# Chapter 2 Solution Manual

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#### 2.1:

For a wavelength of  $\lambda = 1 \mu m$ , an energy density of  $2 \times 10^{-9} \frac{J}{cm^3}$  corresponds to a photon density:

$$N_p = \frac{2 \times 10^{-9}}{\left(\frac{hc}{1\mu m}\right)} = 1.01 \times 10^{10} \frac{\text{photons}}{cm^3} \tag{1}$$

The rate of absorption events per atom with its lower state occupied  $(f_1 = 1, f_2 = 0)$  for a rate constant  $C = 1.5 \times 10^{-7} \frac{cm^3}{s}$  is then:

$$\frac{1}{N}\frac{dN_p}{dt} = -CN_p$$

$$= 1508.296\frac{\text{transitions}}{s}$$
(2)

which corresponds to a time of 0.663ms between transitions.

#### 2.2:

We first note that two levels corresponding to light emitted at  $\lambda = 880nm$  corresponds to an energy difference of:

$$\Delta E = \frac{hc}{\lambda} = 2.26 \times 10^{-19} J \tag{3}$$

If the Fermi level is midway between the levels, it is separated from each level by  $1.13 \times 10^{-19} J$ . At a temperature of T = 300K, the occupation probability for each level is given by:

$$f(E_{upper}) = \frac{1}{1 + exp\left(\frac{1.13 \times 10^{-19}}{k_B(300)}\right)} = 1.4 \times 10^{-12}$$
(4)

$$f(E_{lower}) = \frac{1}{1 + exp\left(\frac{-1.13 \times 10^{-19}}{k_B(300)}\right)} = 1 - 1.4 \times 10^{-12}$$
(5)

For a atomic density of  $N_a = 10^{18} cm^{-3}$ , the electron densities in each level are given by:

$$N_{upper} = 10^{18} (1.4 \times 10^{-12}) = 1.4 \times 10^6 cm^{-3}$$
(6)

$$N_{lower} = 10^{18} - N_{upper} \tag{7}$$

For a photon density of  $N_{ph} = 10^{10} cm^{-3}$ , the net transition rate in a volume V is given by:

$$\frac{dN_{ph}}{dt} = C \left( N_a \right) \left( f_2 - f_1 \right) N_{ph}$$
$$\approx -CN_a N_{ph}$$
$$= -1.5 \times 10^{21}$$
(8)

where we ignore the downward transitions rate since it is 12 orders of magnitude smaller than the upward transition rate. Decreasing photon population corresponds to upward transitions.

#### 2.3:

We first note that at equilibrium (when the Fermi level is in the middle of the bandgap), the difference between the two levels and the Fermi energy is  $E_2 - E_F = E_F - E_1 = 0.7 eV$ . Out of equilibrium, the energy differences become:

$$E_2 - E_{Fc} = 0.7 - \frac{1}{2}(E_{Fc} - E_{Fv})$$
(9)

$$E_{Fv} - E_1 = 0.7 - \frac{1}{2}(E_{Fc} - E_{Fv})$$
(10)

The occupation probabilities are then:

$$f(E_2) = \frac{1}{1 + exp\left(\frac{0.7 - \frac{1}{2}(E_{Fc} - E_{Fv})}{k_B T}\right)}$$
(11)

$$f(E_1) = \frac{1}{1 + exp\left(\frac{\frac{1}{2}(E_{Fc} - E_{Fv}) - 0.7}{k_B T}\right)}$$
(12)

Plotting the difference:



The point at which the curve crosses  $f_2 - f_1$  represents the onset of population inversion, the critical condition necessary for optical gain.

## 2.4:

The condition for gain is for the occupation probability of the conduction band to be greater than that of the valence band:

$$f(E_{cb}) > f(E_{vb}) \quad \leftrightarrow \quad \frac{1}{1 + exp\left(\frac{E_{cb} - E_{Fc}}{k_B T}\right)} > \frac{1}{1 + exp\left(\frac{E_{vb} - E_{Fv}}{k_B T}\right)} \tag{13}$$

This condition can be rewritten as follows:

$$\frac{1 + exp\left(\frac{E_{vb} - E_{Fv}}{k_B T}\right)}{\left[1 + exp\left(\frac{E_{vb} - E_{Fv}}{k_B T}\right)\right] \left[1 + exp\left(\frac{E_{cb} - E_{Fc}}{k_B T}\right)\right]} > \frac{1 + exp\left(\frac{E_{vb} - E_{Fv}}{k_B T}\right)}{\left[1 + exp\left(\frac{E_{vb} - E_{Fv}}{k_B T}\right)\right] \left[1 + exp\left(\frac{E_{cb} - E_{Fc}}{k_B T}\right)\right]}$$

$$1 + exp\left(\frac{E_{vb} - E_{Fv}}{k_B T}\right) > 1 + exp\left(\frac{E_{cb} - E_{Fc}}{k_B T}\right)$$

$$E_{vb} - E_{Fv} > E_{cb} - E_{Fc}$$

$$E_{Fc} - E_{Fv} > E_{cb} - E_{vb} \qquad (14)$$

We have thus derived the Bernard-Duraffourg condition.

2.5:

We first note that from the curve in problem (2.3), the occupation probability difference at  $\Delta E_F = 1.45 eV$  is found to be:

$$f_2 - f_1 = 0.4487 \quad (\text{at } \Delta E_F = 1.45eV)$$
 (15)

For the parameters given, the gain is thus:

$$G = \frac{CNn}{c}(f_2 - f_1) = 897.4\tag{16}$$

which is also the maximum gain since only one photon energy can induce transitions between the two-level system.

### 2.6:

The intensity of the light as a function of distance traveled L is given by:

$$I(L) = I_0 e^{GL} \tag{17}$$

For the intensity to be tripled, we have the condition:

$$3I_0 = I_0 e^{GL}$$
  
 $\to L = \frac{\ln(3)}{G} = 0.055 cm$  (18)