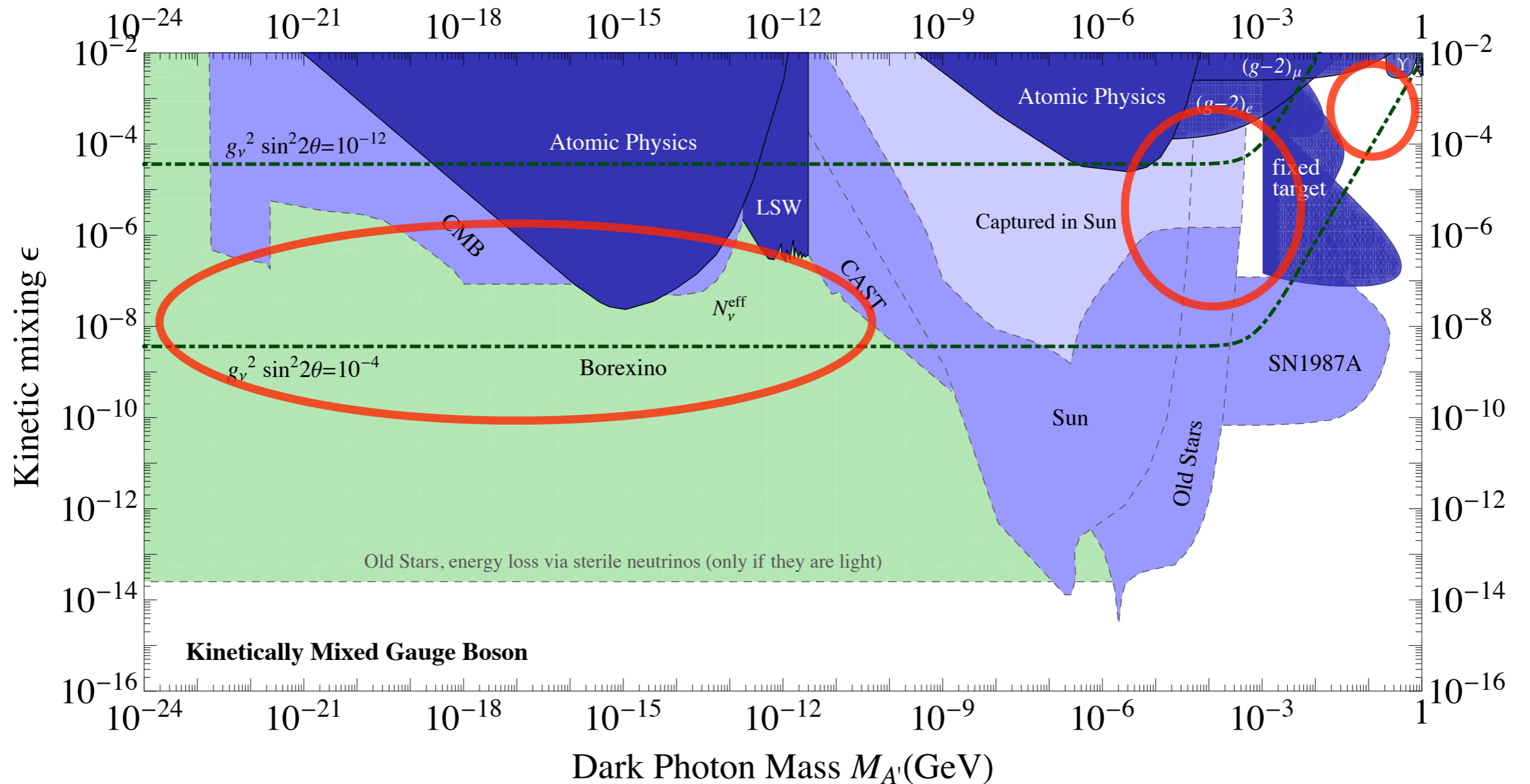
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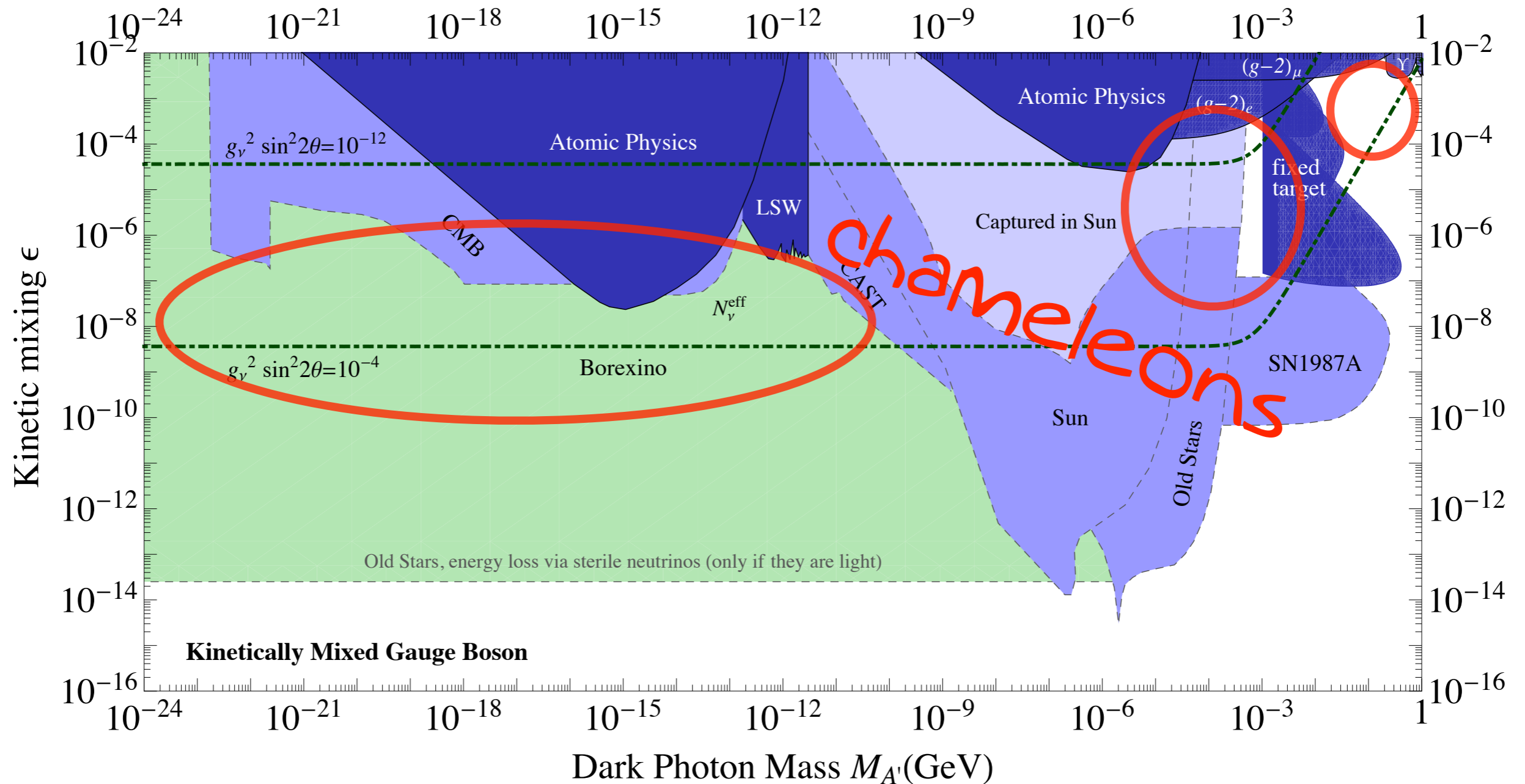
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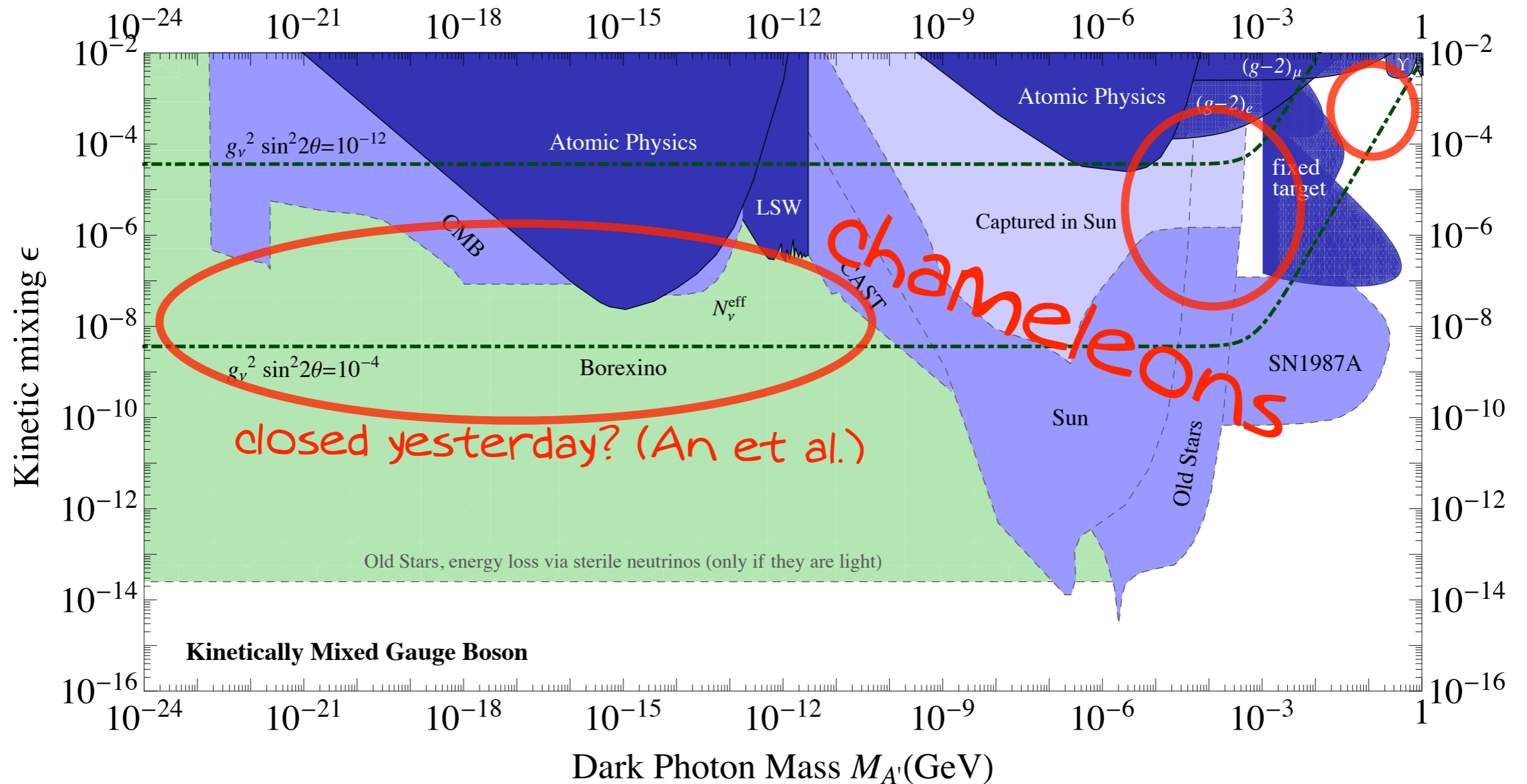
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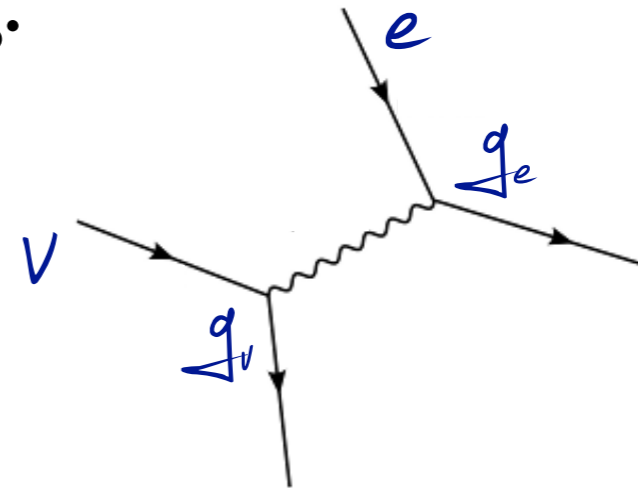
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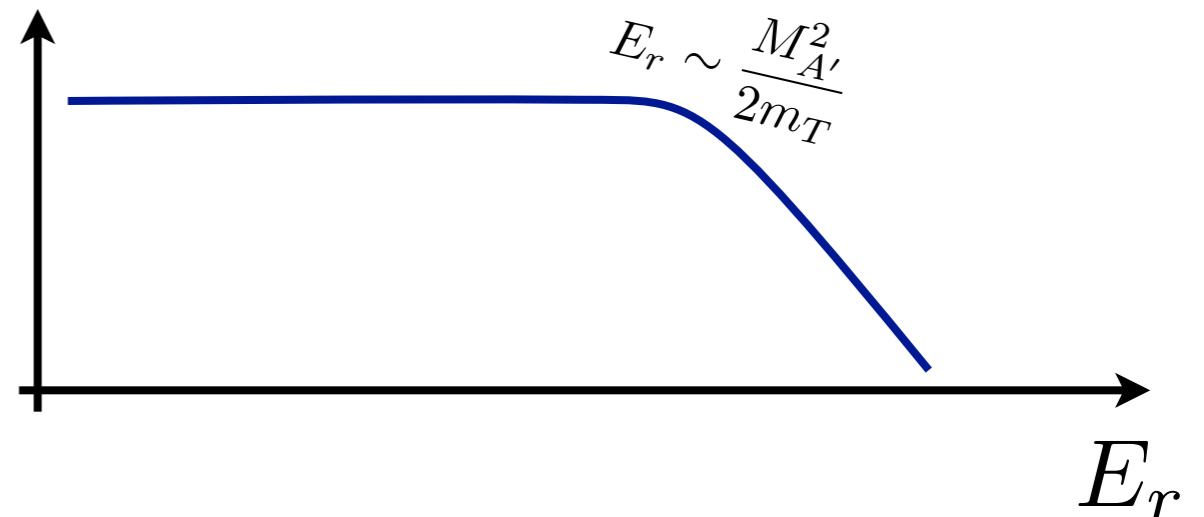
A Light Gauge Boson

