

Higgs@125 GeV: Implications for the MSSM and GMSB

David Shih

Rutgers University

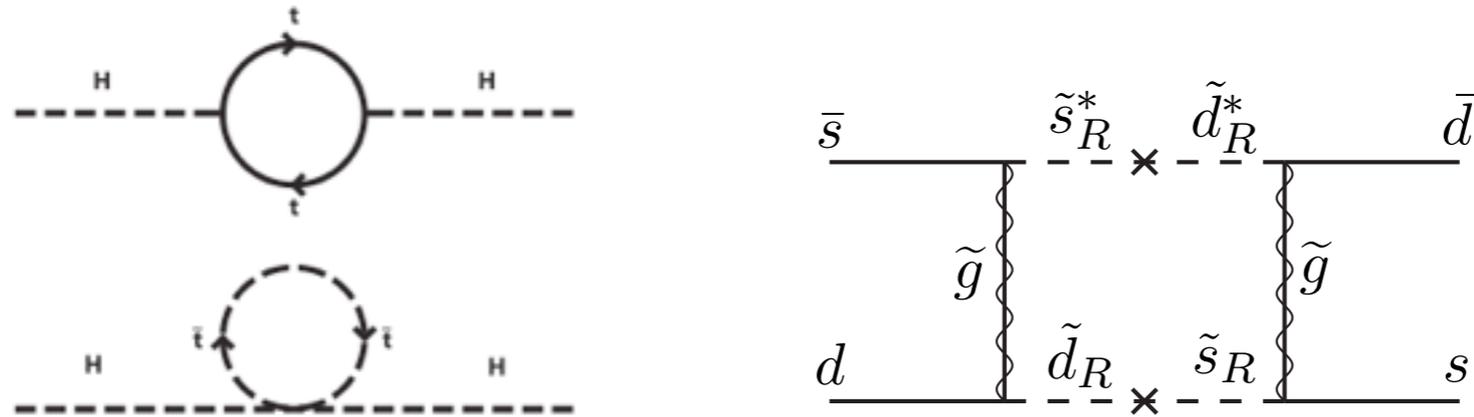
MCTP Symposium on Higgs Physics

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Based on:

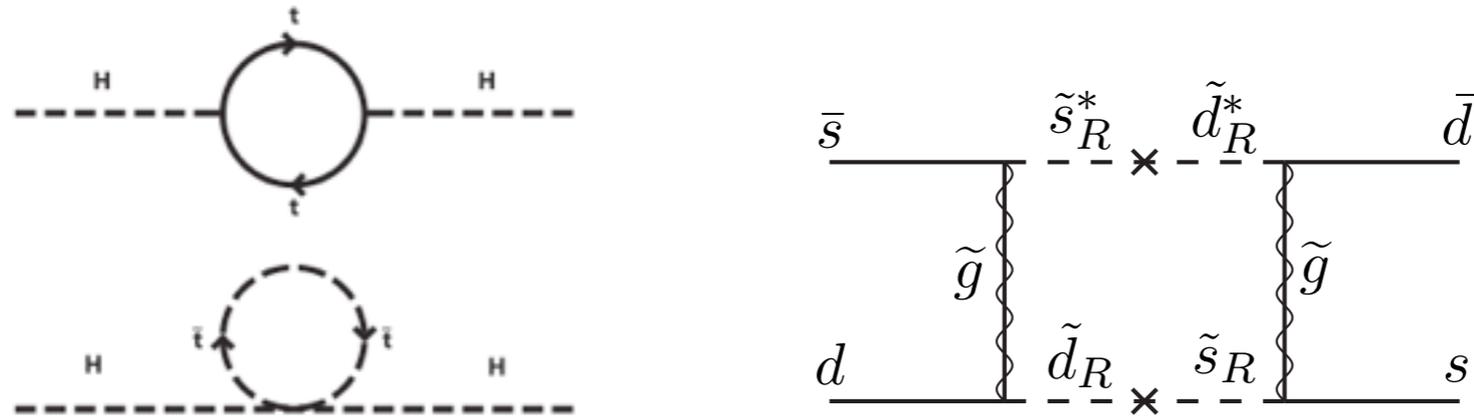
[Draper, Meade, Reece & DS \(1112.3068\)](#)

Higgs@125 GeV



- A SM-like Higgs at 125 GeV would renew the urgency of the hierarchy problem. What new physics protects the Higgs mass??
- Supersymmetry is still an attractive idea for solving the hierarchy problem. Its minimal realization is the MSSM.
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Question for this talk:

MSSM + GMSB + Higgs@125 GeV \Rightarrow ???

Implications for the MSSM

Higgs mass in the MSSM

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We will set $m_A = 1$ TeV in today's talk.

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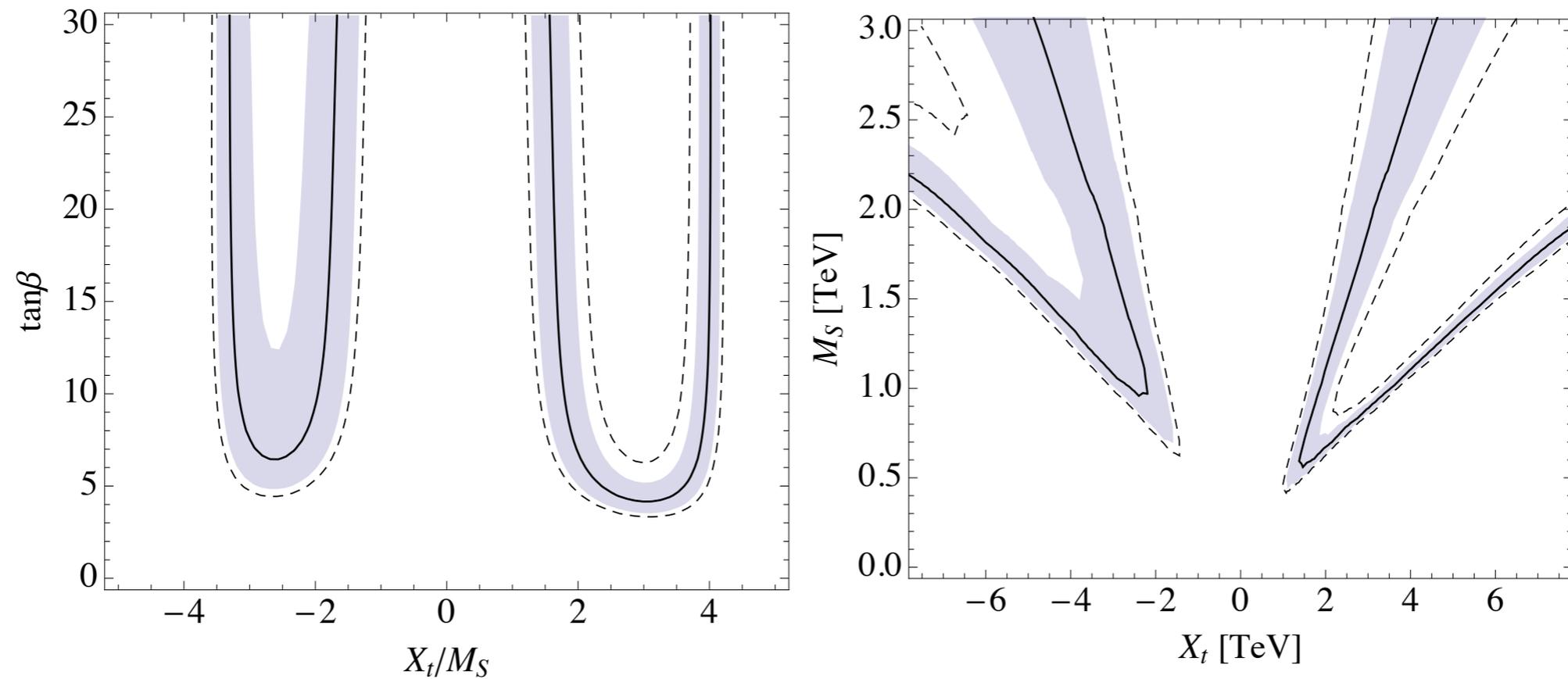
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- Higgs mass decreases with $\tan \beta$. Expect lower bound on $\tan \beta$.
- Quartic polynomial in X_t/M_S . Expect four preferred values for X_t/M_S .
- Log dependence on M_S . Expect lower bound on M_S .

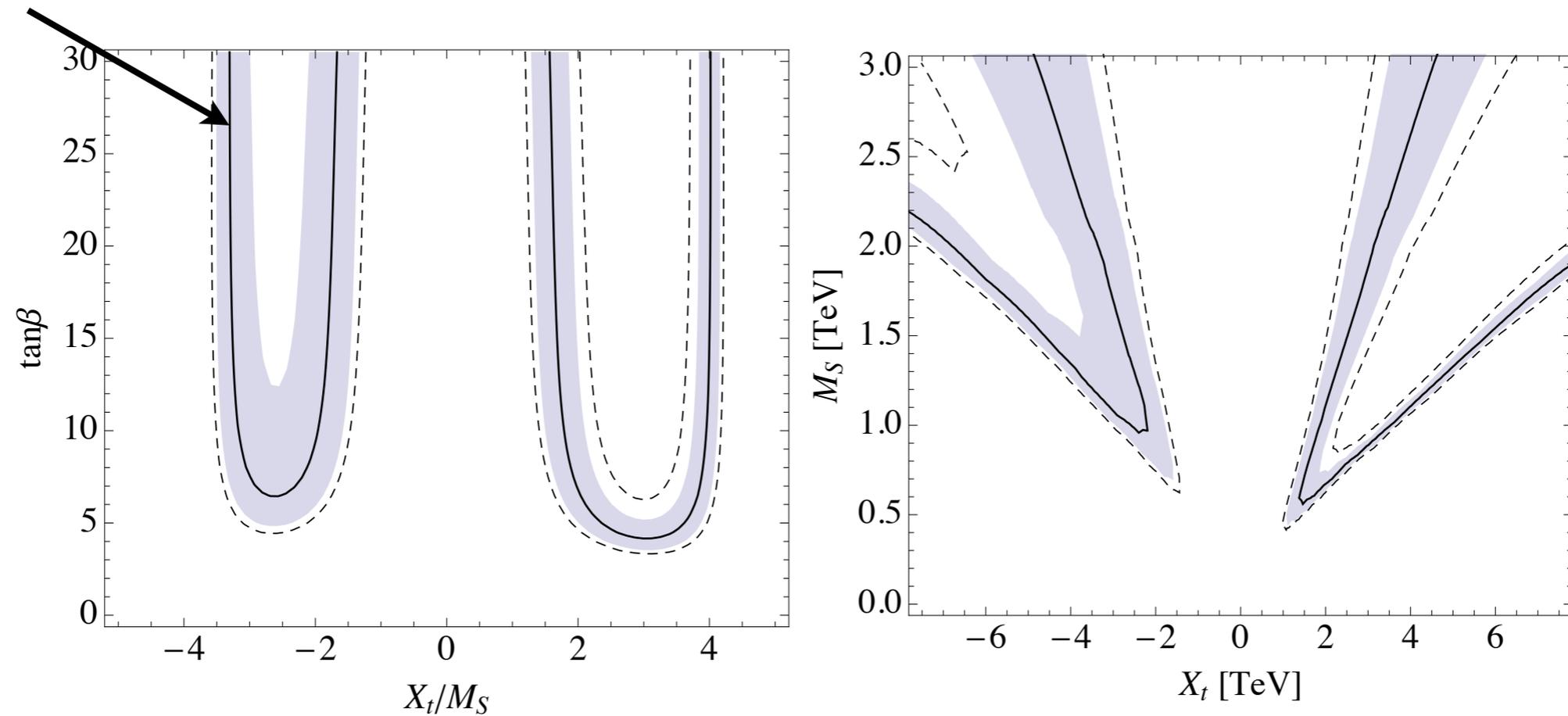
Higgs@125 GeV in the MSSM



$$\tan\beta \gtrsim 3.5 \quad \frac{X_t}{M_S} \approx -3, -1.7, 1.5, \text{ or } 3.5 \quad M_S \gtrsim 500 \text{ GeV}$$

