

III.A - Point Sources of Radiation (11 charts)

A) Radiation Units
1) Units (ICRU)
2) Solid Angle
3) X-ray emission
4) Radiation Exposure

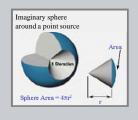
III.A.1 - Units (ICRU)

The International Commission on Radiation Units and Measurements (ICRU) was established in 1925 by the International Congress of Radiology. Since its inception, it has had as its principal objective the development of internationally acceptable recommendations regarding quantities and units of radiation and radioactivity

Name	Symbol	SI unit	Alternate units
'Particle' number	N	1	-
'Particle' flux	Ň	s-1	-
'Particle' fluence	Φ	m-2	-
'Particle' fluence rate	$\dot{\phi}$	m-2 s-1	-
Energy fluence	Ψ	J m ⁻²	
Energy fluence rate	ψ̈́	W m ⁻²	-
Exposure	X	C kg ⁻¹	Roentgen
Exposure rate	X	C kg-1 s-1	Roentgen/sec
Decay constant	λ	s-1	-
Activity	A	s-1	Becquerel/Curie
/BIOE 481 - 2019		See ICRU Repo	ort #85a, Oct 2011

III.A.2 - solid angle definition

For physical processes which are naturally described with a polar coordinate system, it is often necessary to identify the fraction of a unit sphere interior to a surface formed by moving the radial vector to form a conic structure. By convention, the entire unit sphere is defined to have 4π steradians.



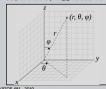
The steradian is the unit used to describe the "solid angle" associated with any portion of the unit sphere.

III.A.2 - differential solid angle

If ϕ is the polar angle from a fixed zenith direction (z) of a spherical coordinate system, and θ is the azimuthal angle of a projection to a plane perpendicular to the zenith (xy), then a differential quantity of solid angle can be written as;

$$d\Omega = \sin(\phi)d\phi d\theta$$

The $sin(\phi)$ term is required because of the shorter arc traced by $d\theta$ for angles of ϕ near the poles. The total solid angle of the unit sphere can then be computed by integration of $d\Omega$ to show that this definition of $d\Omega$ leads to the unit sphere having 4π steradians:



$$\Omega = \int_{0}^{2\pi} \int_{0}^{\pi} \sin\phi d\phi d\theta$$
$$= 2\pi \int_{0}^{\pi} \sin\phi d\phi = 4\pi$$

III.A.2 - fluence in photons/steradian

. Point Sources:

For sources which emit radiation from a region small enough to be considered a point source, the radiation travels in all directions. Typical radionuclide sources emit radiation with no bias as to the direction and are said to have isotropic emission.

- For a source which isotropically emits N photons,
- The fluence is N photons per 4π steradians (N/ 4π photons/sr).
- Fluence at distance r:

If one considers a sphere with a radius of r mm, this source will produce a fluence of photons traveling through the surface of the sphere equal to:

• N/4πr² photons/mm².

Fluence Units:

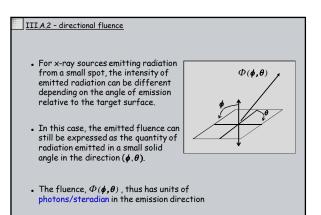
Radiation fluence can either be expressed in terms of photons/steradian or photons/mm². To convert from photons/steradian to photons/mm², simply divide by r^2 , as seen in the above example.

photons/mm² = (photons/steradian) / r², for r in mm

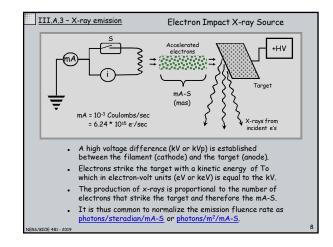
It is often more convenient to describe the fluence from a source in photons/steradian since it is independent of the distance (i.e. radius) from the emission point.