- * The exchange of A' contributes to neutrino-nucleus and neutrino-electron scattering:



$$\frac{d\sigma}{dE_r} = \frac{g_\nu^2 g_T^2 m_T}{4\pi p_\nu^2 (M_{A'}^2 + 2E_r m_T)^2} [2E_\nu^2 + E_r^2 - 2E_r E_\nu - E_r m_T - m_\nu^2]$$

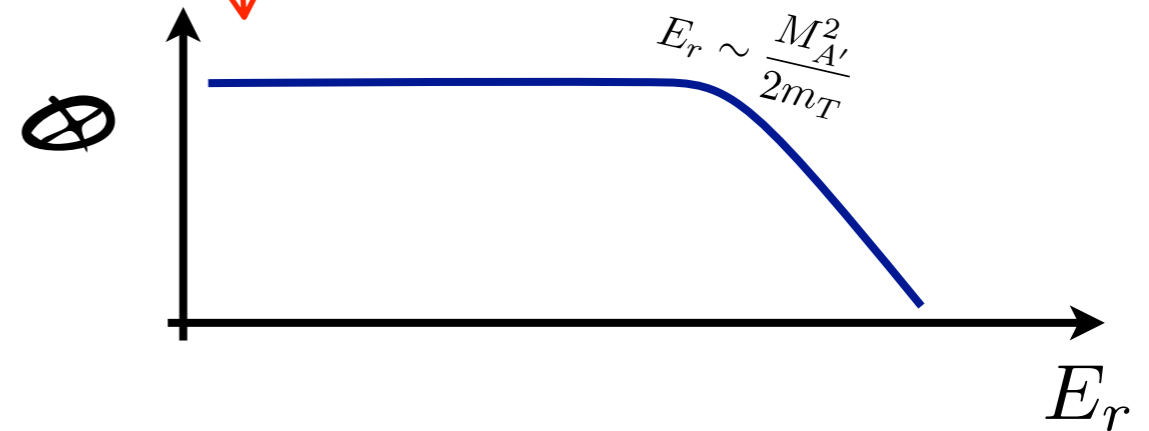
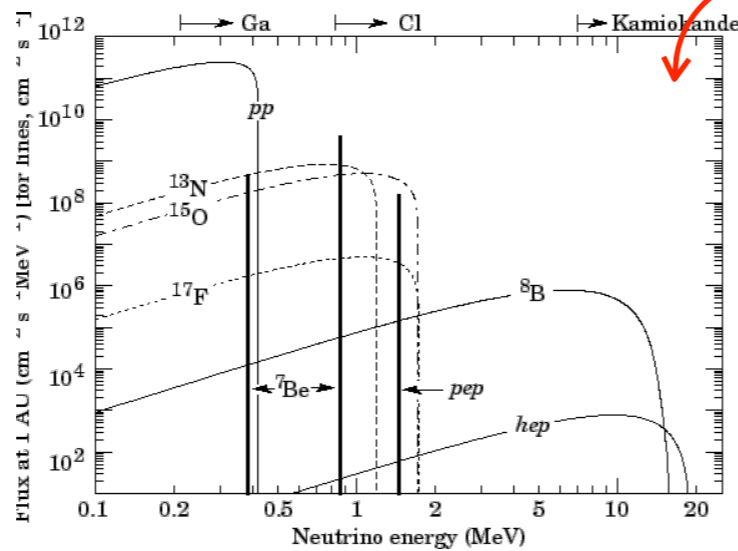
propagator: $q^2 - M^2$



A Light Gauge Boson

- * What does a direct detection experiment see?
back to the master formula:

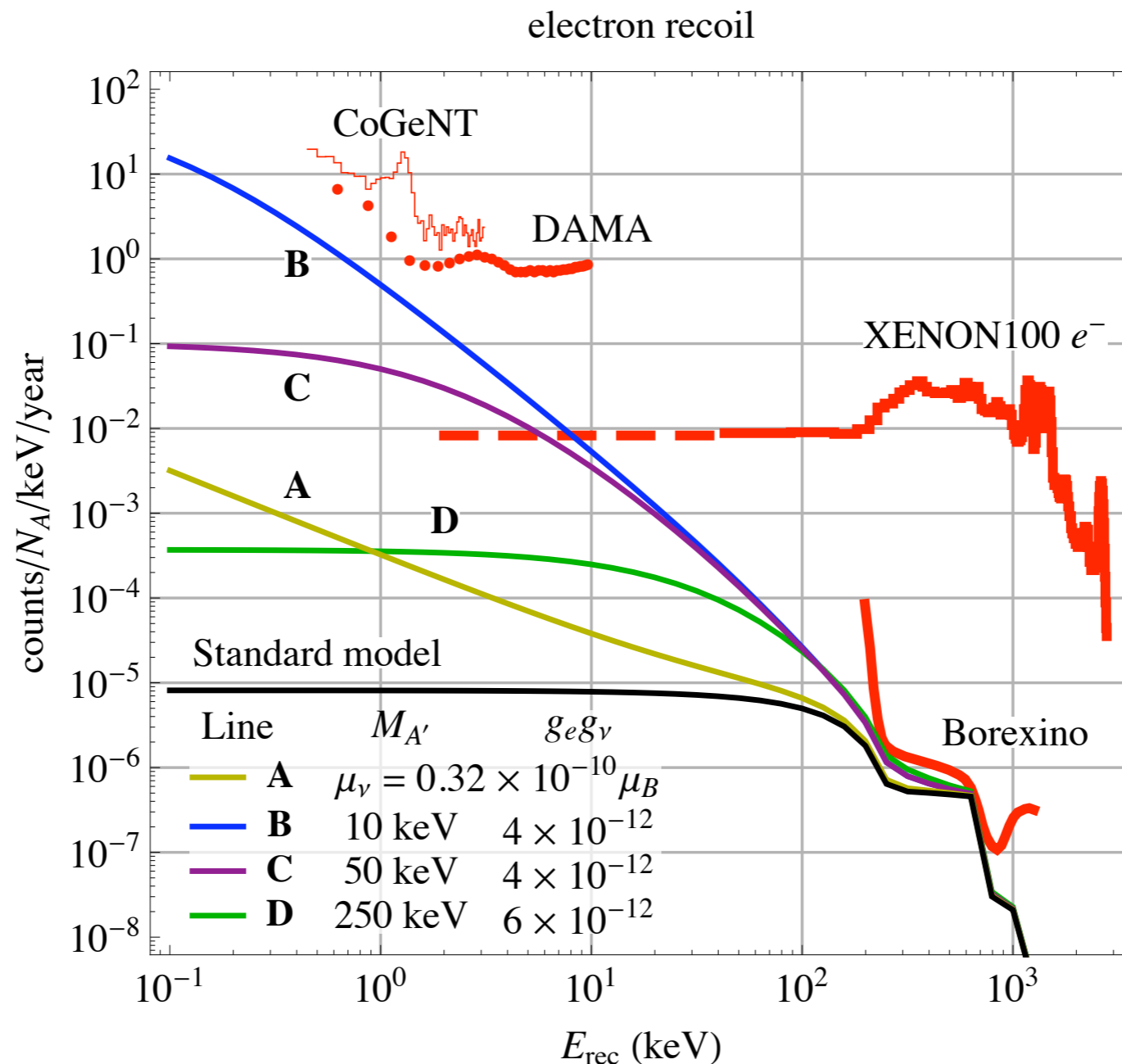
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with
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CoGeNT or DAMA?

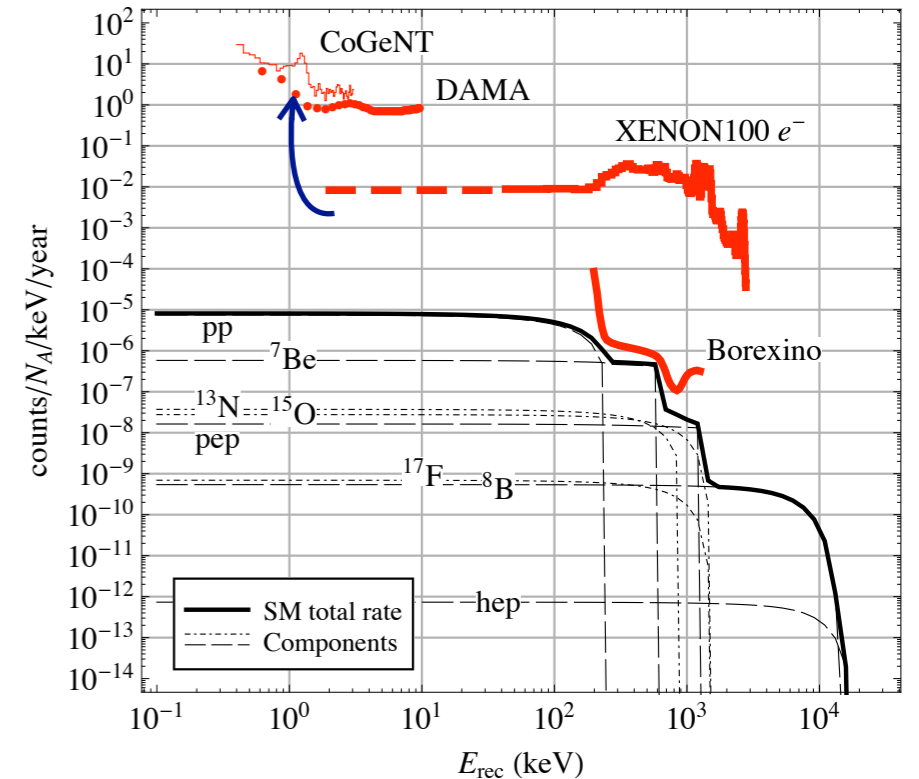
- * If we want to do CoGent or DAMA there is clear tension with XENON100 (and also Borexino).



Heavy Sterile

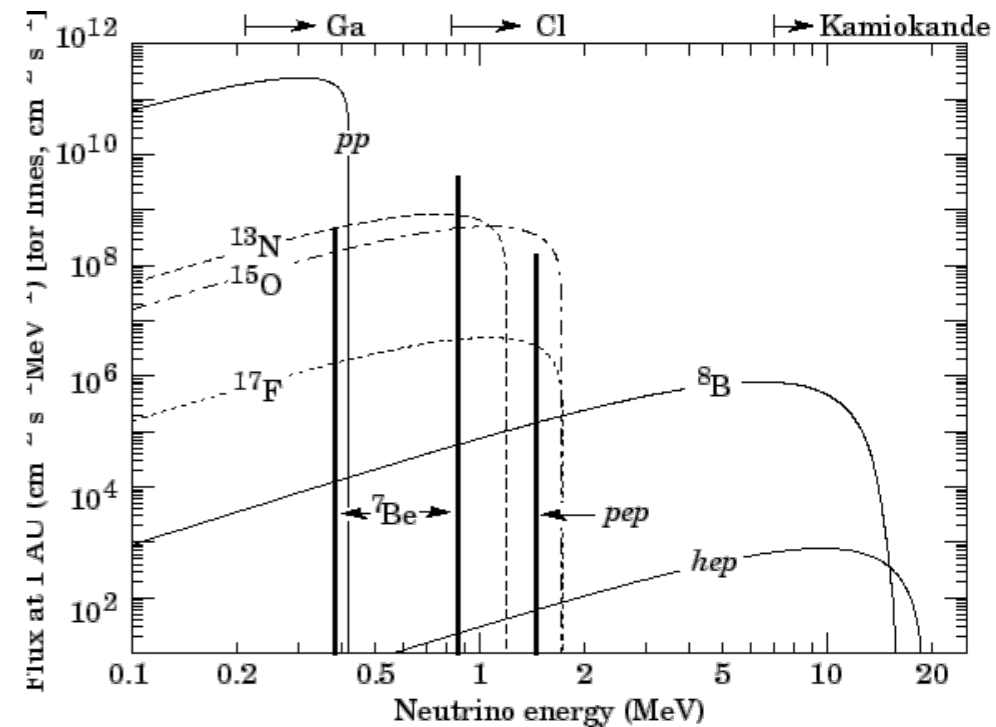
* Is there a way around XENON?

Can we get a sharp threshold?



* The solar flux has lines....

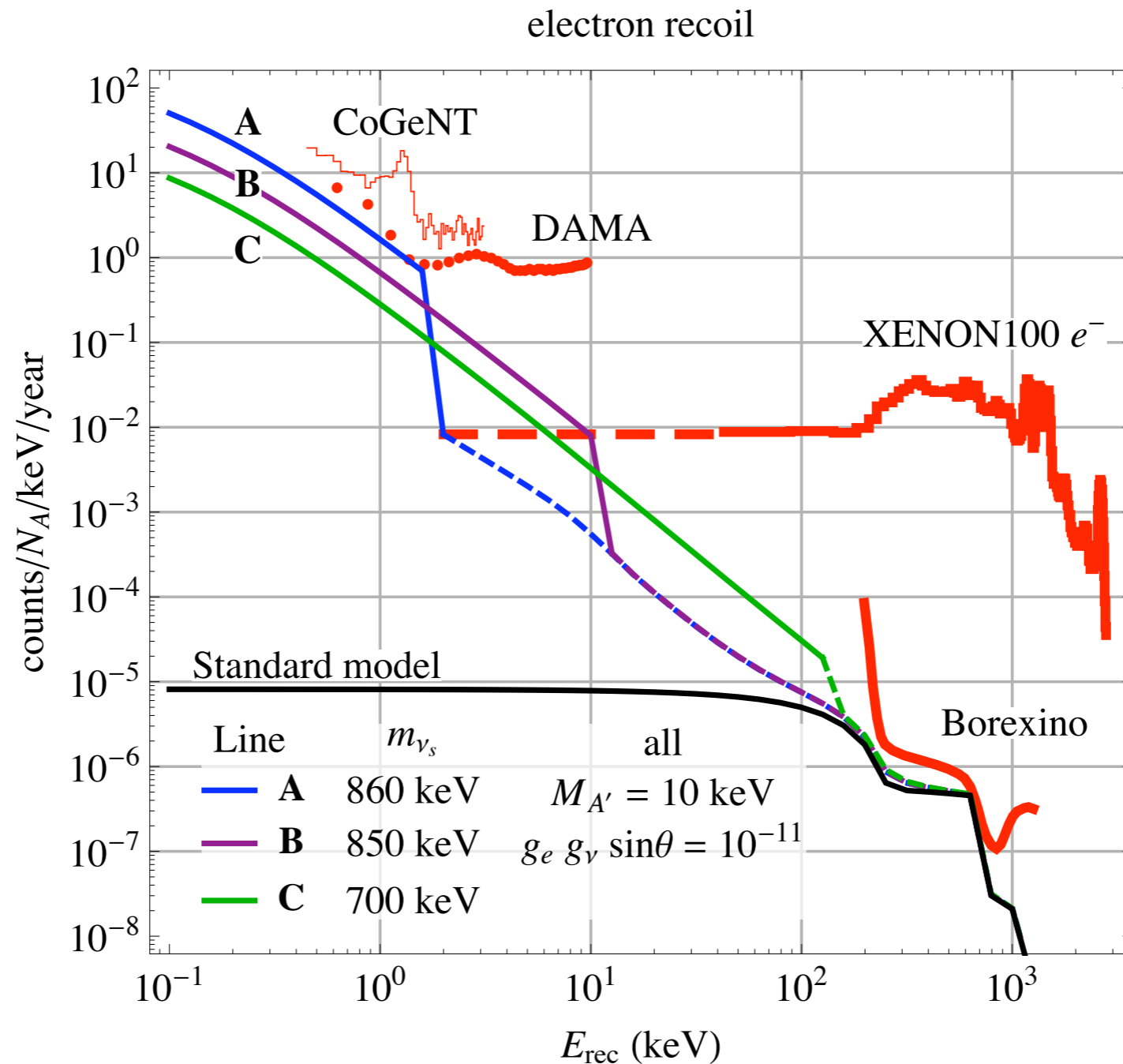
*A kinematic threshold near the ⁷Be line.
It's tuned...**