Higgs@125 GeV in the MSSM

mh=125 GeV
m_{top}=173.2 GeV

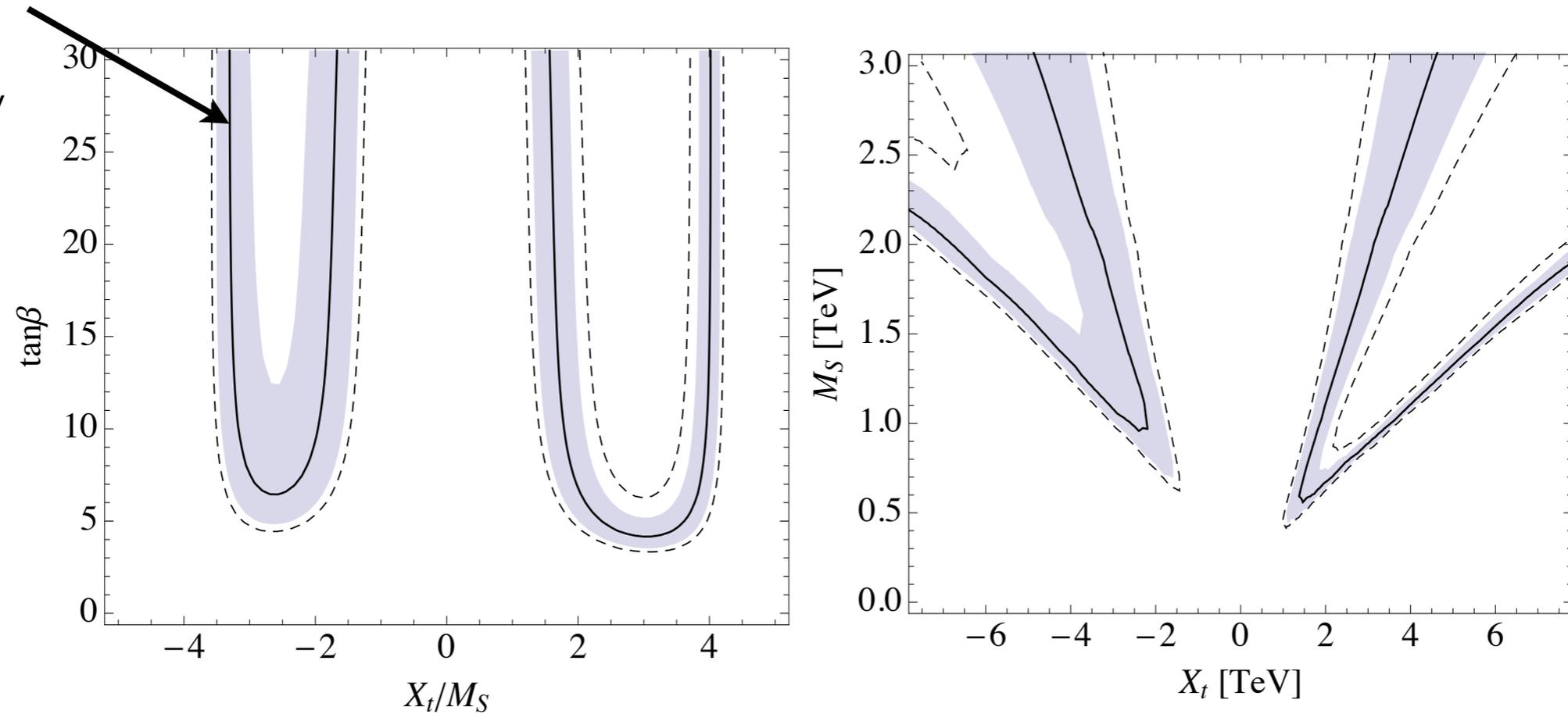


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 $m_{\text{top}}=173.2$ GeV

Grey band:
 $m_h=123-127$ GeV

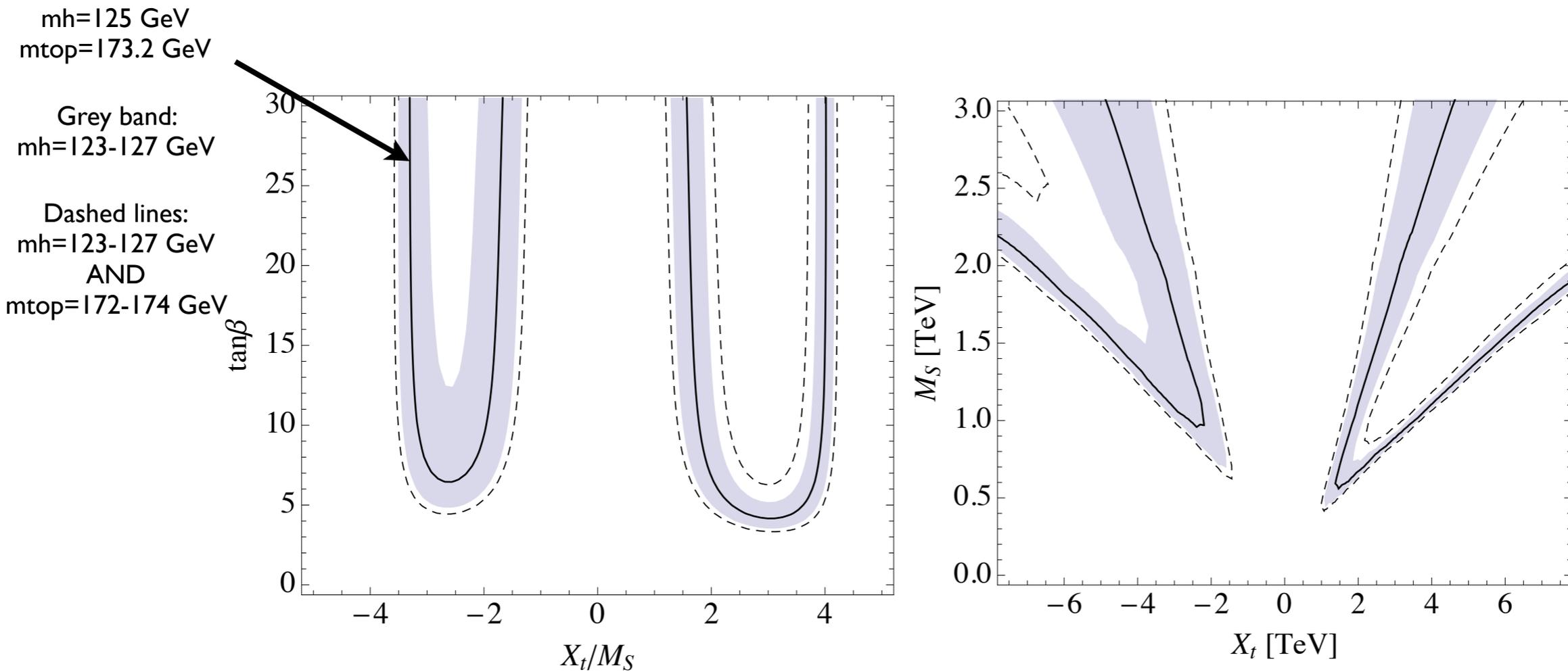


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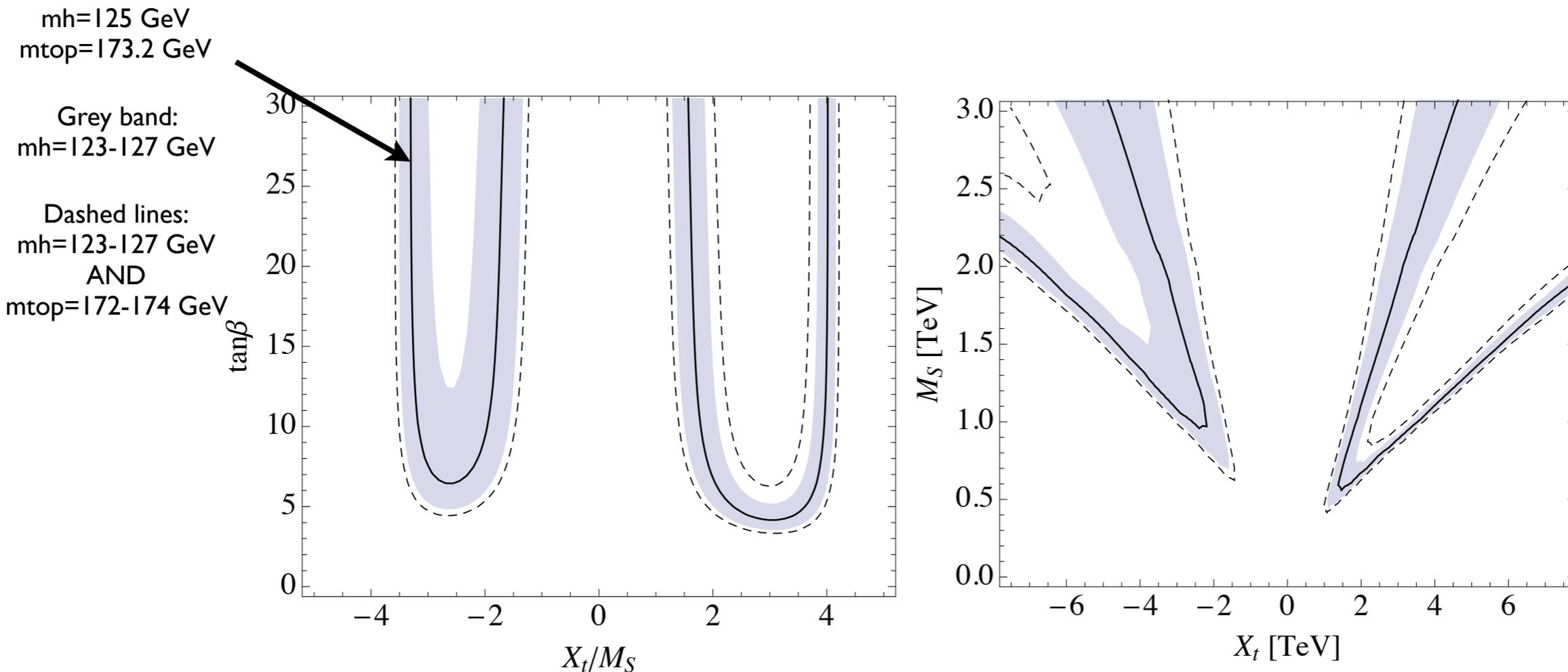
