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$$p(q_e, E)dE = \left[\int_{0}^{E} p(q_e, e)p(e, E)de\right]de$$

- For monoenergetic radiation of energy Ei, the charge signal from  $N_i$  detected photons is deduced from integration of the charge production probability.

$$Q_e = N_i \int_{a}^{q_{\text{max}}} q_e p(q_e, E_i) dq_e = N_i \overline{Q}_E$$

• This is equivalent to considering the average deposited charge from the discrete sum of all events.

$$Q_e = N_i \frac{\sum_{i=1}^{N_i} q_n}{N_i}$$

IOE 481 - 2019





















V.A.3.e leakage current noise
• The leakage current contributes to the signal
in relation to the signal integration time, $t_{int}$ . $Q_l = i_l t_{int}$
<ul> <li>While the signal can be corrected to eliminate the leakage current contribution, their remains an added noise from the number of electrons associated with the collected leakage current.</li> </ul>
$n_{le} = Q_l / 1.602E$ -19 $\sigma_{le} = (Q_l / 1.602E$ -19) <sup>1/2</sup>

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Polycrystalline I HgI <sub>2</sub> , and lead been investigat semi-conductor high x-ray abso	mercuric iodide, F ed as lar materia rption. 2003	c iodide, PbI <sub>2</sub> , have rge area Ils with	HgI	2, Sel	lin2001	
<ul> <li>Relative to a-Se, imp</li> <li>Charge collection collection</li> </ul>	proved abso nsistent wi	orption (high) th Hecht rela	Z) and r tion.	reduce	ed Weff.	
<ul> <li>Relative to a-Se, imp</li> <li>Charge collection collection</li> </ul>	proved absonsistent wi	orption (high) th Hecht rela	Z) and r tion.	reduce	ed Weff.	7
Relative to a-Se, imp     Charge collection col     Atomic Number (Z)	proved abso nsistent wi Poly-HgI <sub>2</sub> 80, 53	orption (high) th Hecht rela Poly-Pbl <sub>2</sub> 82,53	Z) and r tion. a-Se 34		ed Weff.	7
Relative to a-Se, imp     Charge collection co     Atomic Number (Z)     Energy Band Gap (Eg) eV	proved abso nsistent wi Poly-HgI <sub>2</sub> 80, 53 2.1	orption (high) th Hecht rela <u>Poly-Pbl</u> 2 82,53 2.3	Z) and r tion. <u>a-Se</u> <u>34</u> 2.2	reduce	ed Weff.	7
Relative to a-Se, imp     Charge collection co     Atomic Number (Z)     Energy Band Gap (Eg) eV     Effective Charge Pair Formation     Energy (W), eV	Poly-HgI2 80, 53 2.1 -5	Poly-Phi <sub>2</sub> 82,53 2.3 -5.5	Z) and r tion. a-Se 34 2.2 -42	reduce	ed Weff.	]
Relative to a-Se, imp     Charge collection co     Atomic Number (Z)     Energy Band Gap (E) eV     Effective Charge Pair Formation     Energy (W), eV Mobility Life-time Product (µ7)     em <sup>2</sup> V	Poly-Hgl <sub>2</sub> 80,53 2.1 -5 1.5x10 <sup>-5</sup>	Poly-Pbl <sub>2</sub> 82,53 2.3 -5.5 (hole) 1.8x10 <sup>6</sup> (electron) 7x10 <sup>3</sup>	Z) and r ttion. <u>a-Se</u> <u>34</u> 2.2 -42 10 <sup>6</sup> - 10 <sup>3</sup>	educe 8.0 4.0 4.0 4.0 4.0 4.0 4.0 4.0 4.0 4.0 4	ed Weff.	

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27





## V.A.3.f -CdTe, CdZnTe (CZ, CZT)

- The heterogeneous material structure of PbI<sub>2</sub> prohibits consistent measures of the charge from each x-ray.
- The lower resistivity of CZT prohibits its use as an integrating x-ray detector. However, the crystalline nature of CZT makes it attractive for photon counting detectors and radioisotope imaging cameras.

