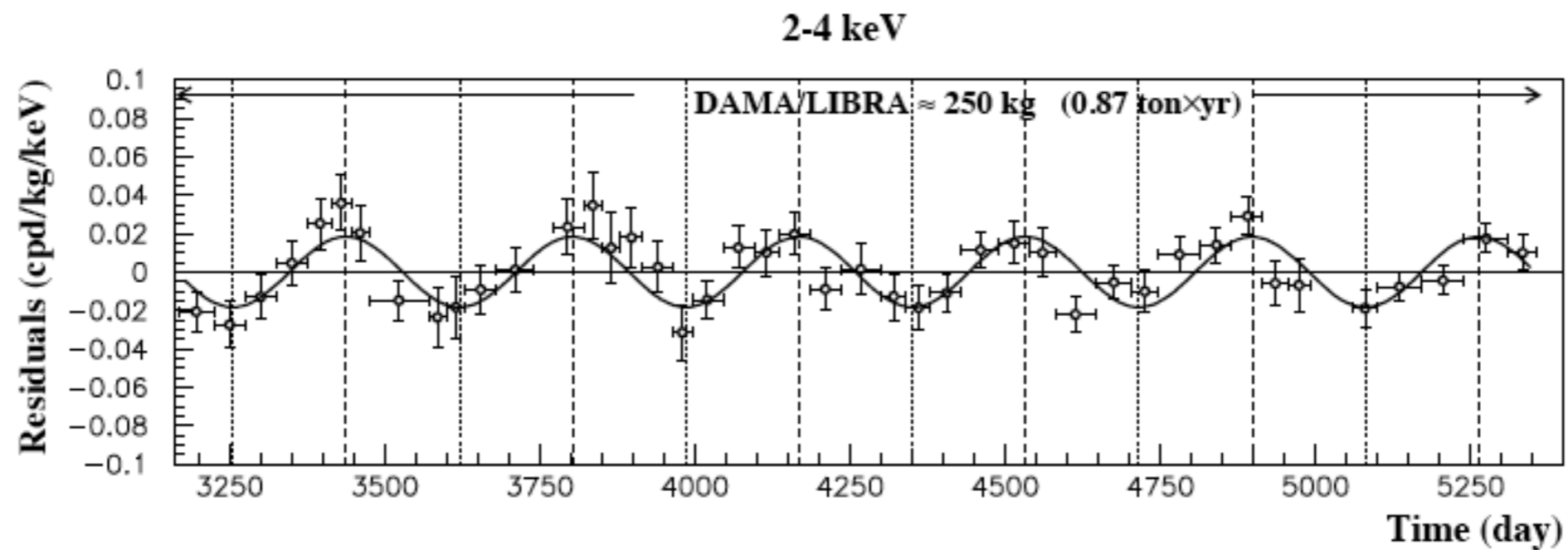
* I would like to thank IDM and its variants for the moral license to do this.

Heavy Sterile

* Interesting spectra can arise:



Modulation



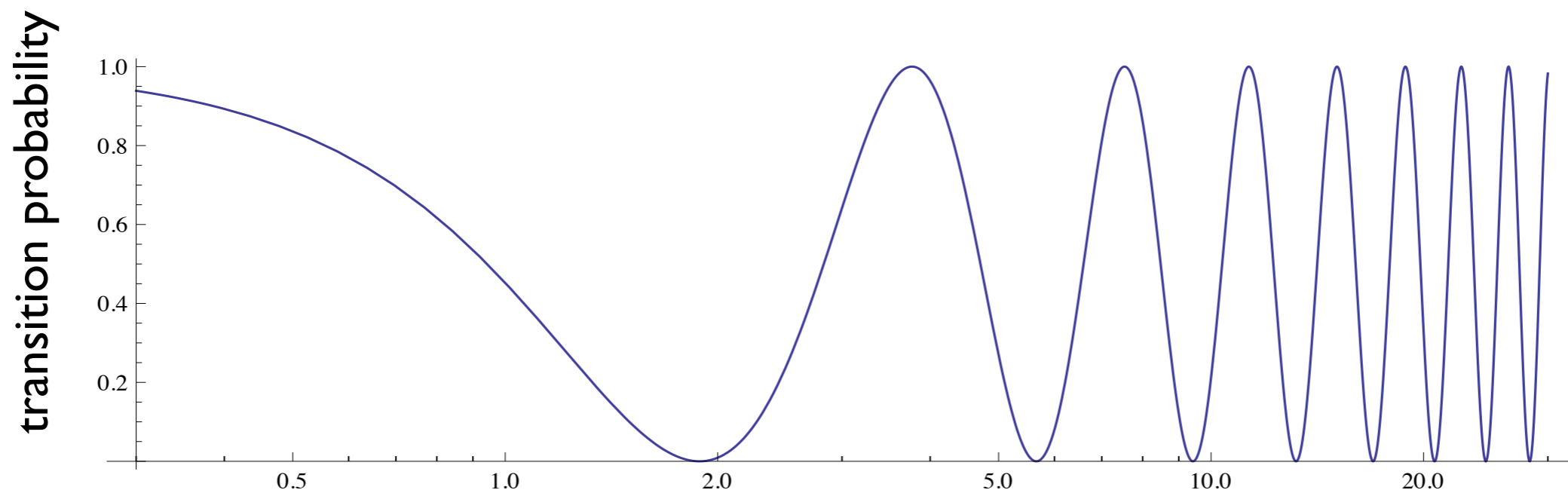
Modulation and the Sun

- * Annual modulation is an important part of past and (hopefully) future anomalies.
- * What modulation signals can the Sun produce?
- * **Many possibilities for modulation:**
 - daily.
 - annually.
 - even semi-annually.

Elliptical Orbit & Just-So

- * We are on an elliptical orbit. Wikipedia:
 - o Closest to the Sun on **Jan 3rd** (wrong phase for DAMA).
 - o The amplitude is 1.6% (flux is double that).
- * Introduce oscillation on AU scale (with sterile):