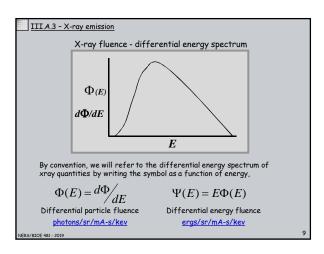
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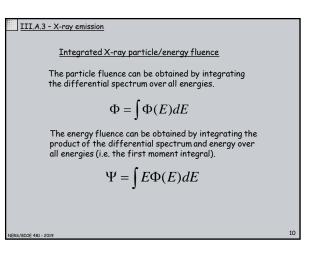
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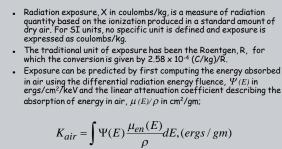


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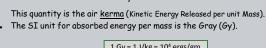


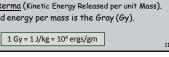


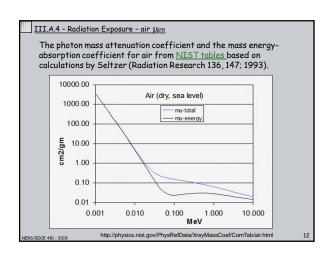


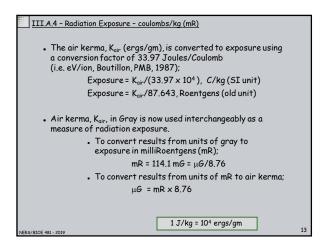


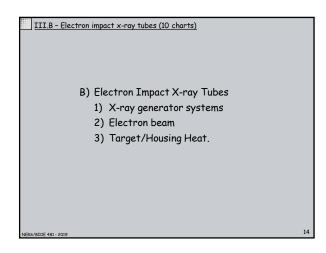
III.A.4 - Radiation Exposure - air kerma, ergs/gm

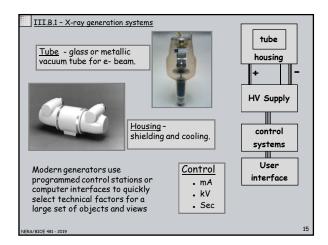


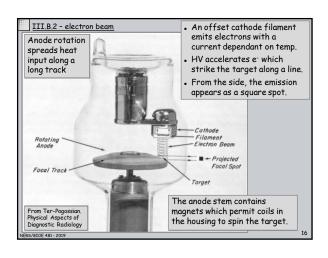


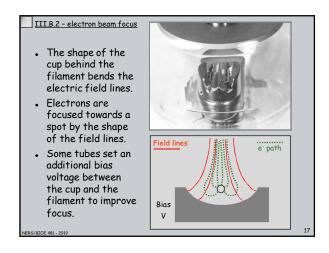


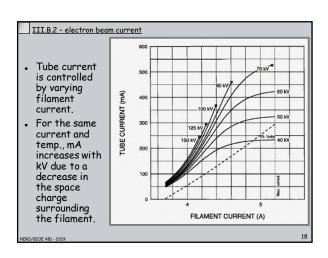


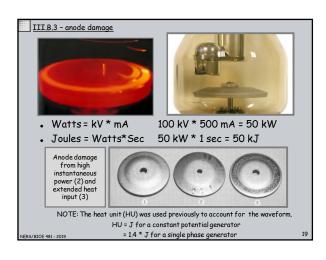


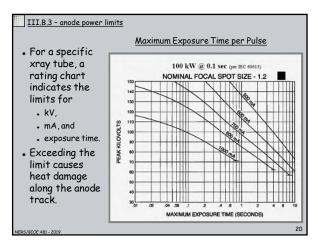


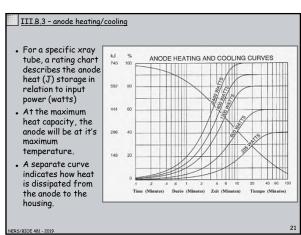






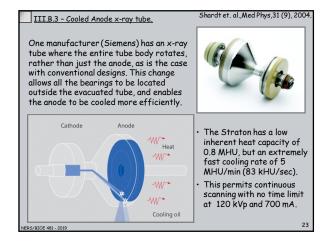


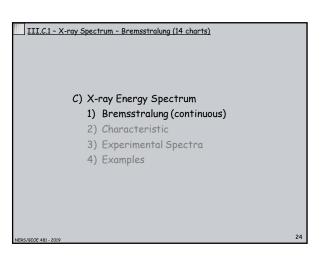




Some radiation imaging devices require that the x-ray tube be run at high power for extended times.