V.A.4.b - Scintillation material properties								
From: van Eijk 2002, pg 89								
	Density (g cm <sup>-3</sup> )	$\rho Z^4_{eff}$ (10 <sup>6</sup> )	Attenuation length at 511 keV (mm)/ prob. phot. eff. (%)	Hygro- scopicity	Light yield (photons/ MeV)	Decay time (ns)	Emission maximum (nm)	
CsI:Na	4.51	38	22.9/21	Yes	40 000	630	420	
CsI:Tl	4.51	38	22.9/21	Slightly	66 000	$800 -> 6 \times 10^3$	550	
CaWO <sub>4</sub>	6.1	89	13.6/32	No	20 000 <sup>b</sup>		420	
YTaO <sub>4</sub> :Nb	7.5	96	11.8/29	No	40 000 <sup>b</sup>		410	
Gd <sub>2</sub> O <sub>2</sub> S:Tb	7.3	103	12.7/27	No	60 000 <sup>b</sup>	$1 \times 10^{6}$	545	
Gd2O2S:Pr,Ce,F	7.3	103	12.7/27	No	35 000 <sup>b</sup>	$4 \times 10^{3}$	510	
Gd <sub>2</sub> O <sub>2</sub> S:Pr (UFC)	7.3	103	12.7/27	No	50 000 <sup>b</sup>	$3 \times 10^{3}$	510	
Y <sub>1.34</sub> ,Gd <sub>0.60</sub> O <sub>3</sub> :(Eu,Pr) <sub>0.06</sub> <sup>c</sup> (Hilight)	5.9	44	17.8/16	No	42 000 <sup>b</sup>	$1 \times 10^{6}$	610	
Gd3Ga5O12:Cr,Ce	7.1	58	14.8/18	No	40 000 <sup>b</sup>	$140 \times 10^{3}$	730	
CdWO <sub>4</sub>	7.9	134	11.1/29	No	20 000 <sup>b</sup>	$5 \times 10^{3}$	495	
Lu2O3:Eu,Tb	9.4	211	8.7/35	No	30 000 <sup>b</sup>	$>10^{6}$	611	
NER5/BIOE 481 - 2019 48								

From: van Eijk 2002, pg 89							
	Density (g cm <sup>-3</sup> )	$\rho Z^4_{eff}$ (10 <sup>6</sup> )	Attenuation length at 511 keV (mm)/ prob. phot. eff. (%)	Hygro- scopicity	Light yield (photons/ MeV)	Decay time (ns)	Emission maximum (nm)
CaHfO3:Ce	7.5	139	11.6/30	No	$\sim \! 10.000^{b}$	40	390
SrHfO3:Ce	7.7	122	11.5/28	No	$\sim 20\ 000^{b}$	40	390
BaHfO3:Ce	8.4	142	10.6/30	No	$\sim 10.000^{b}$	25	400
NaI:TI	3.67	24.5	29.1/17	Yes	41 000	230	410
LaCl3:Ce	3.86	23.2	27.8/14	Yes	46 000	25 (65%)	330
LaBr3:Ce	5.3	25.6	21.3/13	Yes	61 000	35 (90%)	358
Bi4Ge3O12 (BGO)	7.1	227	10.4/40	No	9 000	300	480
Lu2SiO5:Ce (LSO)	7.4	143	11.4/32	No	26 000	40	420
Gd2SiO5:Ce (GSO)	6.7	84	14.1/25	No	8 000	60	440
YAIO3:Ce (YAP)	5.5	7	21.3/4.2	No	21 000	30	350
LuAlO3:Ce (LuAP)	8.3	148	10.5/30	No	12 000	18	365
Lu2Si2O7:Ce (LPS)	6.2	103	14.1/29	No	30 000	30	380









ERS/BIOE 481 - 2019

(Arbitrary Vertical Axis Units)

53

V.A.4.c - Indirect conversion efficiency									
Conversion of gamma energy to electrons.									
• photons/keV(NaI) 40									
<ul> <li>Light collection efficiency</li> </ul>	.50								
Photo-conversion efficiency (PMT) .20									
<ul> <li>Thus for 140 keV gamma rays (Tc 99m), we will collect 4 electrons/keV for a total of 560 electrons.</li> </ul>									
<ul> <li>The standard deviation in the signal associated with 560 electrons is 4.2% (i.e. 1/N<sup>‡</sup>) which corresponds to a FWHM of about 9.7%.</li> </ul>									
<ul> <li>This is typical of the energy resolution of nuclear medicine Anger cameras using NaI crystals and PMT detectors.</li> </ul>									
ERS/BIOE 481 - 2019	54								