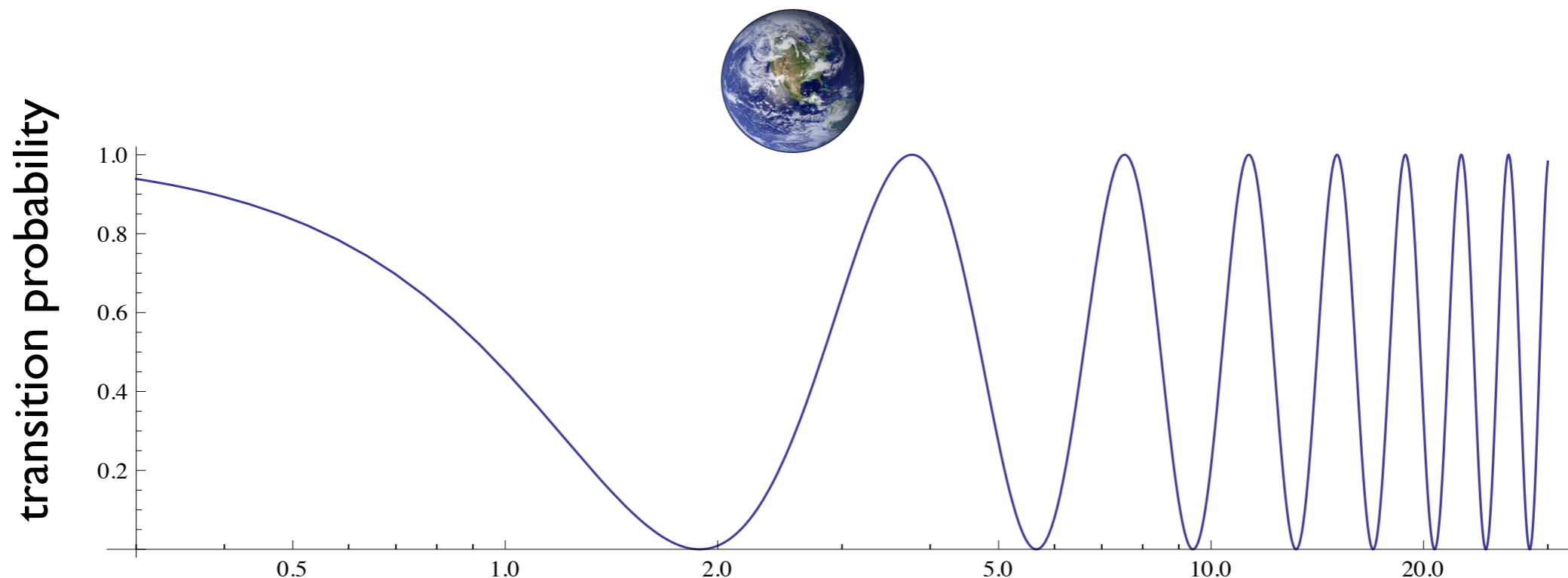
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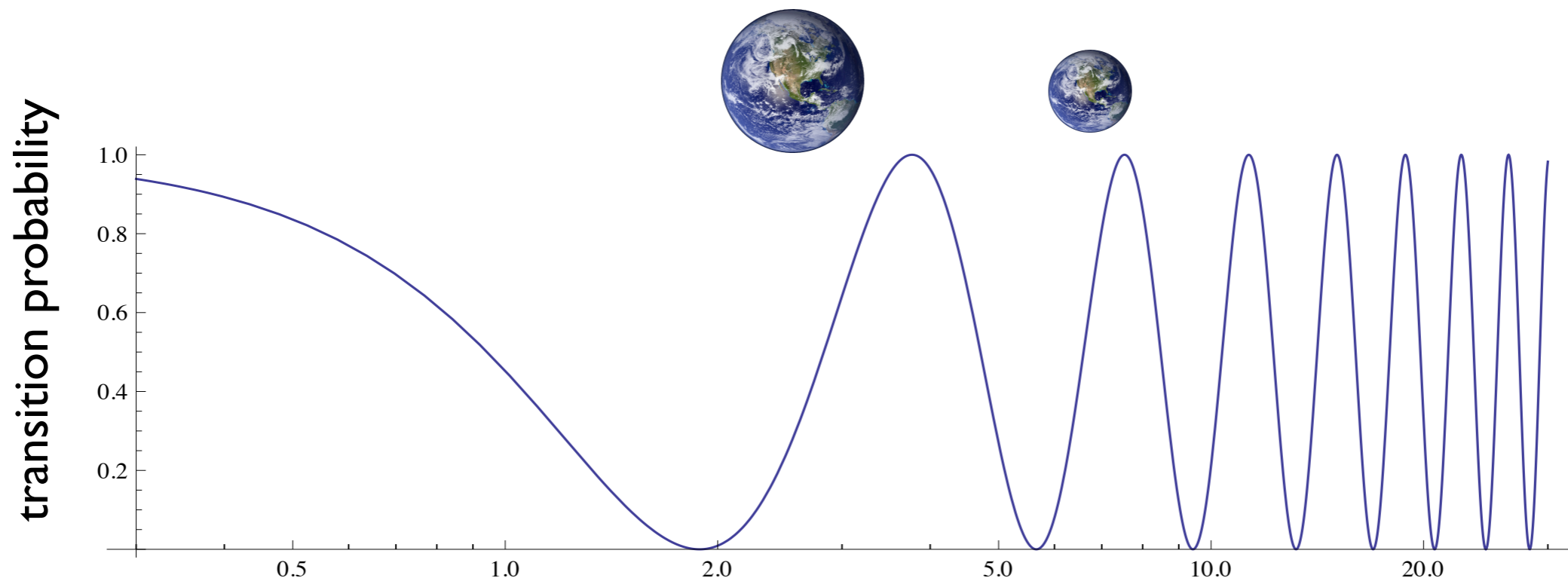
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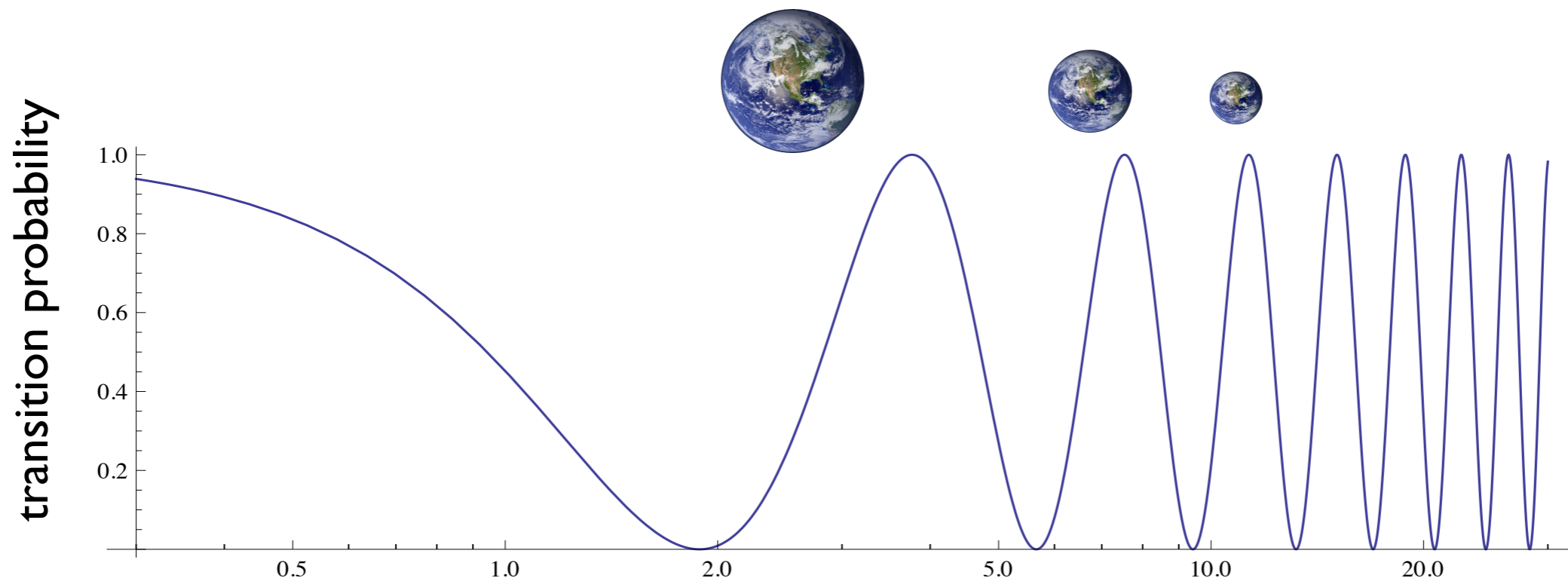
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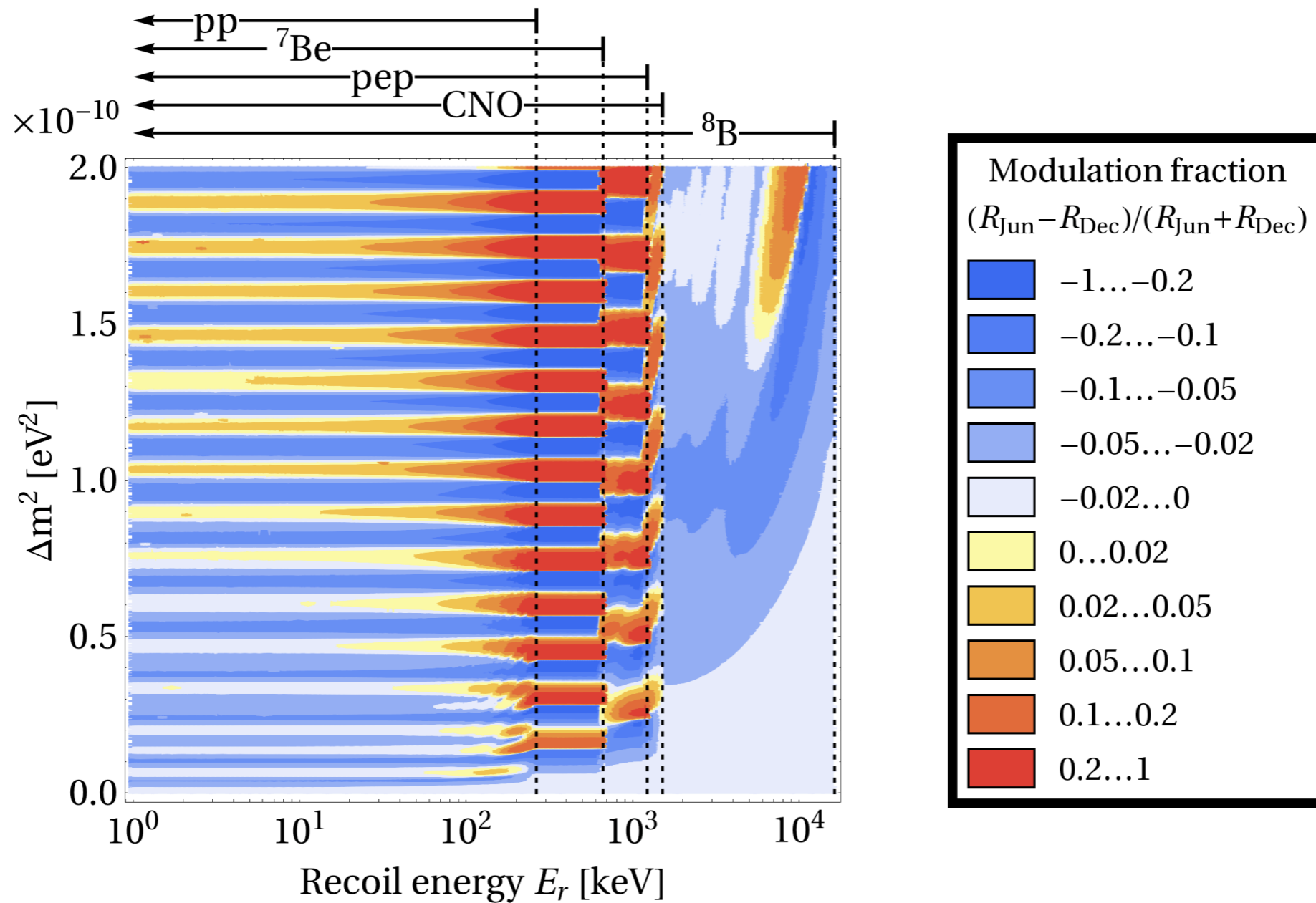
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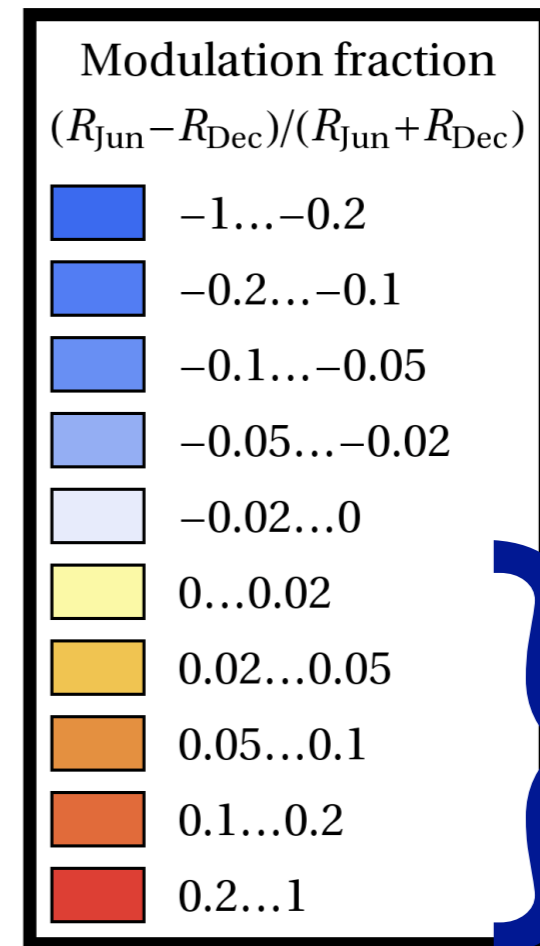
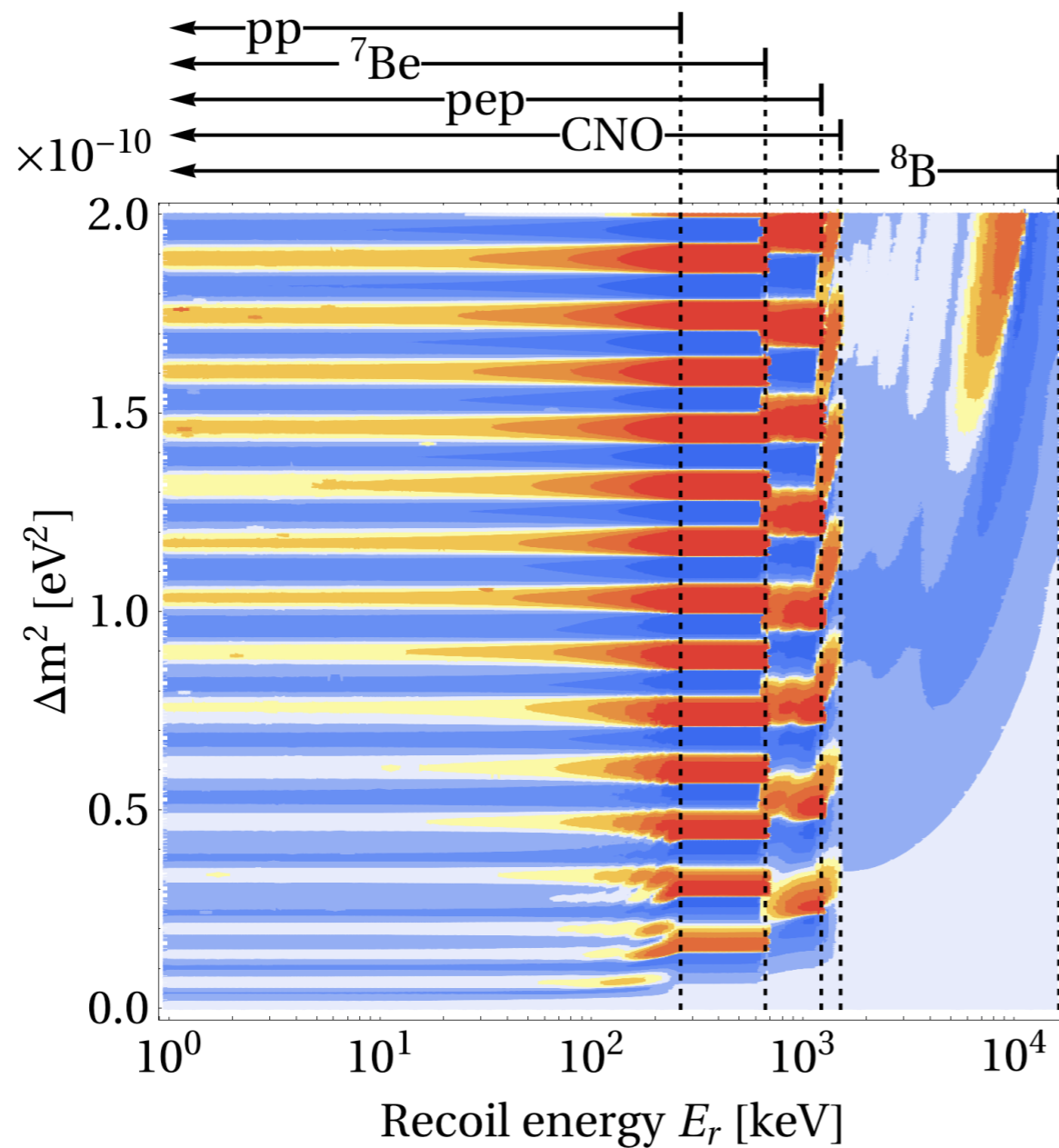
Just-So