CT scanner, 100kW ~30 sec
Angiography, 120kW 100s pulses
These systems require excellent heat transfer from the anode to the housing.
Circulating oil and a heat exchanging transfers heat out of the tube housing.





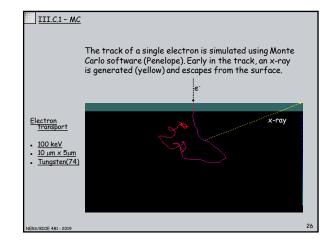
III.C.1 - Bremsstrahlung

- Bremsstrahlung, German for braking radiation, is electro-magnetic radiation produced by the acceleration of a charged particle, such as an electron, when deflected by another charged particle, such as an atomic nucleus.
- An electron gradually looses energy as it slows down in a material. At any point along it's path, a bremsstrahlung photon may be created.

e *7*E ∢ T

In an individual deflection by a nucleus, the incident particle can radiate any amount of energy from zero up to its total kinetic energy T.

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III.C.1.a -Kramers & Kuhlenkampff

In 1923, Hendrik Antonie Kramers (1894-1952) published a significant theoretical paper which included a derivation of the continuum energy spectrum. Kramers began with the quantum theory of Bohr to provide the theoretical basis for his relationship. The paper is one of the first applications of the then new quantum theory to a practical physics problem.

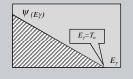
$$\varphi_{(E_{\gamma})} = KZ^{(T_{o} - E_{\gamma})} / E_{\gamma}$$

 E_{γ} = emitted xray energy

 $T_o = incident electron energy, i.e. kV$

K, phots/keV/mA-s/sr

- $K = 6.64 \times 10^8$ @ 30 keV
- K = 6.31 X 10⁸ @ 40 keV K = 4.99 X 10⁸ @ 180 keV



 $\psi_{(E_\gamma)} = E_\gamma \phi_{(E_\gamma)} = KZ (T_{\rm o} - E_\gamma)$

This theoretical result agreed well with the experimental results published by Kuhlenkampff the year before (1922, Ann. Physik)

Kramers HA, On the theory of X-ray absorption and of the continuous X-ray spectrum, Philos. Mag., 46(275):836-871N, Nov. 1923. (Communicated by Prof N. Bohr, Copenhagen)

III.C.1.a - Brems. production efficiency

Kramer's relationship is easily integrated to compute the total radiative energy produced by a thick target.

$$\int \psi(\varepsilon_{\gamma})dE_{\gamma} = \int E_{\gamma}\phi(\varepsilon_{\gamma})dE_{\gamma}$$

$$= \int_{0}^{T_{o}} KZ(T_{o} - E_{\gamma})dE_{\gamma}$$

$$= \frac{1}{\pi}KZT^{2}$$

• Using: To = 100 keV and Z = 74

K = 6 x 108 photons/keV/mA-S/sr and 2π steradians (sr) the radiated energy is

$$E_{rad} = 1.391 \times 10^{15} \text{ keV/mA-S}$$

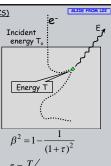
- Using 1 mA-S = 6.24×10^{15} electrons, this becomes $E_{rad} = .22 \, kev/electron$
- Since we assumed 100 keV/electron, the efficiency for converting the energy in the electron beam to radiation is 0.2%

III.C.1.b -Brems. Differential Cross Section (DCS)

- The probability per atom that an electron traveling with energy T will produce an x-ray within the energy range from E to E+dE is known as the differential radiative cross section, $d\sigma_r/dE$.
- Theoretic expressions indicate that the bremsstrahlung DCS can be expressed as;

$$\frac{d\sigma_r}{dE} = f_r(T, E, Z) \frac{Z^2}{\beta^2} \frac{1}{E}$$

- Where β is the velocity of the electron in relation to the speed of light.
- The slowing varying function, $f_r(T,E,Z)$, is often tabulated as the scaled bremsstrahlung DCS. (SHOWN IN LOZ)



Seltzer SM & Berger MJ, Atomic Data & Nucl. Data Tables, 35, 345-418(1986).