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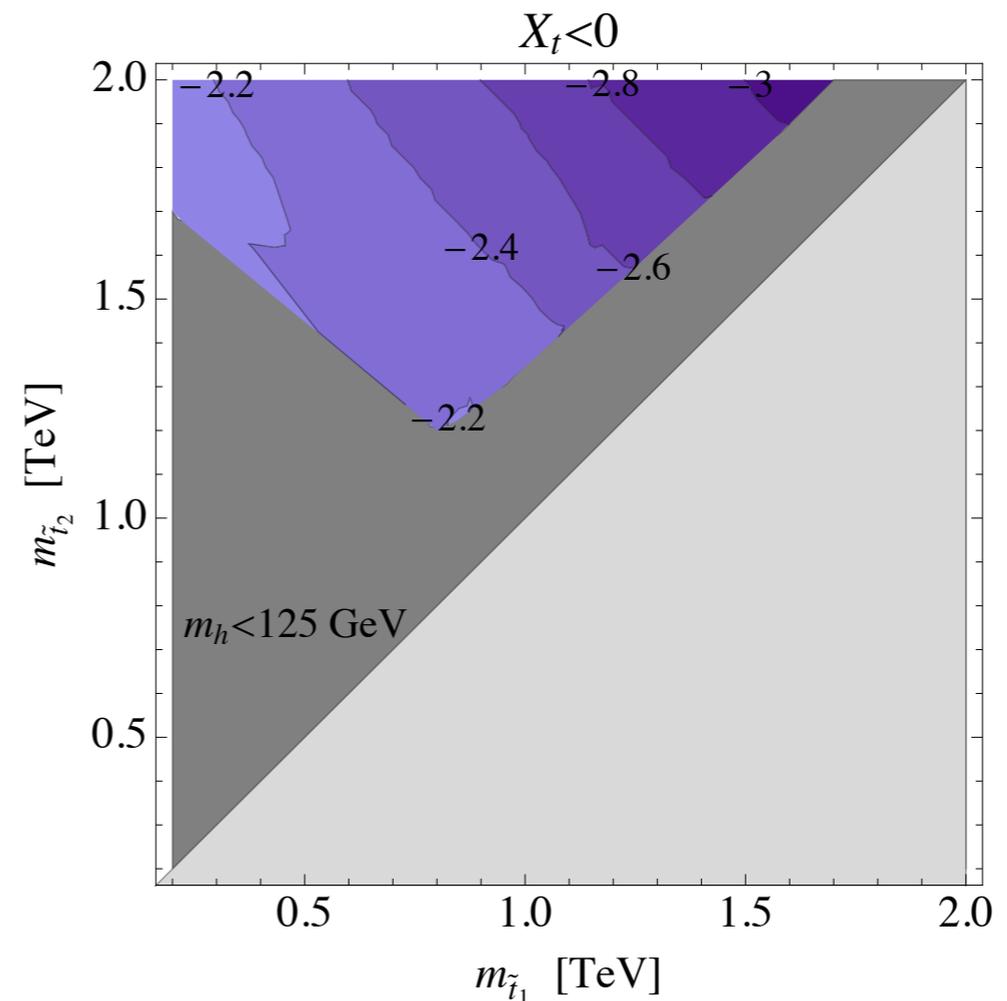
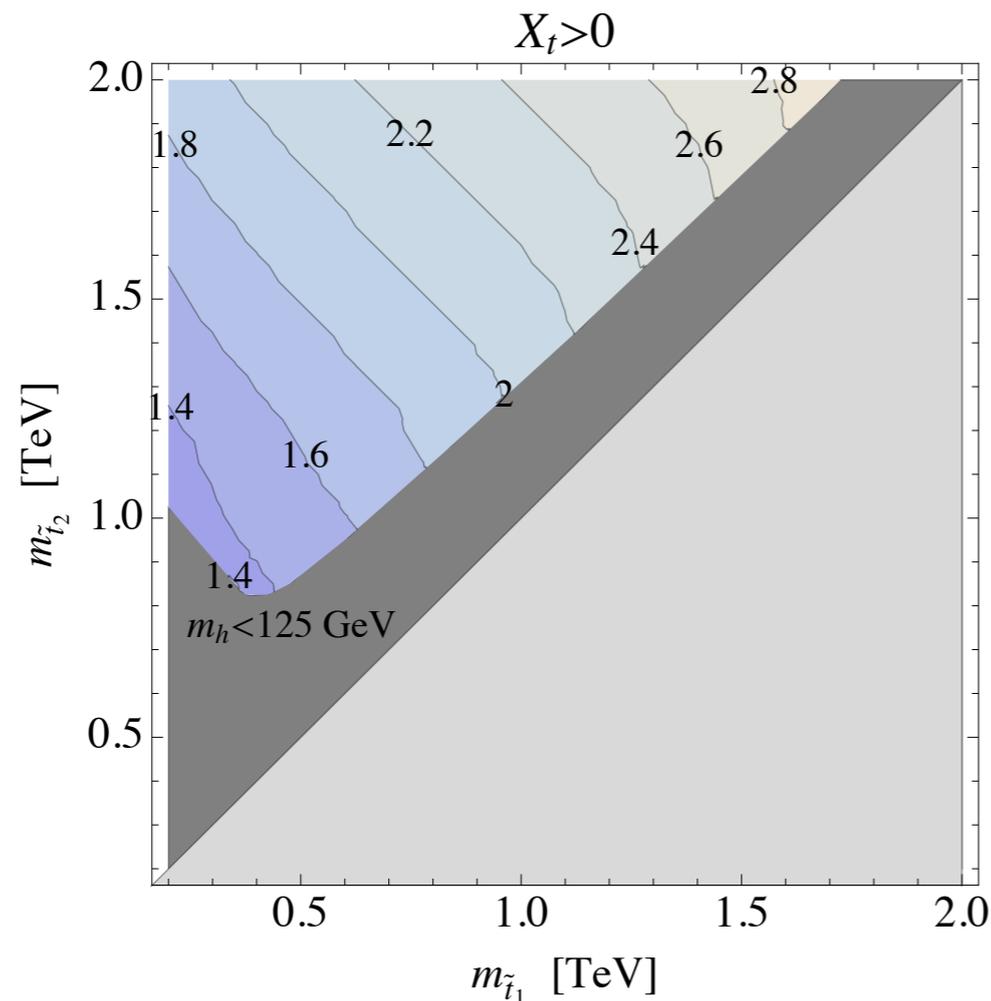
$$\Rightarrow |X_t| \gtrsim 1000 \text{ GeV}$$

Stop mixing (A-terms) must be large in the MSSM!

Higgs@125 GeV in the MSSM:

X_t vs (m_{t1}, m_{t2})