- * A variety of modulation amplitudes are possible



Just-So

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DAMA

Matter Effects

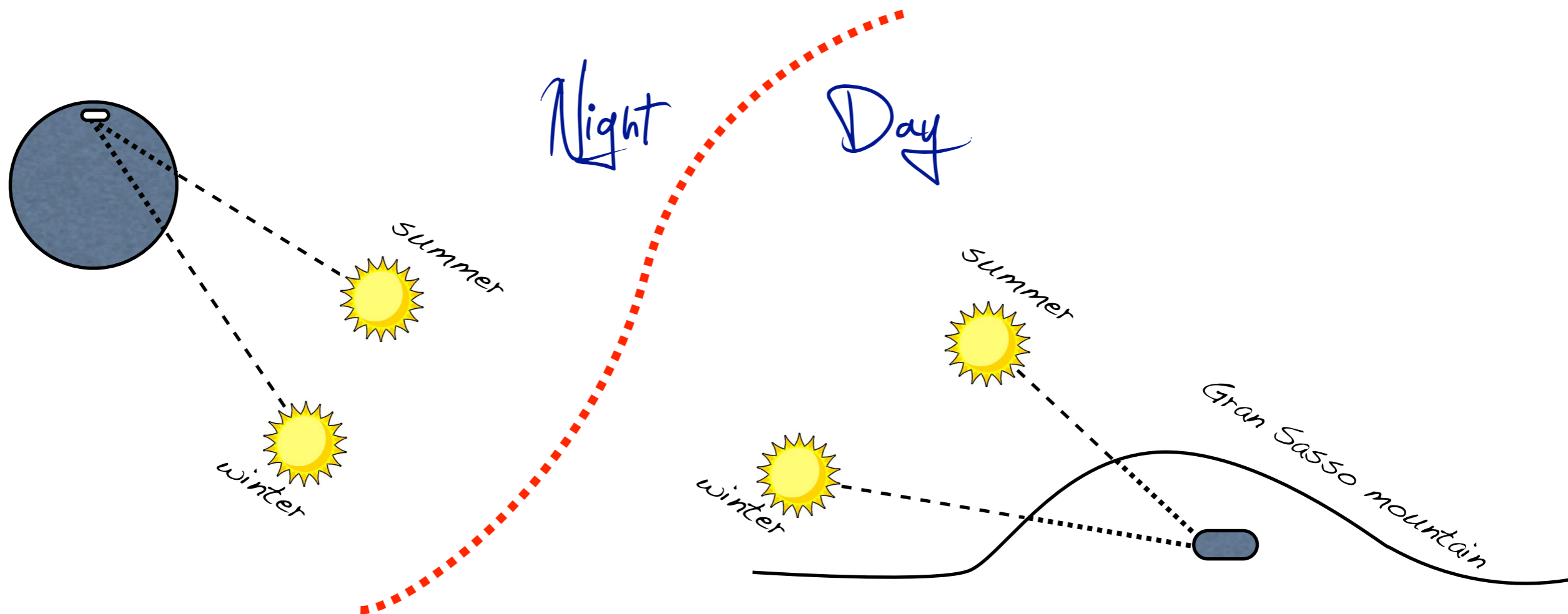
- * The new gauge boson can lead to new “MSW-like” matter effects:

$$V_{\text{matter}} = \frac{g_\nu}{M_{A'}^2} (g_e n_e + g_p n_p + g_n n_n)$$

- * Active-sterile oscillations in matter can be very different from those in vacuum.
- * **Day-night asymmetry** due to an oscillation b/w among sterile species in matter. This asymmetry can be large.
- * The matter oscillation length, $L_{\text{osc}} = 4\pi E / \Delta m^2$, can be anywhere between a kilometer and the earth radius.