III.C.1.b -Integral solution for bremsstrahlung production

• The total radiative production of xrays with energy in the range from E to E+dE can be found by integrating the production per unit pathlength over the path of the electron.

$$\mu_{rs}(\tau, E) = N\sigma_{r}(\tau, E) = \frac{d\mu_{rs}}{dEdS}$$

 $\mu_{rs}(T, E) = 0, T \le E$

Probability per cm per keV

$$\phi(E) = \frac{d\phi}{dE} = \int_{0}^{S(E)} \mu_{rs}(T(s), E) dS$$

· Using the electron stopping power, dT/dS, this can be converted to an integration over the energy of the electron as it slows down.



T=0 S = S(T₀)

S=0

T to T+dT

S to S-dS

III.C.1.c - A simplified integral solution

An early quantum-mechanical theory of radiative collisions (Evans, chapter 20) results in the following expression for the radiative DCS.

$$\frac{d\sigma_r}{dE} = \sigma_o B Z^2 \frac{T + m_o c^2}{T} \frac{1}{E}, cm^2 / nucleus$$

$$\sigma_o = \frac{1}{137} \left(\frac{e^2}{m_o c^2}\right)^2 = 0.580, millibarns / nucleus$$

- Where ${\it B}$ is a very slowly varying function of ${\it Z}$ and the electron energy, ${\it T}$, with a value of approximately 10
- The term $(T+mc^2)/T$ is equal to $0.5/\beta^2$ for small T. This expression is thus consistent with the scaling of the cross shown in the prior slide.
- At values of T small relative to mc^2 and for a constant value of B , this can be used to deduce an approximate expression for the bremsstrahlung spectra.

$$\phi(E) = N_o \frac{\rho}{A} \sigma_o B Z^2 \int_{E}^{T_o} \left[\frac{m_o c^2}{T} \frac{1}{E} \right] / \left(\frac{dT}{ds} \right) dT$$

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III.C.1.c - Integrating the inverse stopping power

The stopping power can be approximated by an expression proportional to the inverse of the electron energy ($\sim 1/T$);

$$dT/dS = -k_a \left(\frac{Z}{A}\right) \left(\frac{\rho}{T}\right), kev/cm$$

Note: in lecture 2, we saw that a better approximation is $1/T^{0.65}$. We use 1/T now to permit integration.

The integration of the inverse of the stopping power can be used to estimate the pathlength of the electron. For a stopping power proportional to I/T, the pathlength is proportional to the incident electron energy squared.

The <u>Thomson-Whiddington</u> law described electron range as proportional to energy squared (Whiddington, Proc. Roy. Soc. London, A86, 1912)

III.C.1.c - Equivalent Kramers model

Using the approximation that the stopping power can be approximated by an expression proportional to $\it I/T$,

$$dT/dS = k_a \left(\frac{Z}{A}\right) \left(\frac{\rho}{T}\right), kev/cm$$

 $\frac{dT}{dS} = k_a \bigg(\frac{Z}{A}\bigg)\bigg(\frac{\rho}{T}\bigg) kev/cm$ The simplified integral solution evaluates to an expression essentially the same as Kramer's equation,

$$\begin{split} \phi(E) &= N_o \frac{\rho}{A} \sigma_o B Z^2 \int_E^T \left[\frac{m_o c^2}{T} \frac{1}{E} \right] / k_o \left(\frac{Z}{A} \right) \left(\frac{\rho}{T} \right) dT \\ \phi(E) &= \frac{N_o \sigma_o B}{L} m_o c^2 Z \frac{1}{E}^T \int_0^T dT \end{split}$$

 $\phi(E) = \left[\frac{N_o \sigma_o B}{k_a} m_o c^2 \right] Z \frac{T_o - E}{E}, xrays/electron/keV$

 $\frac{k_e}{4\pi} \left[\frac{N_o \sigma_o B}{k_o} m_o c^2 \right] = 6.67E08, xrays/mAS/keV/sr$