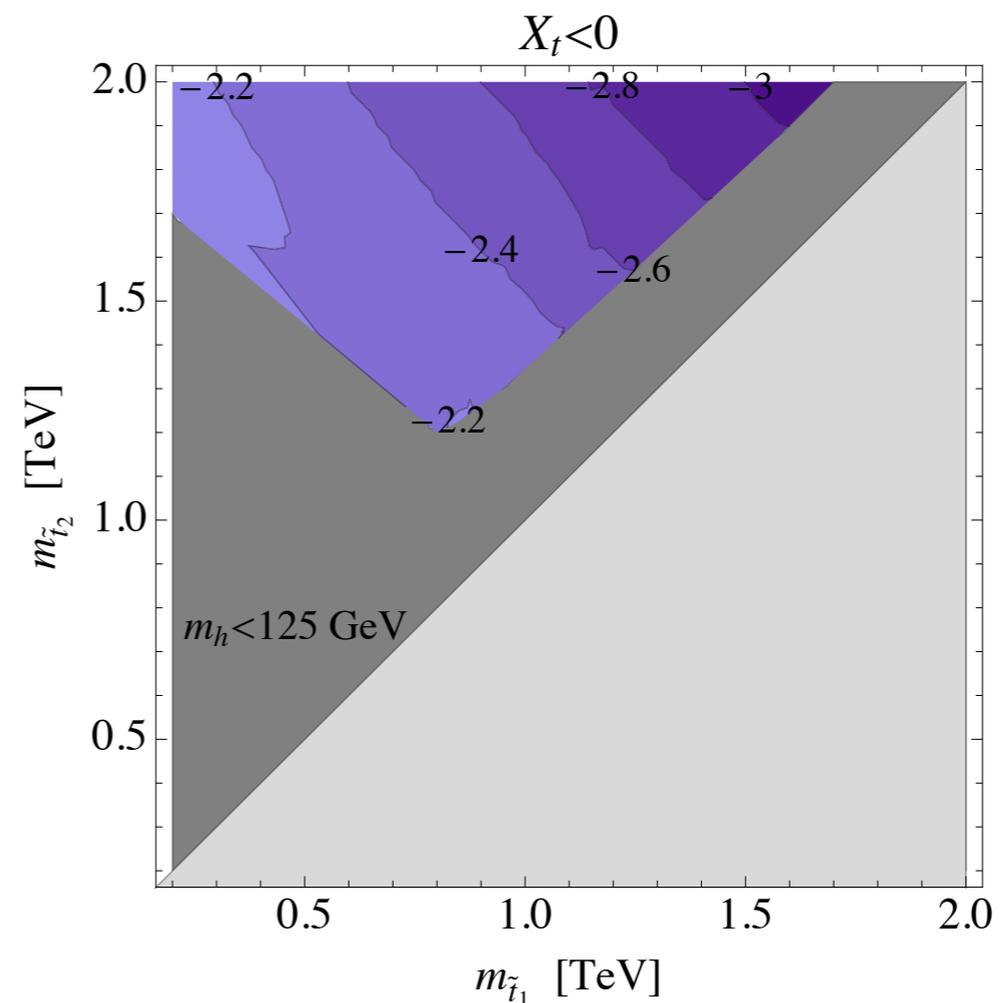
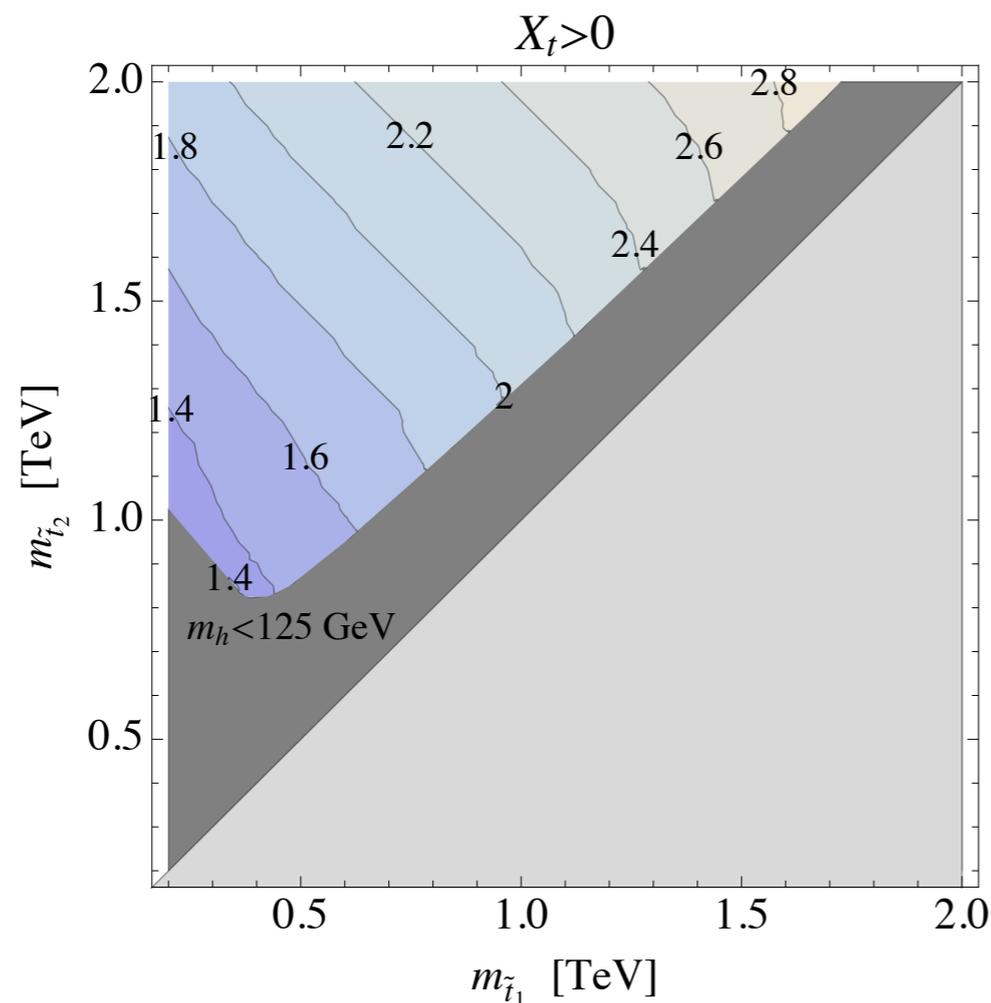
Previous plots assumed $m_Q = m_U \approx M_S$ for simplicity. Easy to generalize to split stop masses. For a given $(m_{\tilde{t}_1}, m_{\tilde{t}_2})$, solve $m_h(m_{\tilde{t}_1}, m_{\tilde{t}_2}, X_t) = 125$ GeV for X_t .



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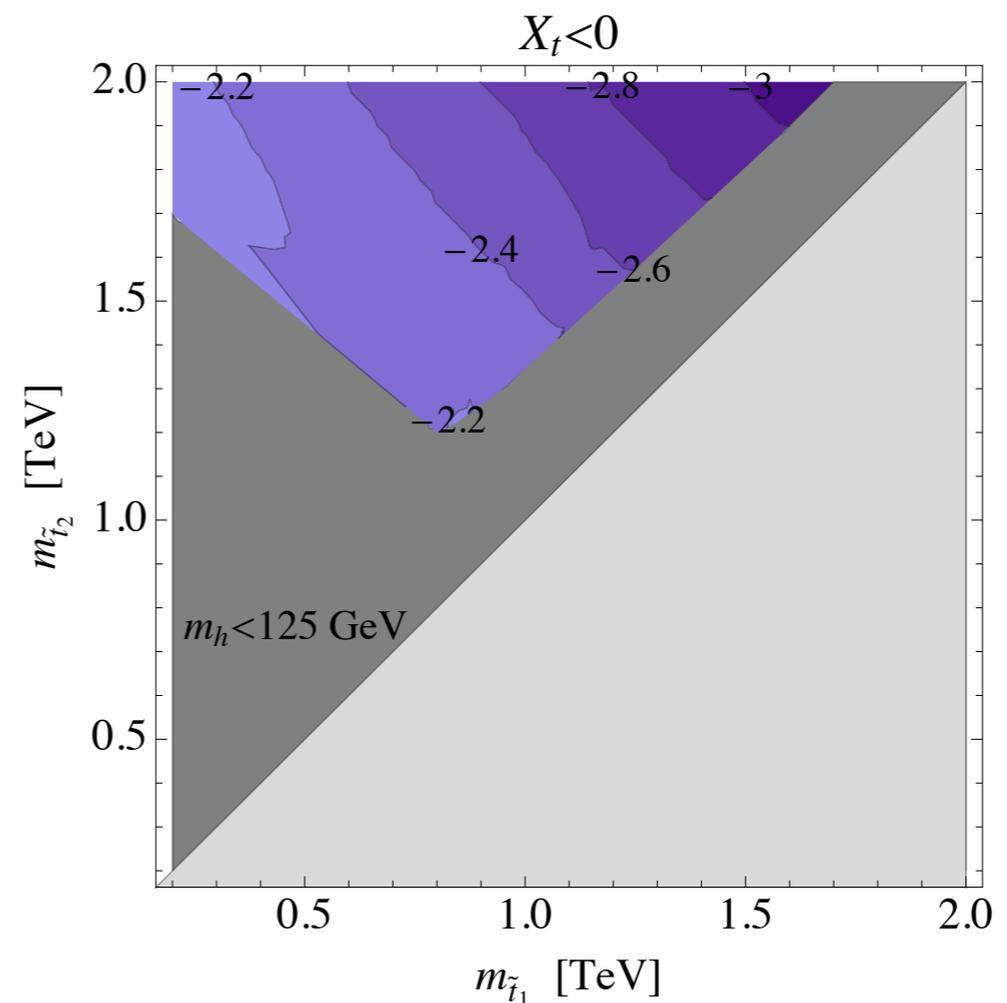
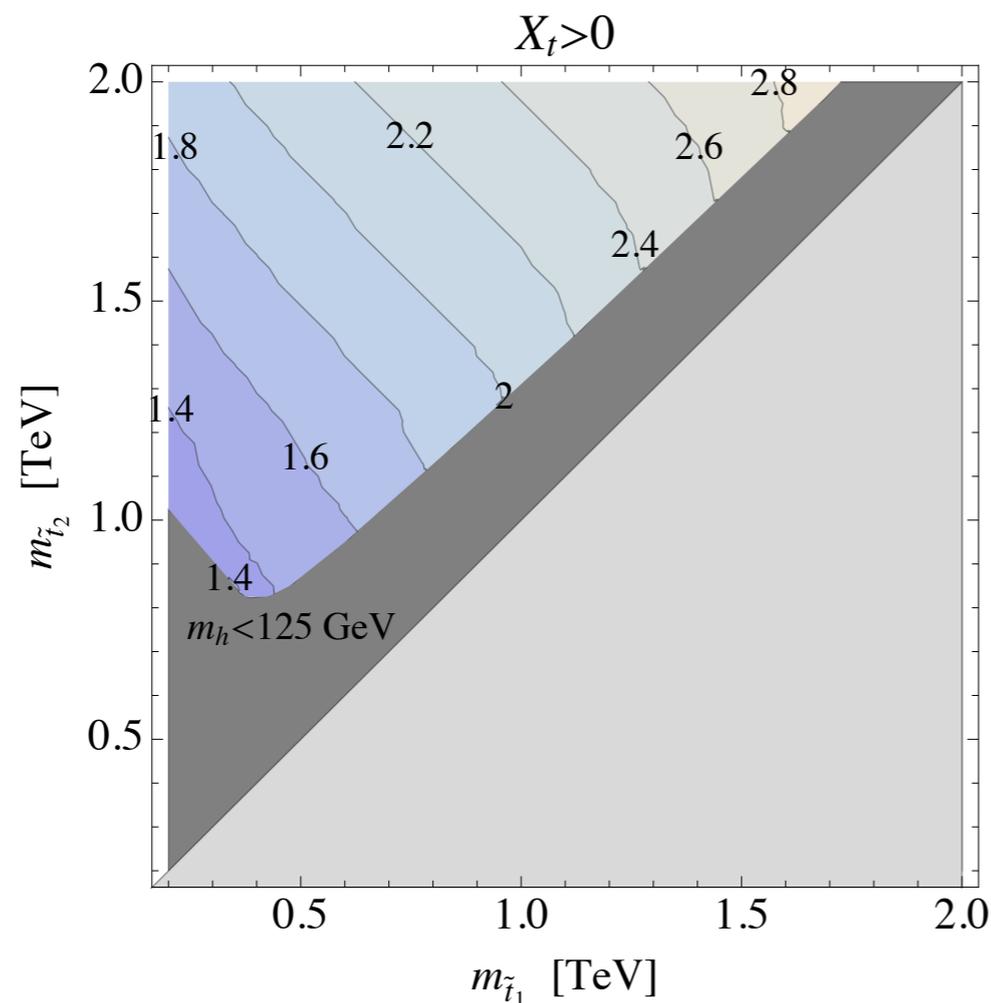


Conclusions about min X_t unchanged.