Zenith Angles

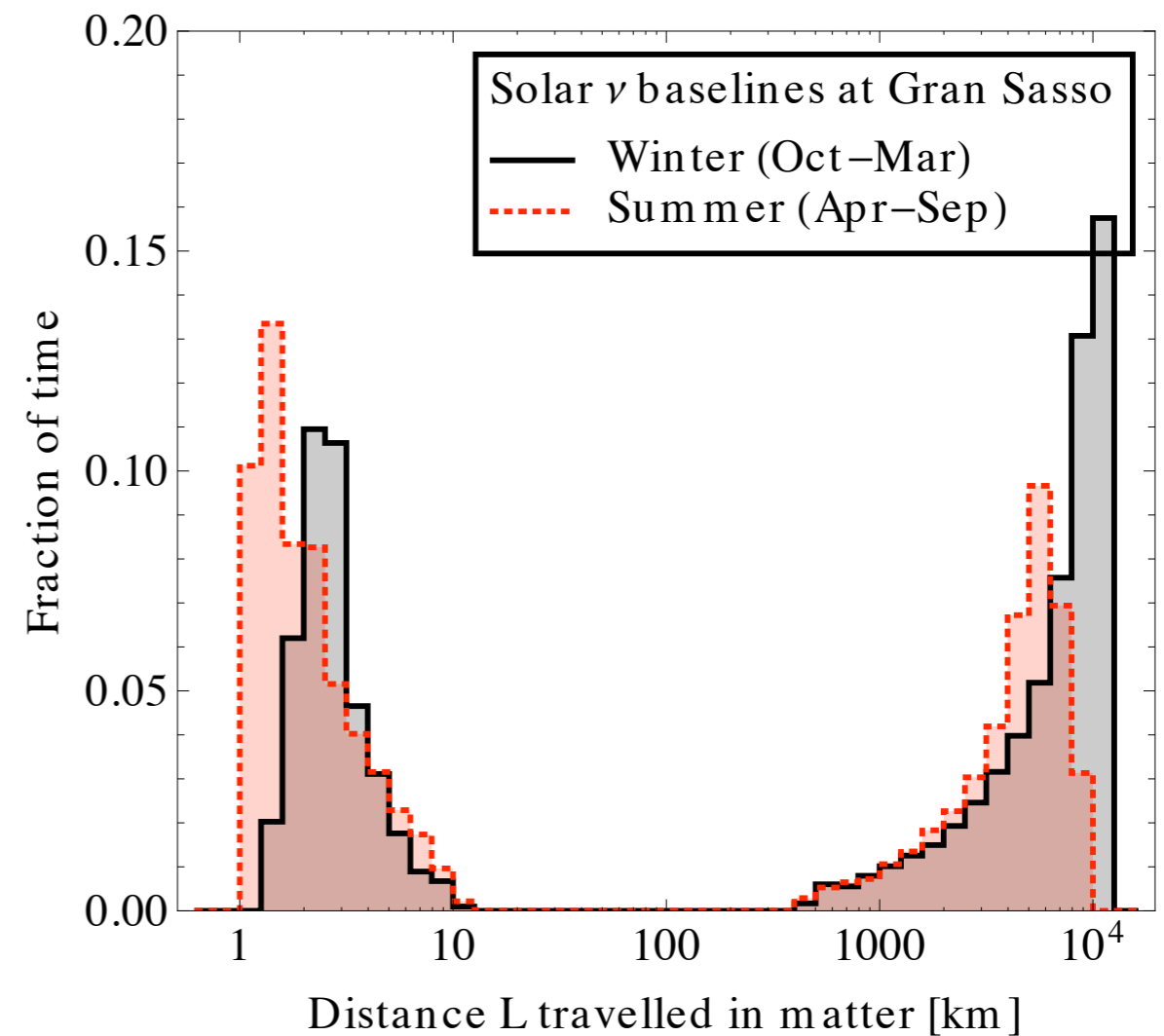
- * At noon, the sun is high in the sky in summer. Low in winter.
- * At midnight, the Sun is lower *below* the horizon in winter. Higher in summer.



Zenith Angle

- * The average baseline in rock for solar neutrinos going to Gran Sasso **modulates**:

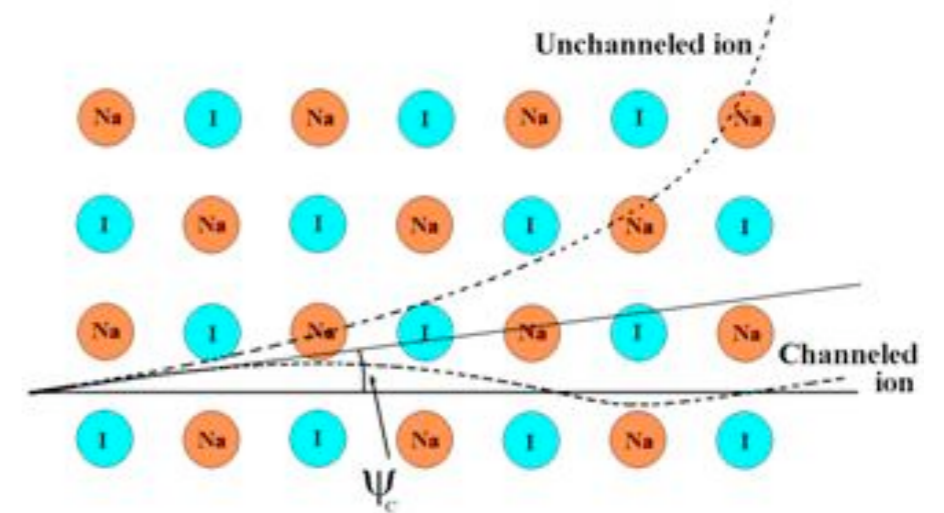
*annual modulation
can be much
stronger asymmetry
in daylight hours.*



- * A strong daily modulation is induced here too.

Channeling

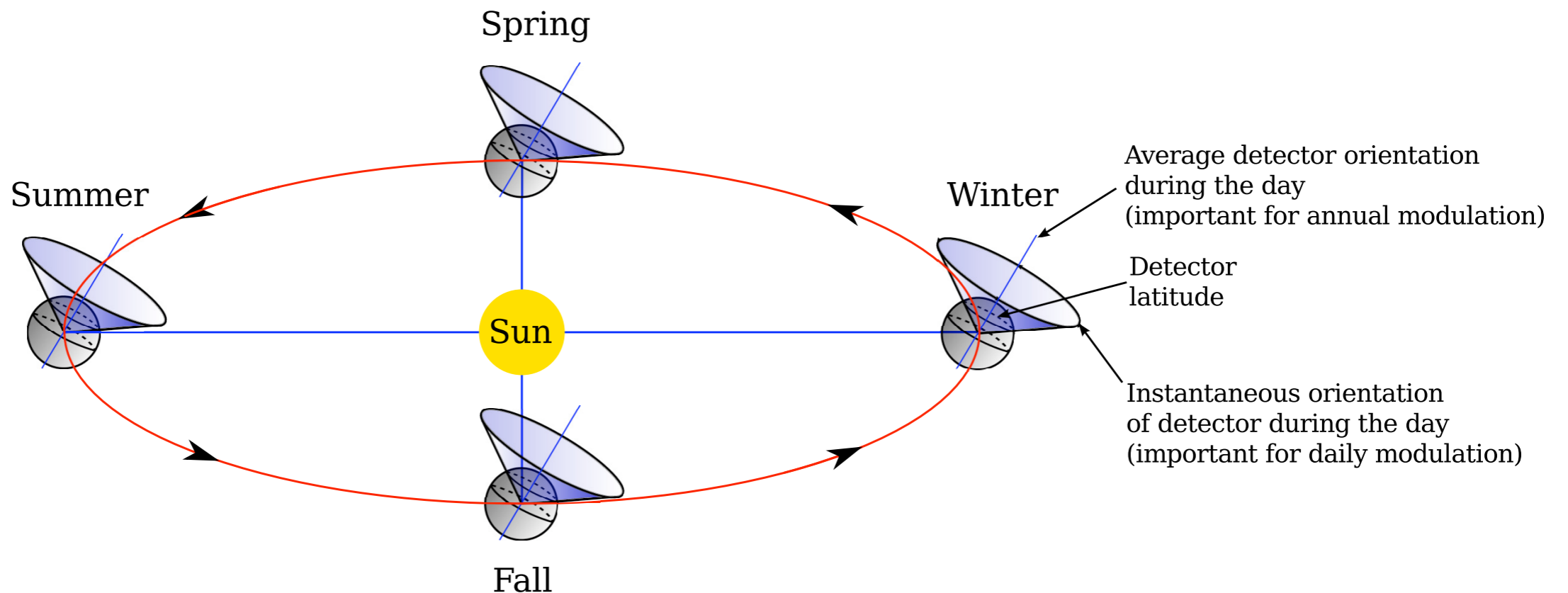
- * The dark matter signal is isotropic to zeroth order.
- * A signal coming from the sun is maximally anisotropic.
- * Imagine channeling occurs in some target crystals:



- * A highly angle dependent effect \rightarrow modulation!

Channeling

- * This can lead to a **daily** modulation, as well as a **annual** or **semi-annual** modulation.



Concluding

- * **Dark matter** and **neutrino** experiments share some features: low backgrounds, large exposure, low thresholds.
- * They can probe similar physics.
- * New physics connecting the SM to neutrinos can lead to **interesting direct detection signals**:
 - o New light gauge bosons.
 - o Neutrino dipole moments.
- * Many possibilities for rich **modulation** signals.