See 'Flynn LO3b' on course website

III.C.1.d - Self Absorption

- X-rays produced at some depth within the target that have a very low energy, are frequently absorbed within the target.
- One approach to account for this self-absorption is to include a term within the integral solution describing the probability of escape to x-rays of energy E produced by electrons of energy T.

$$\phi_{(E)} = \int_{E}^{T_a} \frac{\sigma_{rs}(T, E)}{dT/dS} F_a(E, T) dT$$

- In an integral solution using:
 - · improved B in the cross section
 - improved stopping power

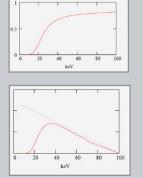
The self absorption term has been computed by considering the mean depth of electron penetration.

- φ(E)E (Kramers)
- $\phi(E)E$ (integral)

See 'Flynn LO3b' on course website

III.C.1.d - Intrinsic Absorption

- The attenuation by the internal materials of the tube and housing is significant below about 40 keV for general radiography tungsten target tubes. This is commonly referred to as 'intrinsic filtration'.
- The effect of intrinsic filtration on the energy fluence spectrum is seen to further reduce low energy emissions such that the spectrum is similar to Kramer's equation above 40keV.
 - $\phi(E)E$ (Kramers)
 - $\phi(E)E$ (integral)



See '<u>Flynn LO3b</u>' on course website

III.C.1.e - Prior integral bremsstrahlung models

- Kramers HA, Philos. Mag. 46(275) 1923. Semi-classical DCS, 1/T dT/dS, no absorption
- Storm E, Phys. Rev. A 5(6) 1972. Born/Sommerfield DCS, Berger&Seltzer dT/dS, fixed depth
- Birch & Marshall, Phys. Med. Biol. 24(3) 1979. polynomial DCS, Bethe dT/dS, T-W penetration
- Tucker et.al., Med. Phys., 18(2&3) 1991. Polynomial DCS, Berger&Seltzer dT/dS, T-W penetration
- For these integral models, electron transport effects $(back scatter, absorption, angular\ distributions)\ are\ approximated$ by simple expressions.
- Dodge has recently developed an advanced integral model (WSU 2008) that uses electron transport distributions determined from Monte Carlo simulations.

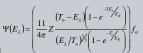
III.C.1.e - The Storm model (xspect 3.5)

- A notable work on the modeling of the continuous spectrum was published by Storm in 1972 (Storm, Phys. Rev. A, 5(6):2328-2338, June 1972).
- Storm formally evaluated several cross sections detailed by Koch and Motz (ref 2). These cross sections have more validity than the Compton and Allison cross section used by most other investigators. He shows that for spectral estimation the best fit to experimental data is obtained with a differential (in energy) cross section derived using the Born approximation with no screening (3BN).
- He then presented a mathematical fit for the bremsstrahlung intensity which specifically accounts for electron backscatter.

The Storm model is used to compute the bremsstralung spectrum in xSpect 3.5 used in the NERS 580 computational lab course. The Dodge model is to be used in xSpect 4.0 (yet to be released).

III.C.2 - X-ray Spectrum - Characteristic (13 charts)

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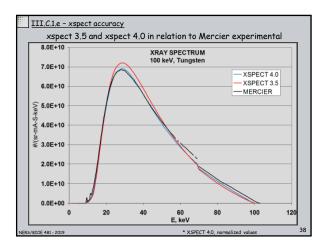
 $\Psi(E\gamma) = \mathsf{diff.}$ energy fluence

 $E\gamma$ = emitted x-ray energy

 T_o = electron energy (high voltage)

 $E_K = K$ binding energy

 $f_a = \mathbf{self}$ absorption



C) X-ray Energy Spectrum 1) Bremsstralung (continuous) 2) Characteristic 3) Experimental Spectra 4) Examples

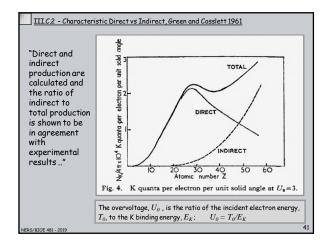
The emission of radiation with energies characteristic of the target material results from atomic shell transitions that occur as a result of a vacancy created in an inner shell, usually the K or L shell.