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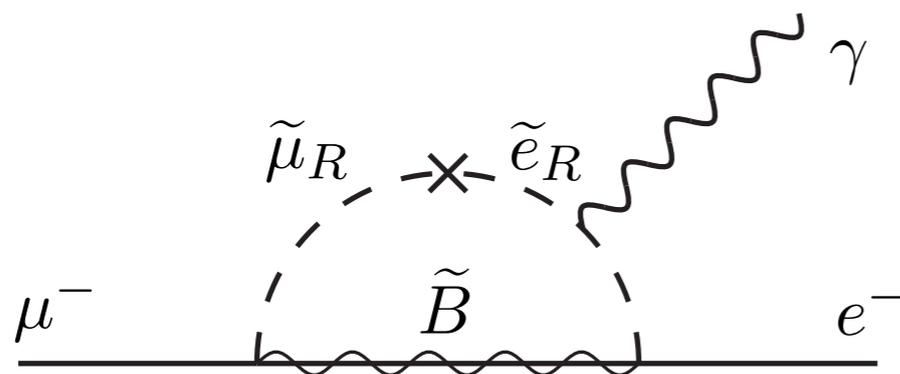
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Stop1 can be light, provided stop2 is sufficiently heavy.

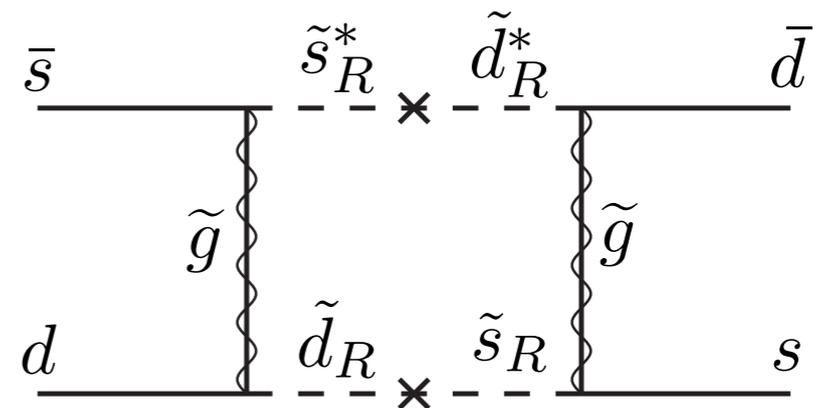
Implications for GMSB

MSSM Flavor Problem

- The ~ 120 soft SUSY-breaking parameters of the MSSM must arise from a simpler underlying source.
- A generic soft Lagrangian suffers from a severe **flavor problem**.
- The mediation of SUSY-breaking must be **flavor universal** to a very high degree of accuracy.

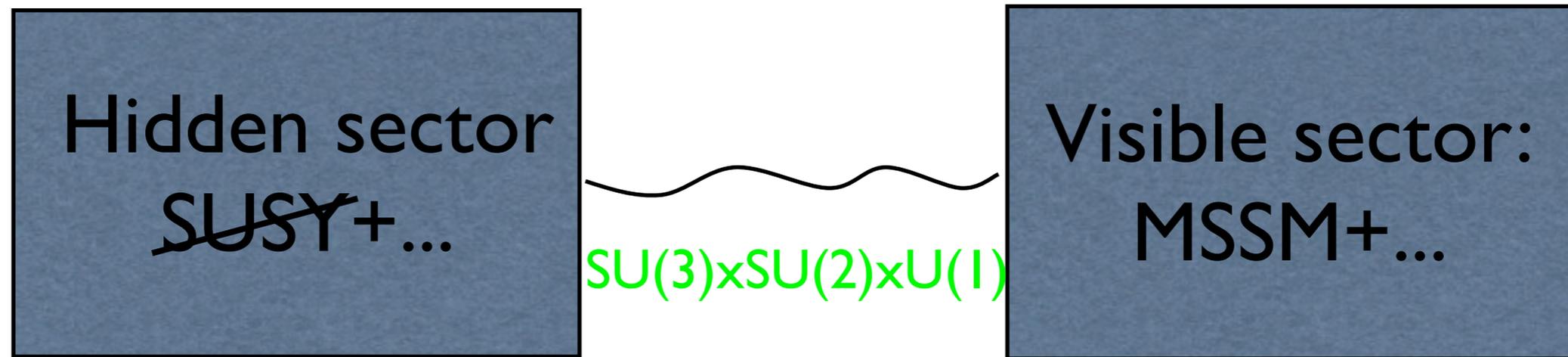


$$\text{Br}(\mu \rightarrow e\gamma)_{exp} < 1.2 \times 10^{-11}$$

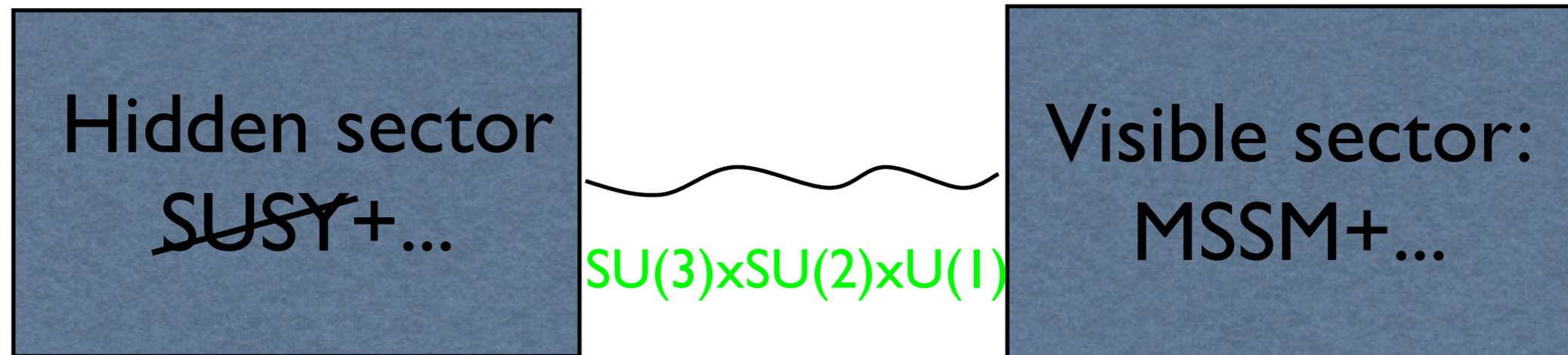


$$(\Delta m_K)_{exp} = (3.483 \pm 0.006) \times 10^{-12} \text{ MeV}$$

Gauge Mediation

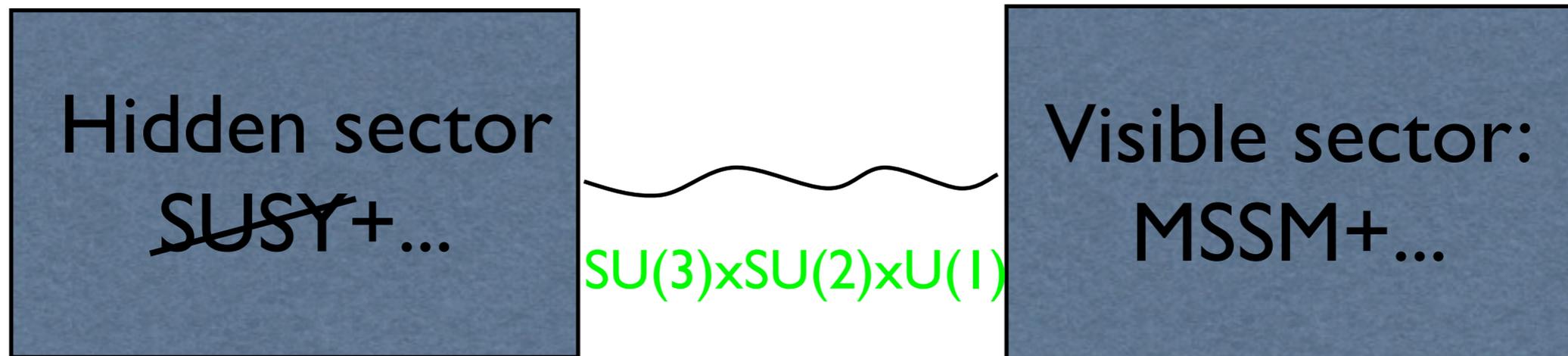


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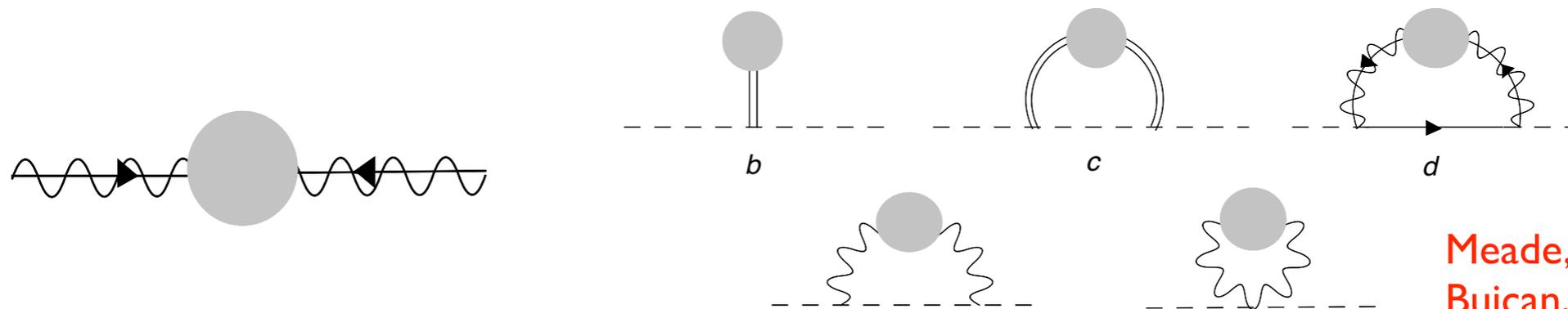


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Gauge Mediation



- Gauge mediation offers the best (only?) explanation for the required flavor universality of the soft Lagrangian.
- In 2008, my collaborators and I invented a model-independent framework for GMSB: **“General Gauge Mediation”**



Meade, Seiberg & DS;
Buican, Meade, Seiberg, DS

Idea: express soft masses in terms of currents and their correlation functions

GGM: Higgs Mass

- Using this, we derived the most general predictions of gauge mediation:
 - Parameter space: $(A_1, A_2, A_3) \rightarrow m_{Q,u,d,L,e}^2$
 $(B_1, B_2, B_3) \rightarrow M_{1,2,3}$
 - Sum rules: $\text{Tr} (B - L)m_{\tilde{f}}^2 = \text{Tr} Y m_{\tilde{f}}^2 = 0$
 - A-terms ≈ 0