Deleted Scenes

Nuclear Recoil

Nuclear Recoil

- * The situation is very different w/ nuclear recoil.
- * SNO has measured Boron 8 neutrinos through deuterium dissociation.
- * SNO is probing a momentum transfer that is only a factor of a few higher than DAMA or CoGent.



- * A light mediator does not buy you much.

But...

Nuclear Recoil

- * Deuterium dissociation is an inelastic process.
- * The standard model rate at SNO is dominated by the axial-vector component of the Z interaction.
- * The vector component is suppressed by...

$$\frac{\sigma_{\nu_b-\text{Nucl}}(\text{elastic})}{\sigma_{\nu_b-\text{Nucl}}(\text{inelastic})} \sim \frac{A^2}{E_\nu^4 R_N^4} \sim 10^8$$

understanding this is in progress.

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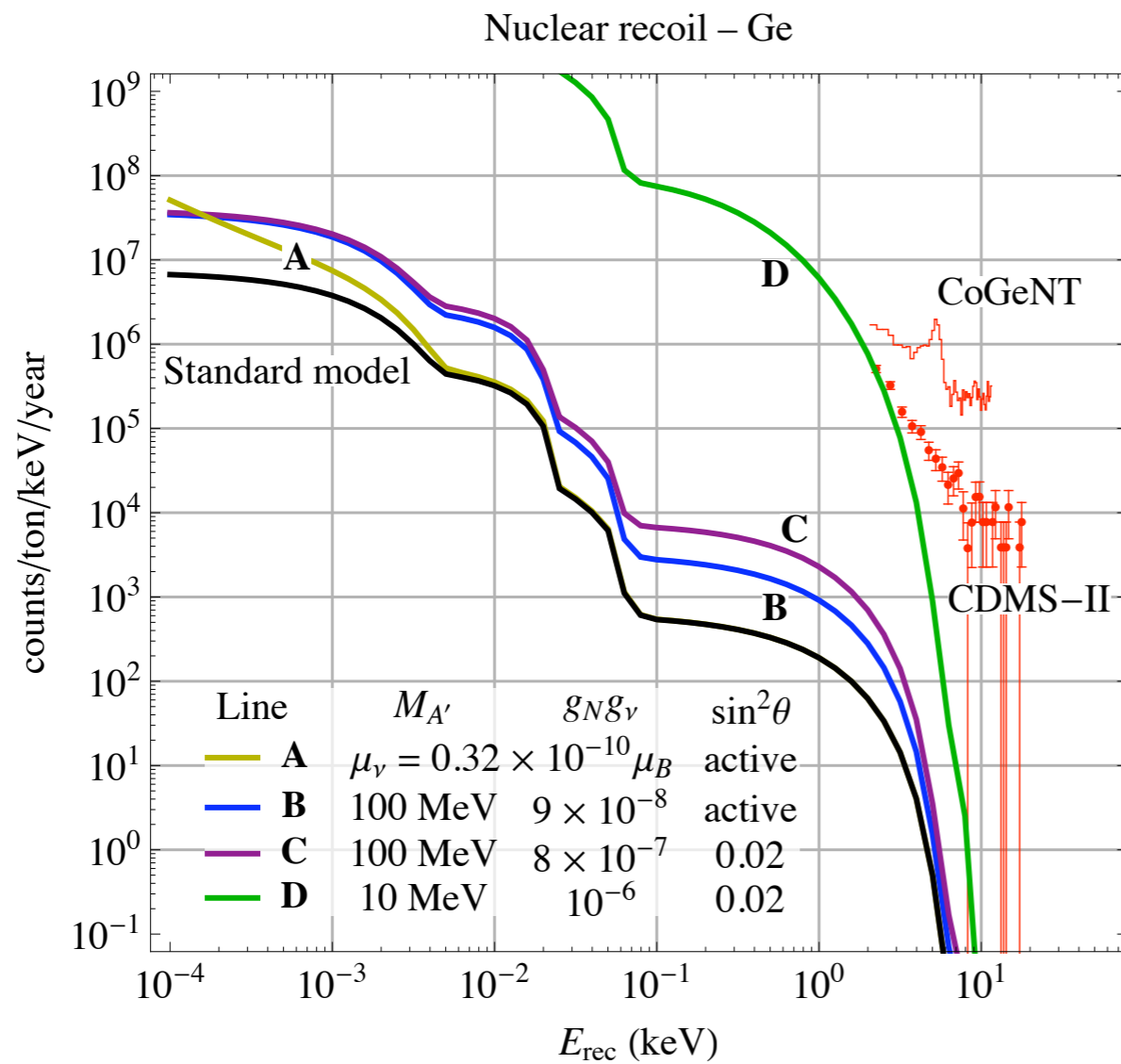
understanding this is in progress.

Pospelov:

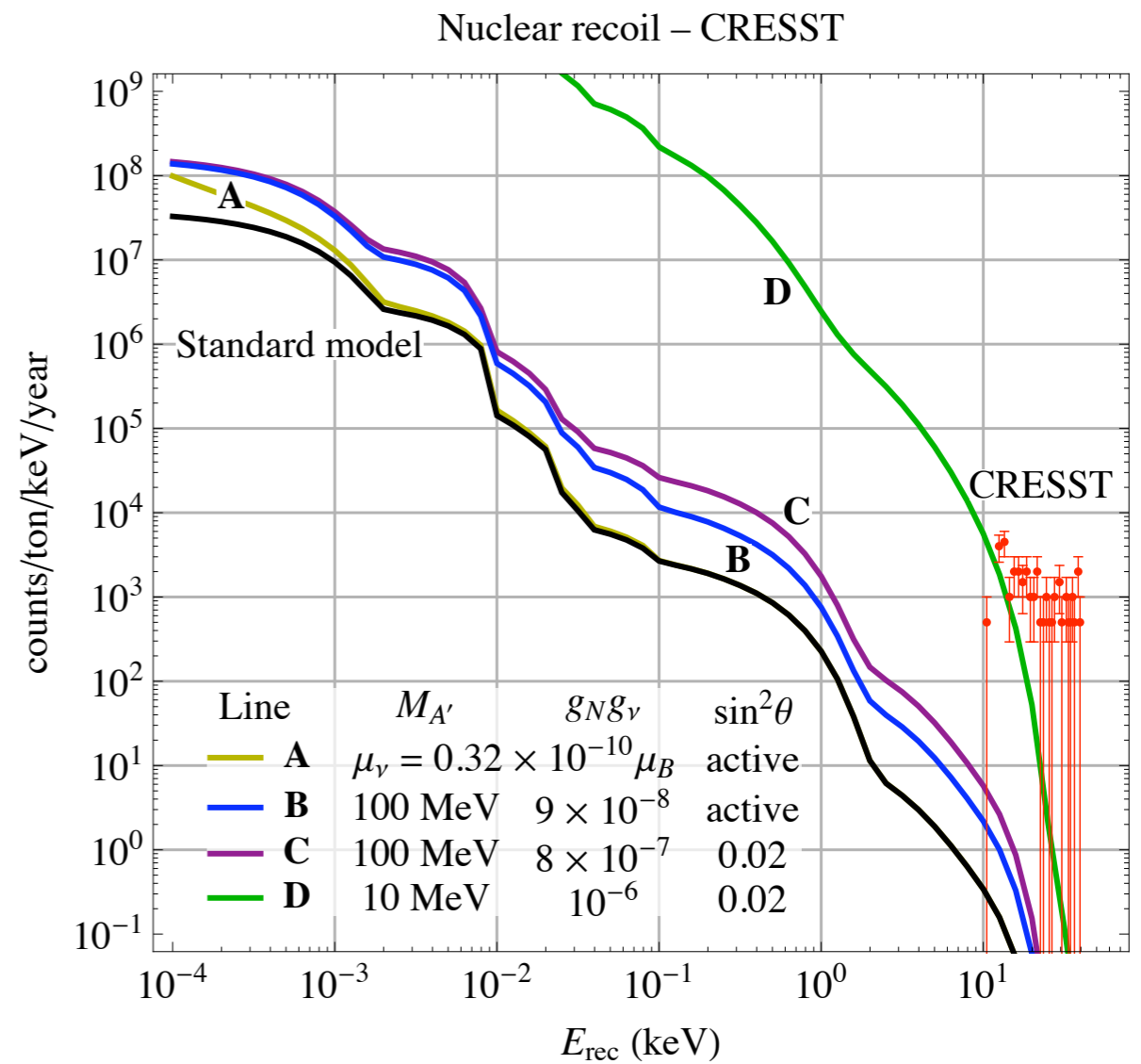
$$\begin{aligned} & \langle d | \exp(i\mathbf{q}\mathbf{r}^{(n)}) + \exp(i\mathbf{q}\mathbf{r}^{(p)}) | np \rangle \\ = & 2\langle d | np \rangle + i\mathbf{q} \cdot \langle d | \mathbf{r}^{(n)} + \mathbf{r}^{(p)} | np \rangle - \frac{q_k q_l}{2} \langle d | r_k^{(n)} r_l^{(n)} + r_k^{(p)} r_l^{(p)} | np \rangle = -\frac{q_k q_l}{4} \langle d | r_k r_l | np \rangle \end{aligned}$$

Nuclear recoil

* Interesting spectra are achievable:



(a)



(b)

SNO constraints may still be too much... (in progress)

Absorption

- * If the sterile scattering cross section is high, its m.f.p may be smaller than earth radius.
- * Neutrinos are captured during the night, but reach the detector during the day.
- * Steriles can still be produced via oscillation outside the sun.
- * The sterile flux may still be adjusted to fit the signal strength in direct detection.