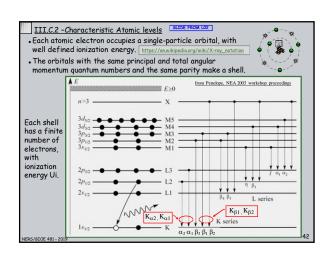
• Direct production:

As each electron penetrates into the target, shell vacancies are occasionally produced by electron-electron interactions in the atoms of the target material.

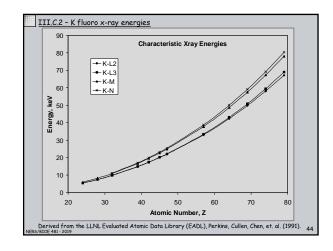
• Indirect production:

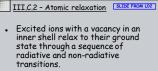
Additionally, many of the bremsstralung x-rays produced by electron-nucleus interactions are absorbed in the target by photo-electric interactions which result in shell vacancies, primarily the K shell.





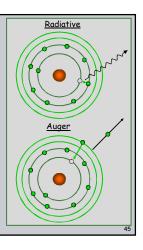
C.2 - Characteristic Energies		NIST X-ray Transition Energy D				
Material	Z	Κ _{α.2}	$K_{\alpha 1}$	$K_{\beta 2}$	$K_{\beta 1}$	
Cr	24	5.40	5.41	6.00	5.95	
Y	39	14.88	14.96	17.01	16.74	
Мо	42	17.37	17.48	19.96	19.61	
Rh	45	20.07	20.22	23.17	22.72	
W	74	57.98	59.32	69.07	67.25	
Pt	78	65.12	66.83	77.83	75.75	
		K-L2	K-L3	K-N2,3	K-M3	
notations vary in the literature. $K_{\alpha 2}$ is the Siegbahn notation.			K-L2 is the IUPAC notation.			
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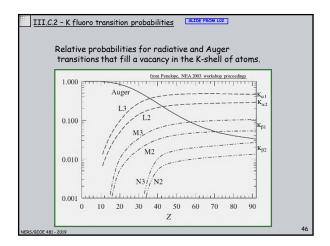


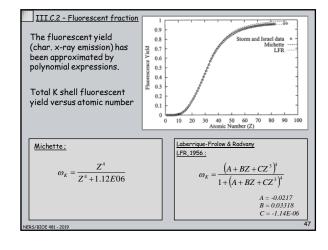


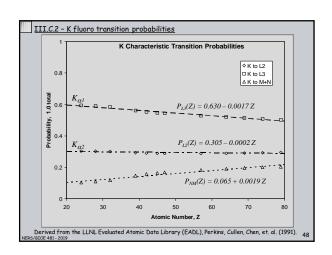
- In a radiative transition, the vacancy is filled by an electron from an outer shell and an x ray with characteristic energy is emitted.
- In a non-radiative transition, the vacancy is filled by an outer electron and the excess energy is released through emission of an electron from a shell that is farther out (Auger effect).
- Each non-radiative transition generates an additional vacancy that in turn, migrates "outwards".

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III.C.2 - Characteristic KL production, Storm 1972 "Webster and Clark were the first of many investigators to report that the K-photon intensity could be described by an empirical formula of the form" $\phi_K = C_K (T_O - E_K)^{\eta_K}$ "The present calculation indicates this formula is good for tungsten up to values of E_O -70 = 100 kV with" $C_K = 4.25 \times 10^8$ MEASUREMENTS HETTINGER & STARFELT (20°CP: STORM, ISRAEL & LIER • NORELCO (22°,CP) • TRIPLETT-BARTON (20°,SR) • PICKER (20°,SR) $\eta_{\scriptscriptstyle K}$ = 1.67 "And C_K in units of 11111 photon/(sec mA sr)." E. - E. (keV) Storm, J. Appl. Phys., Vol. 43, No. 6, June 1972 ERS/BIOE 481 - 2019 • Webster, Proc. Natl. Acad. Sci. US, 3, 181 (1917)