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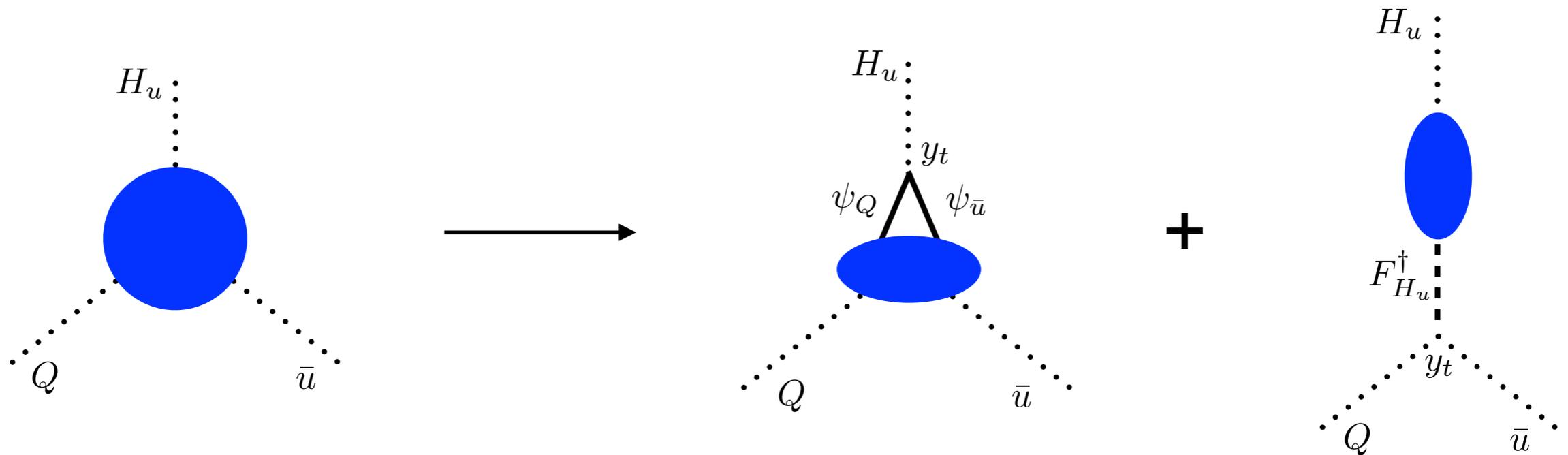
Higgs mass primarily relates to this



A-terms in GGM

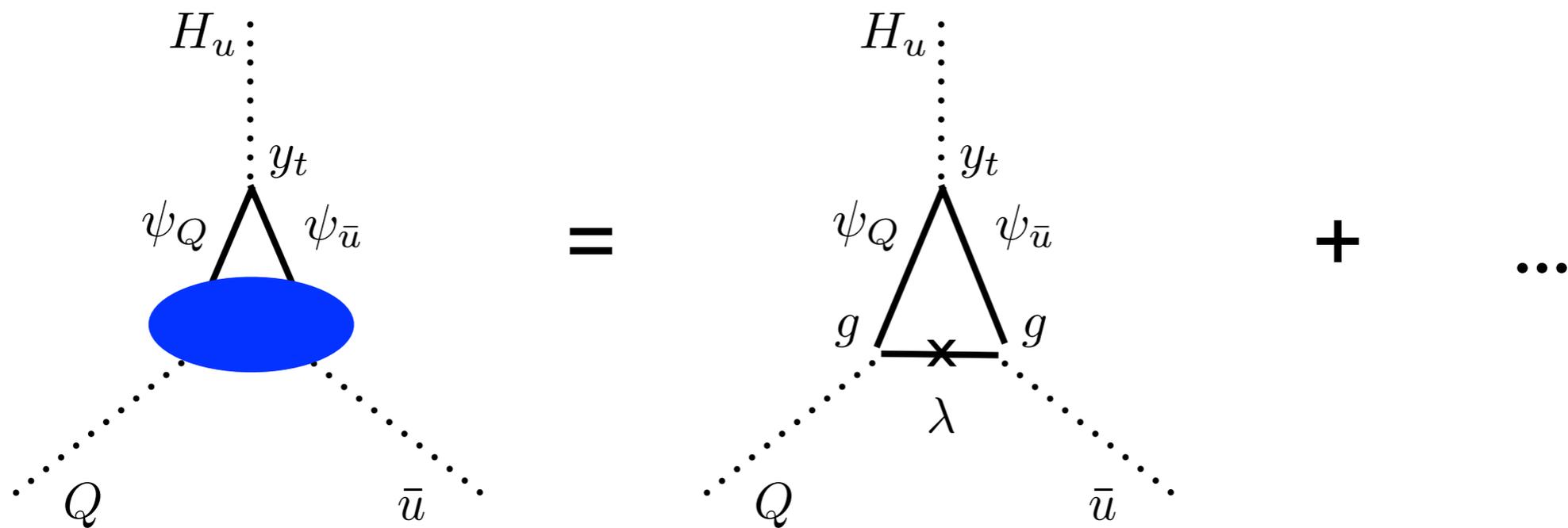
$$\mathcal{L} \supset A_t Q \bar{u} H_u + c.c.$$

- With gauge interactions alone, A terms are not generated at all -- they are protected by phase rotations of the squark fields.
- Need to involve Yukawa couplings.



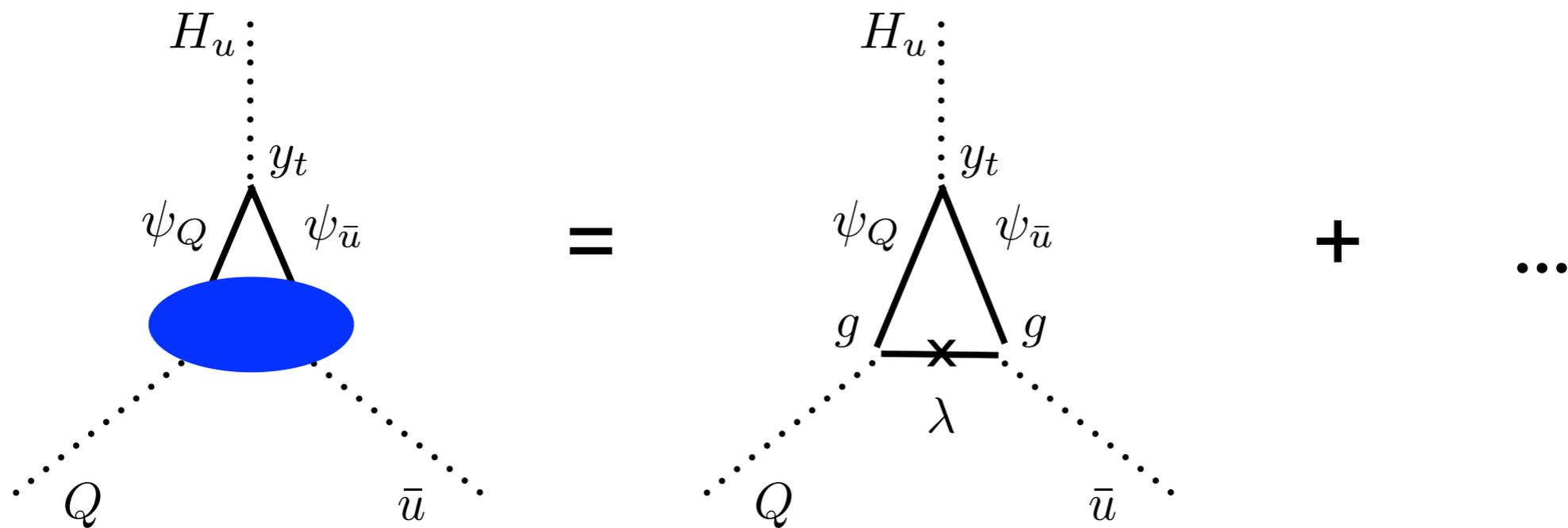
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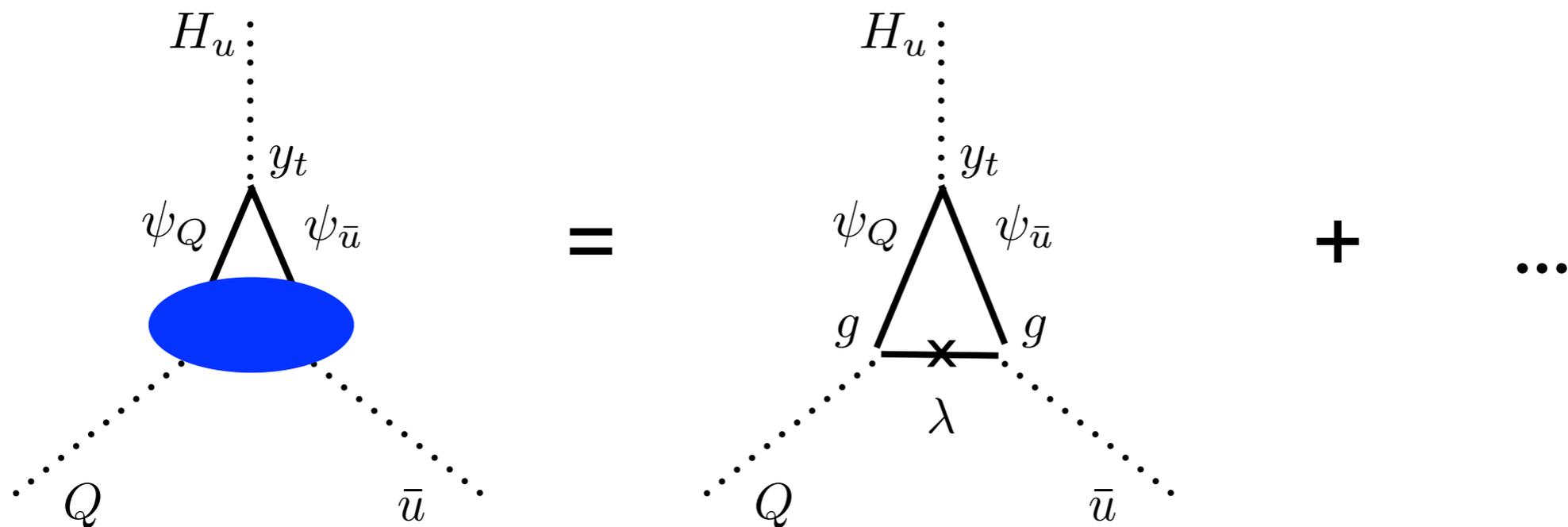
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This is the standard MSSM contribution to A terms!

$$\frac{dA_t}{dt} \sim y_t^2 A_t + g_3^2 M_3$$

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- Higgs@125 in the MSSM: A-terms must be large at weak scale.
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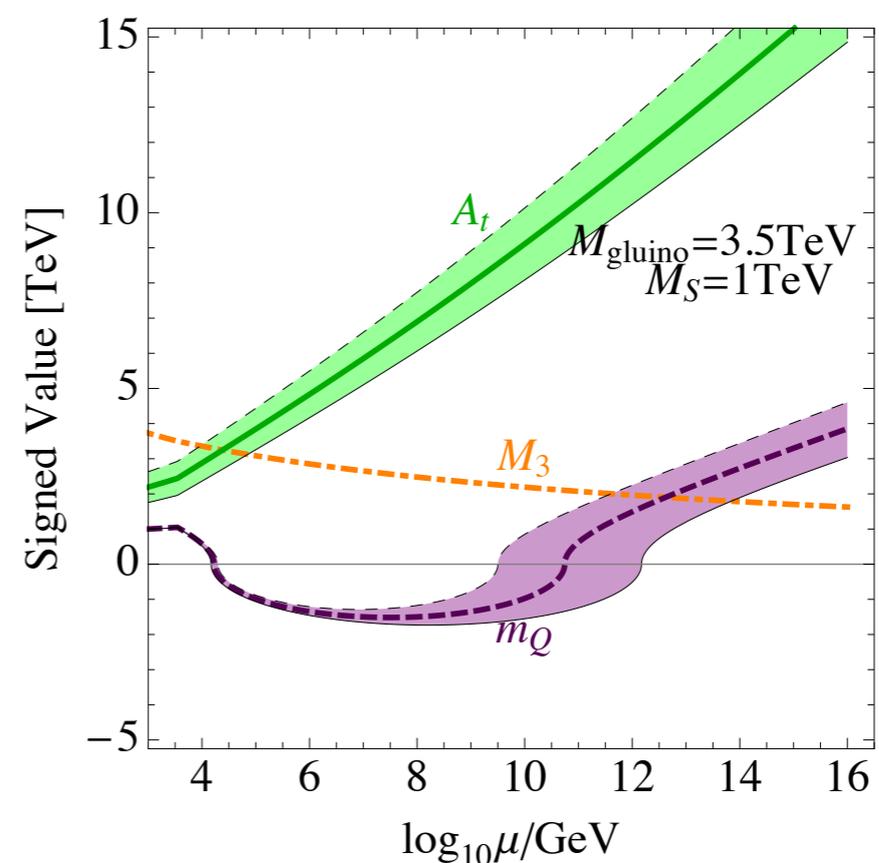
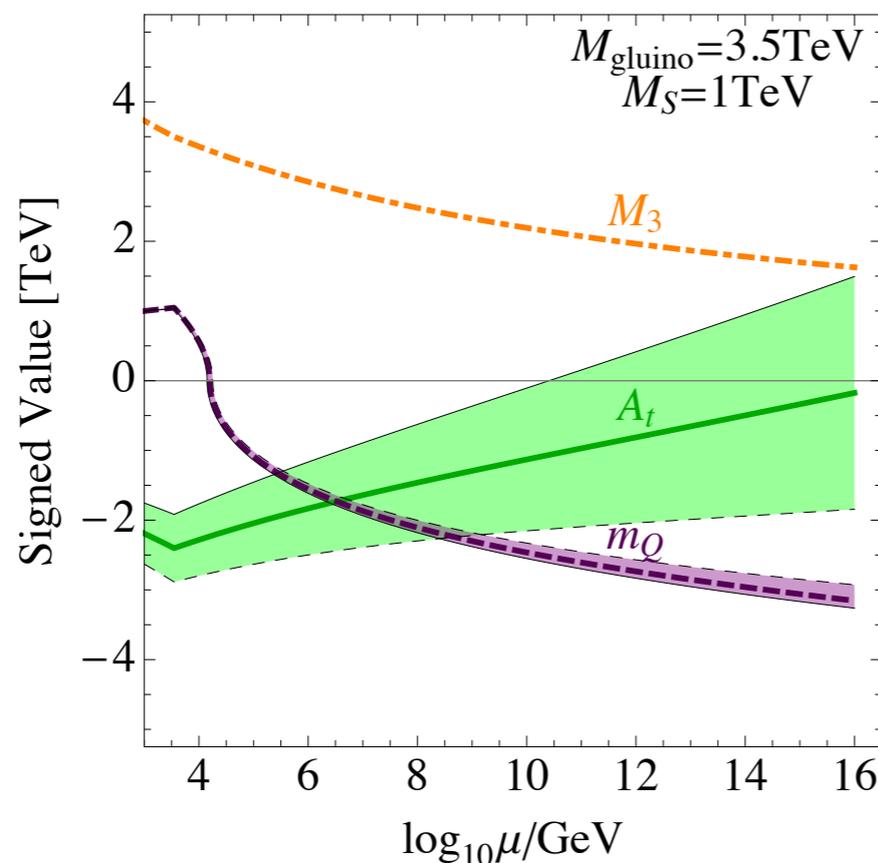
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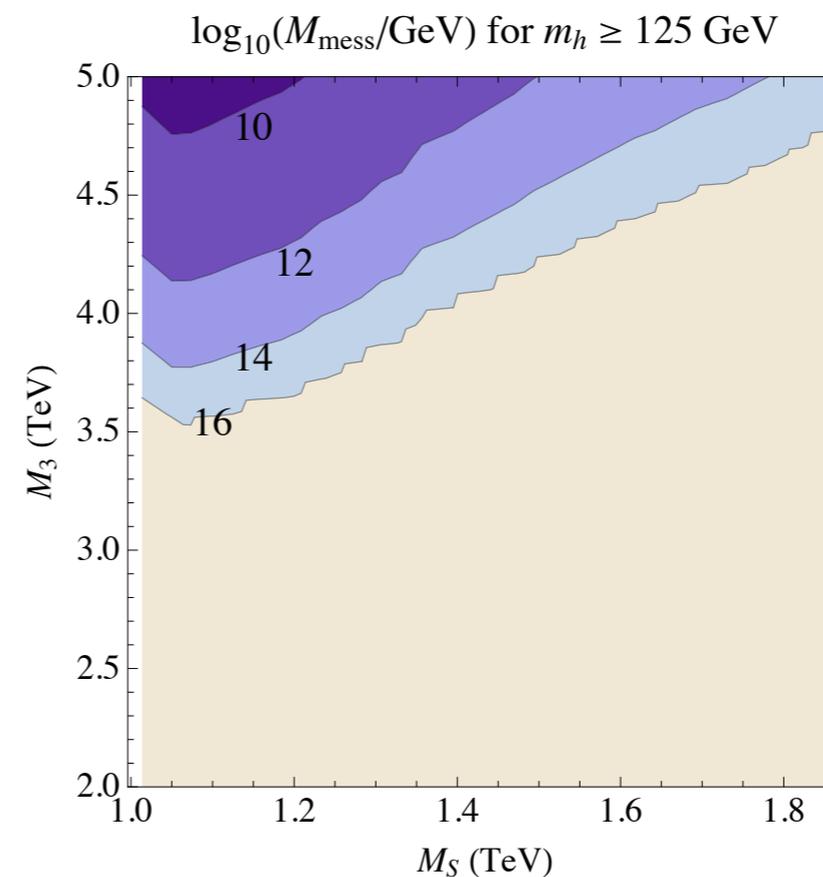
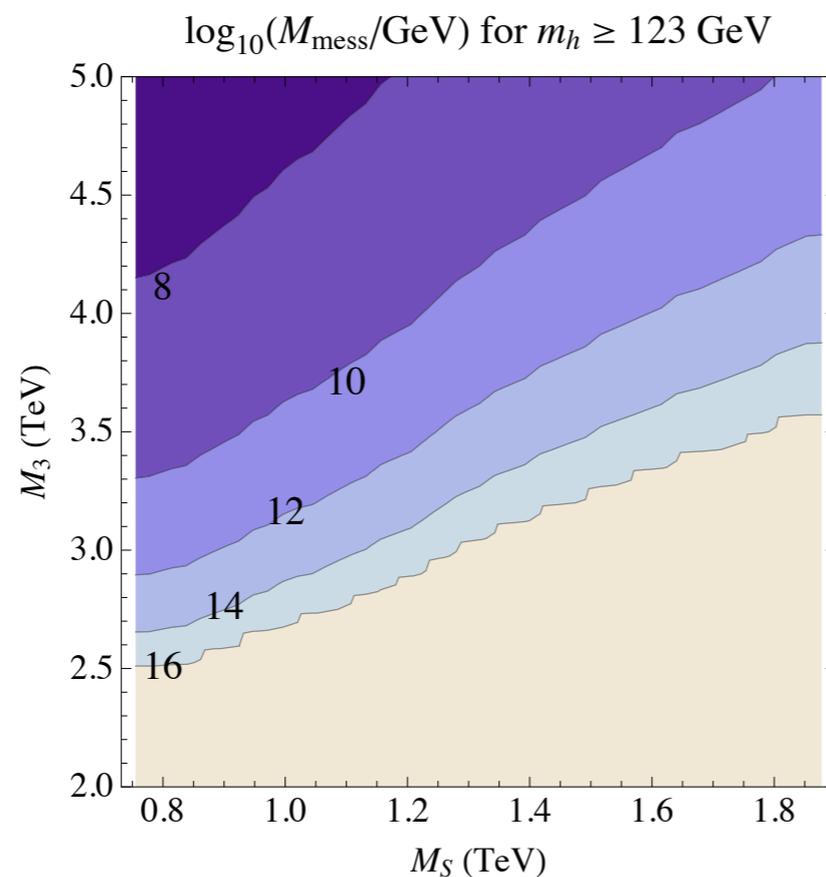
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Higgs@125 GeV in the MSSM +GMSB

- $A_t > 0$ not possible in vanilla GMSB.
- $A_t < 0$ possible, but requires large M_3 and M_{mess} .

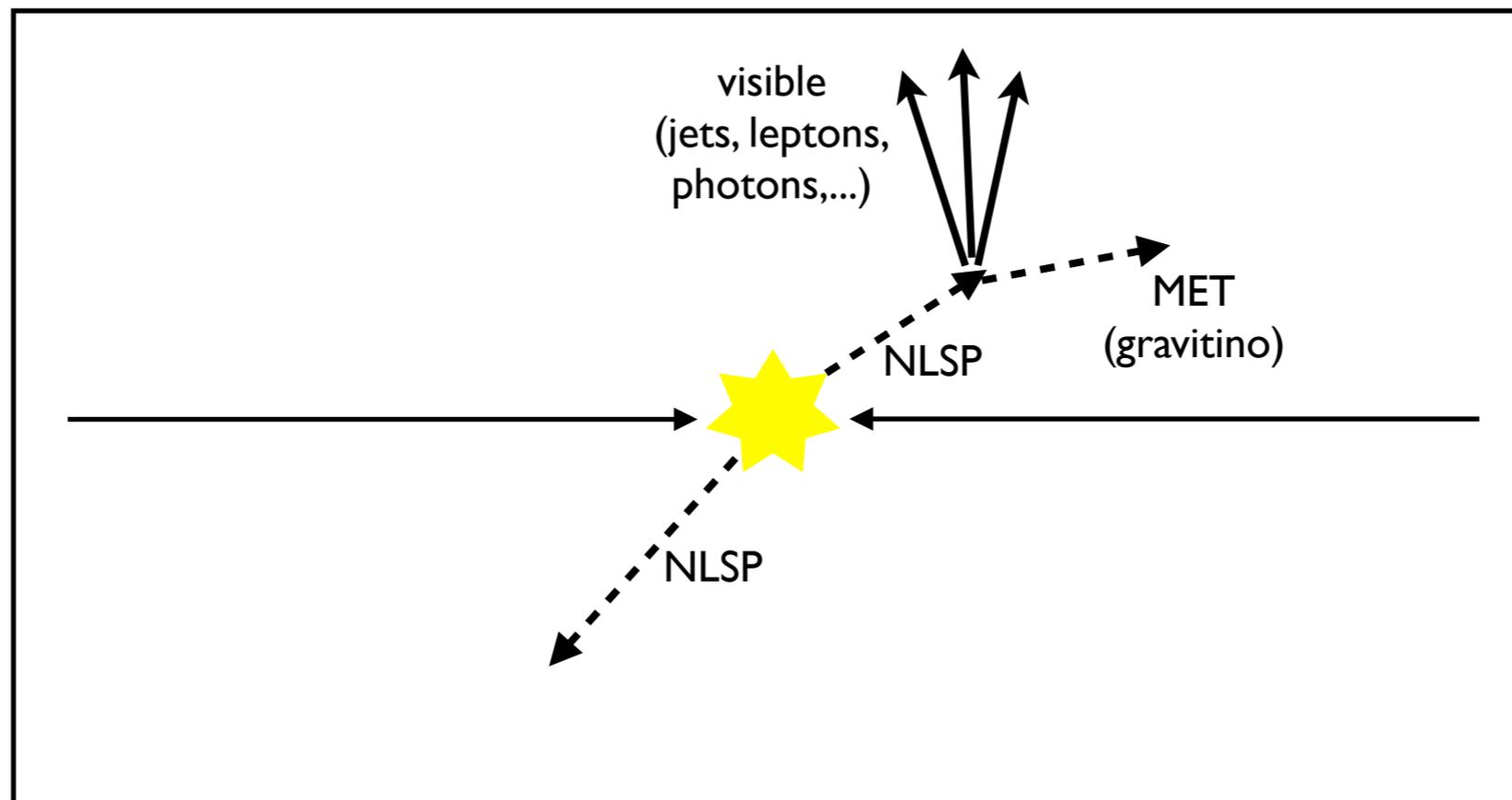


$$M_3 \gtrsim 3 \text{ TeV}, \quad M_{\text{mess}} \gtrsim 10^8 \text{ GeV}$$

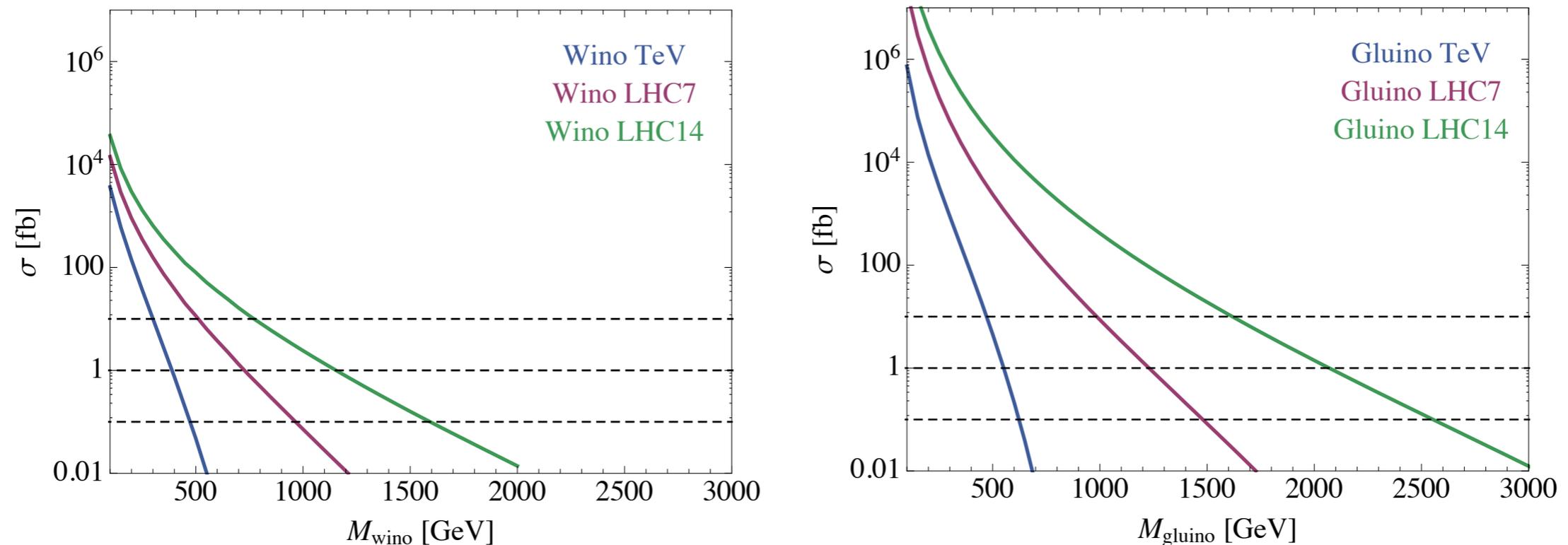
Consequences for Collider Signatures

$$c\tau_{NLSP} \sim 1 \text{ m} \left(\frac{M_{mess}}{10^7 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_{NLSP}} \right)^5$$

For $m_{NLSP} < \sim 1 \text{ TeV}$ and $M_{mess} > \sim 10^8 \text{ GeV}$, NLSP decay is always at least measurably displaced inside the detector.



Consequences for Collider Signatures



For $M_3 > \sim 3$ TeV, gluinos are basically out of reach of the LHC!

Other superpartners (squarks, sleptons, EW-inos) can still be light and accessible.

Summary so far

- “Theorem”: MSSM + GMSB + Higgs@125 => 
- Highly constrained. Huge parts of parameter space ruled out.
 - Messenger scale must be intermediate to high. Displaced or long-lived NLSPs.
 - Gluinos must be heavy.
- Could be the way things are!
- Motivates searches for displaced vertices, CHAMPs, R-hadrons. Also, searches for squark and EW production.
- To evade these conclusions, we must modify one of the starting points.

Evading the constraints

MSSM + GMSB + Higgs@125 => 😞

Options for evading 😞 :

1. The Higgs is not at 125 GeV. The recent hints are just fluctuations.
2. Modify GMSB to achieve large A-terms at the messenger scale.
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But the SM Higgs is now excluded everywhere except around 125 GeV!
Option 1 would require “hiding” the Higgs in the MSSM -- not an easy task!

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Appealing approach: need to extend GGM anyway for μ and B_{μ} :

$$\delta W = \lambda \mathcal{O}_{\mu} H_u H_d + \lambda_u \mathcal{O}_d H_u + \lambda_d \mathcal{O}_u H_d$$

Meade, Seiberg & DS
Komargodski & Seiberg

A-terms beyond GGM

- Many problems with this approach...
- One-loop A-term typically implies one-loop $m_{H_u}^2$

$$A_t \sim \int d^4\theta X^\dagger H_u^\dagger H_u \quad m_{H_u}^2 \sim \int d^4\theta X^\dagger X H_u^\dagger H_u$$

Very analogous to the $\mu/B\mu$ problem!

$$\mu \sim \int d^4\theta X^\dagger H_u H_d \quad B_\mu \sim \int d^4\theta X^\dagger X H_u H_d$$

- Successful solution to $\mu/B\mu$ generally implies small A-terms (Komargodski & Seiberg)
- Currently investigating this in more detail (Craig, Knapen, DS, ... in progress).

Evading the constraints

MSSM + GMSB + Higgs@125 => 😞

Options for evading 😞 :

1. The Higgs is not at 125 GeV. The recent hints are just fluctuations.
2. Modify GMSB to achieve large A-terms at the messenger scale.
3. Modify the MSSM to boost the Higgs mass, allowing for small A-terms at the weak scale.

Higgs beyond the MSSM

$$W \supset \lambda S H_u H_d + \dots$$

- In NMSSM, can have $m_{\text{higgs}} = 125 \text{ GeV}$ without requiring large A -terms.
- Many well-known problems with the NMSSM (Singlet tadpoles, domain walls...) (cf N.Weiner's talk)
- NMSSM+GMSB has even more problems:
 - A -terms for singlet are small for the same reasons as before.
 - Need these A -terms together with negative singlet mass-squared for successful EWSB and extended Higgs spectrum. Generally near-impossible to achieve. (Dine & Nelson '93; de Gouvea, Friedland, Murayama '97; Morrissey & Pierce '08;....)

Higgs beyond the MSSM

- Other extensions of the MSSM?
 - compositeness, “fat Higgs”, non-decoupling D-terms, extra vector-like matter...
 - Luty, Terning, Grant '00; Harnik, Kribs, Larson, Murayama '03; Chang, Kilic, Mahbubani '04; Birkedal, Chacko & Nomura '04; Delgado & Tait '05; Craig, Stolarski, Thaler '11; Csaki, Randall, Terning '12,
 - Generally, unification is problematic in these approaches...

Conclusions

- 125 GeV Higgs in the MSSM+GMSB is highly constrained. Requires high messenger scale and heavy gluinos.
- This motivates the following avenues for exploration:
 - extensions of the MSSM + GMSB
 - MSSM + extensions to GMSB
- Cosmological consequences?
 - High messenger scale \rightarrow gravitino problem?
 - Long-lived NLSPs \rightarrow constraints from BBN?
 - Negative squark mass-squareds in RGEs \rightarrow meta-stable vacuum?

IMPORTANT DISCLAIMER

- Questions I've heard from people:
 - Is GMSB ruled out by the 125 GeV Higgs?
 - Are standard collider signatures of GMSB such as $\gamma\gamma$ +MET suddenly irrelevant?
 - Should I work on something else?
- We can imagine that some modification of the Higgs sector of the MSSM+GMSB boosts the Higgs mass to 125 GeV while preserving all the usual collider signatures.
- Much too early to say! Sensible experimentalists should ignore these theoretical struggles and continue looking for new physics!!

The End